# Multiphase processes in the EC-Earth Earth System model and their relevance to the atmospheric oxalate, sulfate, and iron cycles

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**Abstract.** Understanding how multiphase processes affect the iron-containing aerosol cycle is key to predict ocean biogeochemistry changes and hence the feedback effects on climate. For this work, the EC-Earth Earth system model in its climate-chemistry configuration is used to simulate the global atmospheric oxalate (OXL), sulfate (SO4<sup>2-</sup>), and iron (Fe) cycles, after incorporating a comprehensive representation of the multiphase chemistry in cloud droplets and aerosol water. The model

25 considers a detailed gas-phase chemistry scheme, all major aerosol components, and the partitioning of gases in aerosol and atmospheric water phases. The dissolution of Fe-containing aerosols accounts kinetically for the solution's acidity, oxalic acid, and irradiation. Aerosol acidity is explicitly calculated in the model, both for accumulation and coarse modes, accounting for thermodynamic processes involving inorganic and crustal species from sea-salt and dust.

Simulations for present-day conditions (2000-2014) have been carried out with both EC-Earth and the atmospheric composition component of the model in standalone mode driven by meteorological fields from ECMWF's ERA-Interim reanalysis. The calculated global budgets are presented and the links between the 1) aqueous-phase processes, 2) aerosol dissolution, and 3) atmospheric composition, are demonstrated and quantified. The model results are supported by comparison to available observations. We obtain an average global OXL net chemical production of  $12.615 \pm 0.064$  Tg yr<sup>-1</sup> in EC-Earth, with glyoxal being by far the most important precursor of oxalic acid. In comparison to the ERA-Interim simulation,

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differences in atmospheric dynamics as well as the simulated weaker oxidizing capacity in EC-Earth result overall in a  $\sim$ 30 % lower OXL source. On the other hand, the more explicit representation of the aqueous-phase chemistry in EC-Earth compared to the previous versions of the model leads to an overall  $\sim$ 20\_% higher sulfate production, but still well correlated with atmospheric observations.

40 The total Fe dissolution rate in EC-Earth is calculated at  $0.806 \pm 0.014 \text{ Tg} \text{vr}^{-1}$  and is added to the primary dissolved Fe (DFe) sources from dust and combustion aerosols in the model  $(0.072 \pm 0.001 \text{ Tg} \text{vr}^{-1})$ . The simulated DFe concentrations show a satisfactory comparison with available observations, indicating an atmospheric burden of ~0.007 Tg, and overall resulting in an atmospheric deposition flux into the global ocean of  $0.376 \pm 0.005 \text{ Tg} \text{vr}^{-1}$ , well within the range reported in the literature. All in all, this work is a first step towards the development of EC-Earth into an Earth System Model with fully interactive

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45 bioavailable atmospheric Fe inputs to the marine biogeochemistry component of the model.

#### 1 Introduction

Clouds, fog, and deliquescent aerosols host chemical reactions involving inorganic and organic polar atmospheric compounds (Calvert et al., 1985; Chameides and Davis, 1983; Collett et al., 1999; Donaldson and Valsaraj, 2010; Jacob, 1986; Lelieveld and Crutzen, 1991). These reactions result in the production of species that can neither be formed via gas-phase processes

- 55 directly, nor explained solely by primary sources. These compounds participate in chemical transformations across the gas, aqueous, and solid phases. Such multiphase processes have a significant impact on the atmospheric cycles of important inorganic species like sulfur (e.g., Hoyle et al., 2016; Seinfeld and Pandis, 2006; Tsai et al., 2010), and act as a complementary pathway for the formation of organic particulate matter (e.g., Lin et al., 2014; Liu et al., 2012; Myriokefalitakis et al., 2011). The produced inorganic and organic aerosols serve as cloud condensation nuclei and thus affect the Earth's energy balance
- 60 (Stocker et al., 2013).

Multiphase processes may also impact the global carbon balance indirectly by altering the atmospheric cycles of species that act as nutrients for the marine biota (Hamilton et al., 2022; Kanakidou et al., 2018; Mahowald et al., 2017; Myriokefalitakis et al., 2020a). Nutrient availability in marine ecosystems is key for the primary production that modulates both the surface oceanic concentrations and the uptake of <u>atmospheric</u> CO<sub>2</sub> (e.g., Le Quéré et al., 2007, 2013; Smith, 2019). A large portion of the

- 65 global ocean is found, however, to be iron-limited (Krishnamurthy et al., 2009, 2010), therefore, the importance of iron (Fe) to oceanic productivity is well established (Hamilton et al., 2020; Kanakidou et al., 2020; Meskhidze et al., 2019; Tagliabue et al., 2016). Besides rivers and sea ice, along with sediment dissolution and hydrothermal vents, which are the main sources of bioavailable Fe in the ocean, the atmospheric deposition of nutrients is the most effective external pathway that provides Fe in the open ocean. Fe is a critical micronutrient for marine biota that is mainly utilized in its dissolved form (e.g., aqueous,
- 70 colloidal, or nanoparticulate). Thus, the atmospheric processing of Fe-containing minerals, i.e., the conversion from insoluble to soluble that is readily available Fe for marine organisms, is a central step in the atmospheric and marine Fe cycles and directly connected to atmospheric multiphase processes.

Fe is mainly present in the atmosphere in crystalline lattices of aluminosilicates or as iron-oxides in dust aerosols ( $\sim$ 95\_%; Mahowald et al., 2009), and tends to be rather insoluble when emitted (up to  $\sim$ 1 % solubility; Journet et al., 2008). In fact,

- 75 observed high Fe solubility downwind of dust source regions can be only explained via the atmospheric processing of dust aerosols (Baker and Jickells, 2017; Oakes et al., 2012). Enhanced Fe solubility is observed for biomass burning aerosols (e.g., ranging 2-46 %; Bowie et al., 2009; Guieu et al., 2005b; Mahowald et al., 2018; Oakes et al., 2012; Paris et al., 2010), depending strongly on the source region and/or the type of burned wood. <u>Significantly higher Fe solubilities are found, however, for anthropogenic combustion related Fe-containing aerosols, especially for Fe in oil fly ash from industries and</u>
- 80 shipping which is mainly in the form of ferric sulfates (Chen et al., 2012; Ito, 2013; Rathod et al., 2020; Schroth et al., 2009). The uncertainty in Fe-containing combustion aerosol solubility (e.g., Rathod et al., 2020) is nevertheless also reflected in modeling studies, with some models assuming relatively high solubility at emission (e.g., Hamilton et al., 2019;

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Myriokefalitakis et al., 2011) depending on the aerosol size, and others assuming an almost completely insoluble emitted Fe 90 whose solubility is then enhanced during transport via atmospheric processing (Ito, 2015; Ito et al., 2021). Recent multimodel studies estimate an overall global dissolved Fe (DFe) production rate due to atmospheric processing of dust and combustion aerosols of  $0.56 \pm 0.29$  Tg yr<sup>-1</sup> (Ito et al., 2019; Myriokefalitakis et al., 2018), indicating that a large uncertainty still remains in the impact of atmospheric processing on the mineral Fe solubilization processes.

- During atmospheric transport, inorganic strong acids along with organic ligands may coat mineral aerosols and eventually 95 convert part of the contained insoluble Fe forms (e.g., hematite) to bioavailable forms of Fe for marine biota in the euphotic
- zone (e.g., free ferrous forms, inorganic soluble Fe, and organic Fe-complexes). Mineral dissolution rates depend on the solution's acidity levels, the mineral surface concentration of organic ligands, sunlight, and ambient temperature (e.g., Hamer et al., 2003; Lanzl et al., 2012; Lasaga et al., 1994; Zhu et al., 1993). Although sulfate (SO4<sup>2-</sup>) is the dominant aerosol species that controls the aerosol liquid water content and acidity, oxalate ((COO')<sub>2</sub>; hereafter OXL) acts as an organic ligand for the
- 100 Fe-containing aerosol dissolution processes (e.g., Paris et al., 2011; Paris and Desboeufs, 2013) that can effectively break the Fe-O bonds at the mineral's surface via the formation of ligand-containing surface structures (Yoon et al., 2004). Despite the dominant role of acidity in the mineral Fe dissolution processes, modeling estimates (Ito, 2015; Johnson and Meskhidze, 2013; Myriokefalitakis et al., 2015) show the importance of OXL to atmospheric DFe concentrations (e.g., including the formation of Fe(II/III)-oxalato complexes). The dissolution of Fe by OXL may further contribute to the organic-bounded pool of nutrients
- 105 deposited into the ocean, and thus affects the marine primary production especially in oligotrophic subtropical gyres (e.g., up to 20\_%; Myriokefalitakis et al., 2020a).

Notwithstanding their different roles and efficiencies in Fe solubilization processes, atmospheric observations demonstrate a strong correlation between  $SO_4^{2-}$  and OXL concentrations (Yu et al., 2005), especially above clouds (Sorooshian et al., 2006), indicating common chemical production pathways despite the differences in their precursors and primary sources.  $SO_4^{2-}$  and

- 110 OXL are the most common species formed via aqueous-phase reactions of inorganic and organic origin, respectively, with modeling studies supporting that more than 60% of the sulfates (e.g., Liao et al., 2003) and about 90% of oxalates (Lin et al., 2012; Liu et al., 2012; Myriokefalitakis et al., 2011) are produced in clouds. OXL is the dominant dicarboxylic acid (DCA) in the troposphere (e.g., Kawamura and Ikushima, 1993; Kawamura and Sakaguchi, 1999; Norton et al., 1983) and is formed primarily through cloud processing of glyoxal and other water-soluble products of alkenes and aromatics of anthropogenic,
- 115 biogenic, and marine origin (Carlton et al., 2007; Warneck, 2003). OXL is mostly present in the troposphere in the particulate form (Yang and Yu, 2008), with aerosol concentrations roughly four times larger than in the gas phase (Martinelango et al., 2007; Yao et al., 2002). OXL can be present in urban environments (Yang et al., 2009) as well as in remote regions (Sempére and Kawamura, 1994), and is produced during the photochemical aging of organic aerosols (Eliason et al., 2003). The observed correlation of OXL with ammonium (NH4<sup>+</sup>) (Martinelango et al., 2007) indicates that OXL is mostly present as a salt (i.e., the present of the present of the present and the present as a salt (i.e., the present of the present at the p
- $120 \quad \text{ammonium oxalate; (NH_4)}_2C_2O_4) \text{ in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014). Ortiz-Montalvo et al. (2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Paciga et al., 2014) found that in the atmosphere (Pa$

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presence of  $NH_4^+$  under cloud-relevant conditions, the OXL produced by the aqueous-phase glyoxal oxidation is efficiently converted to ammonium oxalate, with its vapor pressure being several orders of magnitude lower than that of oxalic acid. However, in the presence of metals, such as calcium (Ca<sup>2+</sup>) and magnesium (Mg<sup>2+</sup>) from dust and sea salt aerosols, most of

125 the oxalic acid is found to be present in the form of metal complexes (Furukawa and Takahashi, 2011). Nevertheless, due to their different solubility the stability of oxalate complexes can be rather diverse; while calcium- and magnesium-oxalates precipitate from the solution, other salts, such as sodium or ammonium oxalates, remain in a deliquescent form (Furukawa and Takahashi, 2011).

Laboratory and modeling studies support that OXL is directly produced in atmospheric water via glyoxylic acid (GLX;
 HC(O)COOH) oxidation by hydroxyl (OH) and nitrate (NO<sub>3</sub>) radicals. The estimated net global OXL production rate in atmospheric water ranges between 13 and 30 Tg yr<sup>-1</sup> (Lin et al., 2014; Liu et al., 2012; Myriokefalitakis et al., 2011). However,

- modeling studies where the OXL production is only based on the GLX aqueous phase oxidation tend to underestimate its observed atmospheric concentrations (e.g., Lin et al., 2014; Myriokefalitakis et al., 2011). Based on laboratory experiments, Carlton et al. (2007) proposed that predictions of oxalic acid concentrations could be significantly improved when larger
- 135 multifunctional compounds are allowed to be produced under elevated glyoxal concentrations in typical cloud conditions. These larger multifunctional products can act as precursors for the glyoxylic and oxalic acids via their rapid oxidation by OH radicals (Carlton et al., 2007). When such reactions are included, models tend to predict a higher oxalate atmospheric load and thus match better the observations (e.g., Myriokefalitakis et al., 2011). Note that although small carbonyl compounds, such as glyoxal and methylglyoxal, can undergo oligomerization under concentrated acidic conditions (Ervens and Volkamer, 2010;
- 140 Lim et al., 2010, 2013), the mechanism behind the production of larger multifunctional products in dilute solutions may be rather complex, e.g., for products with alcohol functional groups, covalently bonded oligomers, larger carboxylic acids, and other humic-like substances (HULIS) components (Altieri et al., 2006; Blando and Turpin, 2000; Cappiello et al., 2003; Carlton et al., 2007).

The involvement of Fe chemistry in the aqueous phase decreases overall (by ~57,%) the global OXL net production rates,

- 145 despite the increase in dissolved OH radical sources and thus the oxidation of OXL precursors (Lin et al., 2014). Besides the dissolved H<sub>2</sub>O<sub>2</sub> photolysis that enhances drastically the OH production in the solution during the daytime, the presence of transition metal ions (TMIs) may play a central role in aqueous-phase oxidizing capacity, especially under dark conditions (Tilgner et al., 2013; Tilgner and Herrmann, 2018). Among other metals, Fe is the most efficient for the aqueous-phase oxidizing capacity, since on one hand, it contributes to the OH reactivity via the Fenton reaction and the direct Fe photolysis,
- 150 and on the other hand its dissolved concentrations are high due to the mineral dust contribution. The metal-oxalato complexes formed in the presence of Fe in the solution (Zuo and Deng, 1997), however, can also undergo Fenton reaction and further increase the dissolved OH source, in particular for air masses of continental origin (Bianco et al., 2020), where elevated concentrations of OXL precursors and Fe-containing aerosols both of lithogenic and pyrogenic sources can exist. The
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- 155 photolysis of Fe-oxalato complex [Fe(C<sub>2</sub>O<sub>4</sub>)<sub>2</sub>]<sup>-</sup> eventually transforms C<sub>2</sub>O<sub>4</sub><sup>2-</sup> into CO<sub>2</sub> in the aqueous phase (Ervens et al., 2003). Overall, it is clear that the impact of the Fe redox chemistry on the OXL production, and vice versa, is a rather complex issue, that <u>can</u> also affect the ligand-promoted dissolution process of the Fe-containing minerals under ambient atmospheric conditions.
- For this work, we incorporate a comprehensive aqueous-phase chemistry scheme into a state-of-the-art global climate-160 chemistry model to simulate the atmospheric multiphase processes with respect to iron-containing aerosol dissolution. Section 2 provides an overview of the model, focusing mostly on the new implementations. In particular, we describe the multiphase chemistry scheme used to simulate the atmospheric OXL, SO4<sup>2-</sup>, and Fe cycles, along with the respective developments for the primary soil and combustion sources applied in the model. In Sect. 3, we present the model-derived OXL, SO4<sup>2-</sup>, and Fe containing aerosol atmospheric concentrations and their evaluation with available observations, and in Sect. 4 we discuss the
- 165 impact of the simulated aqueous-phase processes on the DFe deposition fluxes to the global ocean. Finally, in Sect. 5, we summarize the global implications of explicitly resolving multiphase chemistry in a climate-chemistry model for the atmospheric Fe cycle, along with the plans for future model development.

#### 2 Model description

#### 2.1 The EC-Earth3 Earth System Model

- 170 Our tropospheric multiphase chemistry developments have been implemented in the global Earth System Model (ESM) EC-Earth3 (Döscher et al., 2021). EC-Earth3 took part in the Coupled Model Intercomparison Project phase 6 (CMIP6; Eyring et al., 2016). The atmospheric General Circulation Model (GCM) of EC-Earth3 is based on cycle 36r4 of the Integrated Forecast System (IFS) from the European Centre for Medium-Range Weather Forecasts (ECMWF), which includes the land surface model H-TESSEL (Balsamo et al., 2009). The ocean model is the Nucleus for European Modeling of the Ocean (NEMO)
- 175 release 3.6 (Rousset et al., 2015), with sea ice processes represented by the Louvain-la-Neuve sea ice model (LIM) (Rousset et al., 2015; Vancoppenolle et al., 2009). The ESM presents two configurations: <u>1</u>) the carbon cycle one that represents the marine biogeochemistry processes through PISCES (Aumont et al., 2015), the dynamic terrestrial vegetation through LPJ-Guess (Smith et al., 2001, 2014), and the atmospheric cycle of CO<sub>2</sub> through the Tracer Model version 5 release 3.0 (TM5-MP 3.0) and <u>2</u>) the EC-Earth3-AerChem configuration (van Noije et al., 2021) that represents the atmospheric chemistry and
- 180 transport of aerosols and reactive species, also through the TM5-MP 3.0. Most of the information exchange and interpolation between modules is handled through the Ocean Atmosphere Sea Ice Soil version 3 (OASIS3) coupler (Craig et al., 2017). For this work, specifically, we rely on the EC-Earth3-AerChem branch (van Noije et al., 2021).

EC-Earth3-AerChem includes TM5-MP to simulate tropospheric aerosols and the reactive greenhouse gases methane (CH4) and ozone (O3) and allows the coupling of those species to relevant processes in the atmospheric module IFS (e.g., radiation

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and clouds). The model can be executed in an atmospheric mode only, i.e., using prescribed sea surface temperature and sea ice concentration, or coupled to the NEMO-LIM ocean and sea-ice model. In addition, TM5-MP can run as a standalone (offline) atmospheric Chemistry and Transport Model (CTM) driven by meteorological and surface fields (Krol et al., 2005).

190 The present work is structured around a recently released version of TM5-MP that incorporates a rather detailed gas-phase tropospheric chemistry scheme, the MOGUNTIA (Myriokefalitakis et al., 2020b). MOGUNTIA simulates explicitly the organic polar species that partition in the atmospheric aqueous phase and allows for a sophisticated parameterization of the multiphase processes needed for this study.

All major aerosol components such as sulfate, black carbon, organic aerosols, sea salt, and mineral dust aerosols are included in TM5-MP and are distributed (depending on the aerosol type) in seven lognormal modes, i.e., four soluble modes (i.e., nucleation, Aitken, accumulation, and coarse) and three insoluble modes (i.e., Aitken, accumulation and coarse). The aerosol microphysics in the model is calculated by the modal aerosol scheme M7 (Aan de Brugh et al., 2011; Vignati et al., 2004), which represents both the evolution of the total particle number and mass of the different species in each mode. Ammonium, nitrate, and aerosol water are determined based on gas/particle partitioning. M7 uses seven log-normal size distributions with

- 200 predefined geometric standard deviations, with four water-soluble modes (nucleation, Aitken, accumulation, and coarse) and three insoluble modes (Aitken, accumulation, and coarse). Note that the new developments of this work are added to the model on top of the aerosols already represented by M7 and that the new aerosol components are introduced using the existing modes. Primary emissions of anthropogenic, biogenic, and biomass burning processes are defined through a variety of datasets; the most updated being those produced for the CMIP6 project. Natural emissions of mineral dust, sea salt, marine dimethyl sulfide 205 (DMS), and nitrogen oxides from lighting are calculated online, while other natural emissions are prescribed. Details on the
- various parameterizations used for the definition of the gas and aerosol emissions in the model can be found in van Noije et al. (2021).

# 2.2 The EC-Earth3-Iron model

EC-Earth3-Iron is the new version of the model developed and used for this work that builds on EC-Earth3-AerChem. The 210 new features required to determine the global aqueous-phase OXL formation, the atmospheric acidity, and the Fe cycle in the atmosphere can be summarized as:

 Treatment of mineral dust emission that considers soil mineralogical composition variations to account for the emission of Fe-containing minerals (and calcite), along with a detailed speciation of anthropogenic combustion and biomass burning emissions to explicitly account for Fe both in soluble and insoluble forms,

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2. Acidity calculations for water contained in fine and coarse aerosols, as well as, for cloud droplets,

3. A comprehensive aqueous phase chemistry scheme in cloud droplets and aerosol water, and

4. An explicit description of the Fe-containing aerosol dissolution processes of mineral dust, anthropogenic combustion and biomass burning aerosols.

#### 2.2.1 Speciated emissions

- 220 EC-Earth3-Iron includes a characterization of the dust mineralogical composition at emission and explicitly traces the Fe and calcium-containing species. The relative amounts of eight different minerals, namely: illite, kaolinite, montmorillonite, calcite, feldspars, quartz, gypsum, and hematite, are derived from the soil mineralogy atlas of Claquin et al. (1999), including the updates proposed in Nickovic et al. (2012). The atlas provides the soil mineralogical composition in arid and semi-arid regions of the world, distinguishing between two soil size classes (i.e., the clay-size fraction, up to 2 µm, and the silt-size fraction from
- 225 2 to 50 µm diameter). The mineral fractions emitted in the accumulation and coarse insoluble modes of TM5-MP are estimated from the soil mineralogy atlas based on the Brittle Fragmentation Theory (BFT) from Kok (2011). BFT posits that the emitted particle size distribution is independent of wind and soil conditions and additionally allows estimating the size-resolved mineral fractions (Pérez García-Pando et al., 2016; Perlwitz et al., 2015a, 2015b). The resulting mineral mass fractions are then applied to the dust emission fluxes, as calculated online in the model, yielding the corresponding accumulation and coarse mode
- 230 emission of each mineral. We note that although we derive the mineral dust fractions in each mode using BFT, we maintain the dependence of the ratio between the accumulation and coarse mode dust mass at emissions upon wind and soil conditions of the original dust emission scheme (Tegen et al., 2002).

In EC-Earth3-Iron, the different Fe-containing minerals are not prognostic variables (tracers). Instead, we trace the mineral dust Fe according to three dissolution classes, namely fast, intermediate, and slow Fe pools (Ito and Shi, 2016). <u>No relationship</u>

235 of Fe dissolution with other elements is observed, however, for clays and feldspars where the total Fe content of the minerals is very low (< 0.54 %) and the Fe is in the form of impurities (Journet et al., 2008). For this, 0.1 % Fe content in total Fe-containing minerals is here assumed directly soluble as amorphous free iron impurities regardless of mineralogy [Ito and Shi, 2016]. The emitted amounts of calcium (i.e., in calcite) and Fe (i.e., in illite, kaolinite, montmorillonite, feldspars, and hematite) are derived either from the average elemental compositions of minerals or based on experimental analyses (Journet et al., 2008; Nickovic et al., 2013). The respective average fractions applied to mineral dust sources of this work are listed in Table S1.</p>

Fe is also emitted in the model from anthropogenic activities (including fossil and biomass fuels) and biomass burning (excluding biofuel combustion) following J to et al. (2018), The Fe-containing fossil fuel and biofuel combustion emissions are here estimated by applying specific factors (i.e., per emission sector and per particle size) to the total particulate emissions (i.e., the sum of organic carbon, black carbon, and inorganic matter), as derived for this work based on estimates from J to et

245 al (2018), for the Fe content in the sub-micron and super-micron combustion aerosols. The historical anthropogenic emissions are taken here from the Community Emissions Data System (Hoesly et al., 2018) and the historical fire emissions from the BB4CMIP6 data set (van Marle et al., 2017). We note, however, that the estimate of Fe emission from metal smelting remain Deleted: (PSD)

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- 260 highly uncertain and further works are needed (Rathod et al., 2020). As for the biomass burning, the iron fractions in the fine particles are related to the combustion stages of flaming (0.46 ± 0.51 %) and smoldering (0.06 ± 0.03 %) fires, while the averaged iron fraction is used for coarse particles (3.4 %) (Ito, 2011), The global mean ratio of 0.04 gFe gBC<sup>-1</sup> for biomass burning in fine particles is consistent with that of 0.032 in the review paper by Hamilton et al. (2022), Fe-containing aerosol combustion emissions are considered to be insoluble (Ito, 2015), except for ship oil combustion that is assumed to be mostly
- 265 soluble, i.e., 79% on average for the years 2000-2014, We note that the value of 79% represents the high solubility of iron emissions in oil fly ash (Ito et al., 2021). Rathod et al. (2020) proposed a lower solubility in emissions (i.e., 47.5% for iron-sulfates), with an upper value, however, at ~90%. The year-to-year variation in anthropogenic combustion Fe-emission fractions follows Ito et al. (2018). On the contrary biomass burning no such a variation is provided. The average Fe fractions (per sector) for the years 2000 to 2014 applied to the total particulate carbonaceous emissions are also listed in Table S1.
- 270 EC-Earth3-Iron also includes OXL primary emissions from natural and anthropogenic wood-burning processes, that mainly account for its rapid formation in the sub-grid plumes not represented in the model. Indeed, OXL is well correlated with elemental carbon and levoglucosan (Cao et al., 2017; Cong et al., 2015), which are observed at significant levels during biomass burning episodes in the Amazon (Kundu et al., 2010), suggesting that oxalic acid could be either directly emitted or formed rapidly via combustion processes. During biomass burning episodes enhanced emissions of ionic species have been
- 275 generally measured, indicating an average OXL mass concentration measured in plumes of ~0.04-0.07 %\_w/w (Yamasoe et al., 2000). Furthermore, domestic wood combustion is a potential OXL source (Schmidl et al., 2008) since measurements indicate an OXL contribution to the total particulate concentrations of ~0.09-0.28 %\_w/w. Gasoline engines may also contribute to total dicarboxylic acid mass emitted to the atmosphere (Kawamura and Kaplan, 1987), although their direct contribution to ambient OXL concentrations is generally found to be low (Huang and Yu, 2007) and, therefore, neglected here. All in all,
- 280 primary OXL sources are quite uncertain and, given the current estimates, may only have a limited impact on the calculation of its atmospheric concentrations (e.g., Myriokefalitakis et al., 2011).

#### 2.2.2 Thermodynamic equilibrium and atmospheric acidity calculations

The gas/particle equilibrium calculations of  $NH_3/NH_4^+$  and  $HNO_3/NO_3^-$  have been substantially revised in EC-Earth3-Iron. In EC-Earth3-AerChem, EQSAM (Metzger et al., 2002) is used to determine the partitioning of  $NH_3/NH_4^+$  and  $HNO_3/NO_3^-$ . In

- 285 EC-Earth3-Iron, the ISORROPIA II thermodynamic equilibrium model (Fountoukis and Nenes, 2007), replaces EQSAM to determine the equilibrium between the inorganic gas and the aerosol phases. ISORROPIA-II calculates the gas/liquid/solid equilibrium partitioning of the K<sup>+</sup>-Ca<sup>2+</sup>-Mg<sup>2+</sup>-NH<sub>4</sub><sup>+</sup>-Na<sup>+</sup>-SO<sub>4</sub><sup>2-</sup>-NO<sub>3</sub><sup>-</sup>-Cl<sup>-</sup>-H<sub>2</sub>O aerosol system, and is used in the forward mode, assuming that all aerosols are in a metastable (liquid) state. The inclusion of sea salt and dust aerosols in the aerosol thermodynamic calculations has been shown, nevertheless, to substantially affect the ion balance and thus the partitioning of
- 290 HNO<sub>3</sub>/NO<sub>3</sub> and NH<sub>3</sub>/NH<sub>4</sub><sup>+</sup> species, especially in areas with abundant mineral dust and/or sea spray aerosols (Athanasopoulou et al., 2008, 2016; Karydis et al., 2016). In EC-Earth3-Iron nitrate aerosols are calculated for both the accumulation and coarse

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**Deleted:** The Fe-containing combustion emissions are estimated here from the total particulate carbon emissions (i.e., the sum of organic carbon and black carbon), based on the Fe content in the submicron and super-micron combustion aerosols. The scaling factors for each aerosol size (and solubility) are applied to the sectors energy industrial, iron and steel industries, residential and commercial shipping, and waste treatment, as well as to the biomass burning emissions...

**Moved up [1]:** The historical anthropogenic emissions are taken from the Community Emissions Data System (Hoesly et al., 2018) and the historical fire emissions from the BB4CMIP6 data set (van Marle et al., 2017).

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% for Na<sup>+</sup>, K<sup>+</sup>, and Mg<sup>2+</sup>, respectively (Karydis et al., 2016; Sposito, 1989). For sea spray aerosols, mean mass fractions of

55.0 % Clr, 30.6 % Na<sup>+</sup>, 7.7 % SO4<sup>2-</sup>, 3.7 % Mg<sup>2+</sup>, 1.2 % Ca<sup>2+</sup>, and 1.1% K<sup>+</sup> (Seinfeld and Pandis, 2006) are also applied.

The acidity levels of deliquescent aerosols are calculated in the model based on thermodynamic processes for accumulation and coarse particles. Aerosol acidity impacts the scavenging efficiency and the dry deposition of inorganic reactive nitrogen species due to changes in the partitioning of total nitrate and ammonium between the gas and aerosol phases as well as between

- 325 the various aerosol sizes (Pye et al., 2020). Acidity levels also play a fundamental role in the aqueous-phase chemistry by controlling the dissociation reactions and thus, the reactivity of the chemical mechanism. Indeed, aqueous-phase species, such as organic and inorganics acids, are oxidized with higher rates when they are dissociated. Nevertheless, in the case of the forward and reverse reactions, they typically occur fast and thus the concentrations of the reactants and the products are generally assumed to be in equilibrium in the global model due to its relatively long timestep and large model grid. Note,
- 330 however, that recent modeling studies showed that the metastable assumption produces pH values that are different from the stable assumption (e.g., regionally up to 2 pH units in the presence of crustal elements over dust sources, and roughly 0.5 pH units globally; Karydis et al., 2020), However, work to date, such as in Bougiatioti et al. (2016), Guo et al. (2018; 2015) and others identified in the review of Pye et al. (2020), has shown that the metastable solution tends to provide semi-volatile partitioning of pH-sensitive species (e.g., NH<sub>3</sub>/NH<sub>4</sub> and HNO<sub>3</sub>/NO<sub>3</sub>) and aerosol liquid water content that is closer to
- 335 observations at least for when the relative humidity is above 40  $\frac{9}{20}$ . For this reason, we assume that the most plausible estimates of acidity are to be obtained with the metastable assumption, and we base our simulations on that,

Under ambient atmospheric conditions, the water vapor uptake on aerosols depends on both the inorganic and organic components, along with the meteorological conditions (e.g., the temperature and the relative humidity conditions). ISORROPIA II does not, however, include water associated with organic aerosols, possibly leading to an underestimation of

340 the aerosol hygroscopicity especially within the boundary layer where the contribution of water-soluble organics to total aerosol mass can be substantial. For this, we account here for a contribution of aerosol water from organic particles in the acidity calculations, using a hygroscopicity parameter  $\kappa_{org}$ =0.15 (Bougiatioti et al., 2016). In more detail, the particulate water due to the organics (W<sub>org</sub>) that is added to the aerosol water associated with the inorganic aerosol as calculated from ISORROPIA-II (W<sub>inorg</sub>), is determined in the model as:

345 
$$W_{org} = m_s \cdot \frac{\rho_W}{\rho_s} \cdot \frac{\kappa_{org}}{(\frac{1}{RH} - 1)}$$

(1)

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where  $m_s$  is the soluble organic mass concentration ( $\mu g m^{-3}$ ) as simulated by the TM5-MP chemistry scheme,  $\rho_w$  is the water density (1 kg m<sup>-3</sup>),  $\rho_s$  is the organic aerosol density (1.4 kg m<sup>-3</sup>), and RH (0-1) is the relative humidity.

- 355 Cloud acidity is also an important factor for simulating the multiphase processes in the atmosphere. The in-cloud proton concentration is initially determined by the electroneutrality of strong acids and bases (i.e., H<sub>2</sub>SO<sub>4</sub>, SO<sub>4</sub><sup>2-</sup>, methanesulfonate (MS<sup>-</sup>), HNO<sub>3</sub>, NO<sub>3</sub><sup>-</sup>, and NH<sub>4</sub><sup>+</sup>), and then the subsequent dissociations of CO<sub>2</sub>, SO<sub>2</sub>, and NH<sub>3</sub> (Jeuken et al., 2001) are solved iteratively in the model. For the cloud acidity calculations, the liquid water content, and the respective cloud cover fraction (i.e., 0-1) are obtained from meteorology. Note, however, that the effect of mineral dust, and especially calculation, on cloud
- 360 proton concentrations is <u>here</u> neglected. This assumption may result in some overestimation of cloud acidity, although the overall impact should be small particularly in dusty areas with a low presence of clouds. Another limitation in the determination of cloud acidity is the omission of light gaseous organic acids (such as formic and acetic acids), possibly leading to some underestimation in cloud acidity where their concentration is important.

#### 2.2.3 The aqueous phase chemistry scheme

- 365 The aqueous-phase chemistry scheme used in this work is based to a large extent on the Chemical Aqueous Phase Radical Mechanism (CAPRAM) (e.g., Deguillaume et al., 2004; Ervens et al., 2003; Herrmann et al., 2000, 2015). However, CAPRAM includes more than 70 aqueous-phase species, 34 equilibria for compounds that are present both in the gas and the aqueous phases, along with numerous photolytic and aqueous-phase reactions, also covering a large series of acid-base and metalcomplex equilibria. Note that various updates may further extend the mechanism by including among others, the oxidation of
- 370 aromatic hydrocarbons (Hoffmann et al., 2018), the multiphase oxidation of DMS (Hoffmann et al., 2016), and the tropospheric multiphase halogen chemistry (Bräuer et al., 2013). For this, some reactions are considered here in a more simplified way based on various assumptions published in literature. Indeed, the level of chemical complexity of such a detailed mechanism is beyond the computational resources available for three-dimensional global climate-chemistry simulations, and thus simplifications, that preserve however the essential features of the aqueous mechanism, are needed.
- 375 Aqueous-phase chemical transformations are considered at the interface and in the bulk, initiated mainly by free radicals and oxidants produced both via photochemical reactions and in dark conditions (Bianco et al., 2020). The sources of OH radicals in the aqueous phase, however, strongly differ from those in the gas phase, primarily because of the presence of ionic species and TMIs in the solution. OH radicals are the main oxidant in the aqueous phase, either produced directly in the aqueous medium or diffused from the gas phase (i.e., via a gas-to-liquid transfer). However, aqueous phase oxidation can also be
- 380 induced by non-radical species, such as ozone (O<sub>3</sub>) and hydrogen peroxide (H<sub>2</sub>O<sub>2</sub>). A characteristic example is the formation of SO<sub>4</sub><sup>2-</sup> in cloud droplets, via the oxidation of dissolved sulfur dioxide (SO<sub>2</sub>) by O<sub>3</sub> and H<sub>2</sub>O<sub>2</sub>, with H<sub>2</sub>O<sub>2</sub> being nevertheless the most effective oxidant (Seinfeld and Pandis, 2006), especially when the solution becomes acidic. Upon the absorption of SO<sub>2</sub> in cloud droplets, the establishment of the equilibrium between the dissolved sulfur species in oxidation state 4, i.e.,
  - 11

 $SO_2.H_2O$ ,  $HSO_3^{-}$  (pKa<sub>1</sub> = 1.9), and  $SO_3^{2-}$  (pKa<sub>2</sub> = 7.2) (hereafter also as S(IV)) is calculated in the model. Then, depending on the availability of oxidants and the solution's acidity, the different S(IV) species can participate in the formation of S(VI) (i.e., dissolved sulfur in oxidation state 6).

In EC-Earth3-Iron, the aqueous-phase sulfur scheme is applied both in cloud droplets and aerosol water, replacing the S(VI) production through the dissolved S(IV) oxidation in cloud droplets previously included in the EC-Earth-AerChem (van Noije et al., 2014, 2021). In more detail, besides the two classic reactions of bisulfite and sulfite with hydrogen peroxide and ozone

- 390 included in EC-Earth3-AerChem, additional reactions of S(IV) oxidation via methyl hydroperoxide (CH<sub>3</sub>O<sub>2</sub>H), peroxyacetic acid, and with the hydroperoxyl radical (HO<sub>2</sub>)/superoxide radical anion (O<sub>2</sub><sup>-</sup>) are considered. Nevertheless, in acidic solutions, the oxidation by peroxides, and especially H<sub>2</sub>O<sub>2</sub>, is significantly more important than other oxidants (Herrmann, 2003; Jacob, 1986). H<sub>2</sub>O<sub>2</sub> is produced in the gas phase and can be rapidly dissolved in the liquid phase due to its high solubility. The dissolved H<sub>2</sub>O<sub>2</sub> (as well as the organic peroxides, such as CH<sub>3</sub>OOH) can react rapidly with the HSO<sub>3</sub><sup>-</sup>. However, the pH-
- 395 independent reaction of HSO<sub>3</sub><sup>-</sup> with CH<sub>3</sub>OOH (or other organic peroxides) is expected to be less important than H<sub>2</sub>O<sub>2</sub> under typical cloud conditions due to the much lower solubility of CH<sub>3</sub>OOH. Note that the dissociation of H<sub>2</sub>O<sub>2</sub> is here neglected since it is not expected to significantly influence the total H<sub>2</sub>O<sub>2</sub> concentrations under typical tropospheric conditions (Herrmann, 2003; Jacob, 1986). In contrast, at a higher pH, the S(IV) oxidation by ozone tends to dominate the S(IV) oxidation (Seinfeld and Pandis, 2006). O<sub>3</sub> oxidizes rapidly all three S(IV) forms in the aqueous phase, becoming significant at pH higher
- 400 than 4 (Seinfeld and Pandis, 2006), even in the absence of light. S(IV) oxidation by O<sub>3</sub> is also predicted to dominate S(VI) formation during winter in arctic regions due to the lack of photochemical production of OH and H<sub>2</sub>O<sub>2</sub> at high latitudes, as well as to the high anthropogenic SO<sub>2</sub> emissions in the Northern Hemisphere (Alexander et al., 2009). Laboratory studies indicate that S(IV) compounds may be also oxidized in the aqueous phase via other pathways. For example, the aqueous S(VI) production can be enhanced by TMIs (Harris et al., 2013), such as the Mn(II) catalyzed oxidation of S(IV) by dissolved O<sub>2</sub>. In
- a global modeling study, Alexander et al. (2009) attributed 9-17\_% of the total S(VI) production to the latter mechanism. However, such reactions would require several oxysulfur radicals as intermediates (e.g., Deguillaume et al., 2004; Herrmann et al., 2005), like a free radical chain mechanism initiated by reactions of HSO<sub>3</sub><sup>-</sup>, SO<sub>3</sub><sup>2-</sup> with radicals and radical anions, or TMIs catalyzed oxidation of several S(IV)-compounds, which is not considered in our model. Thus, in the case of the sulfate radical anion (SO<sub>4</sub><sup>-</sup>) production via the Fe(III)-sulfato-complex [Fe(SO<sub>4</sub>)]<sup>+</sup> photolysis (Table S2), the sulfate radical anion is simply added to the S(VI) pool.

Gas-phase organics can be also oxidized in the interstitial cloud space, form water-soluble compounds like aldehydes, and rapidly partition into the droplets. In the presence of oxidants such as OH and NO<sub>3</sub> radicals in the solution, the dissolved organics undergo chemical conversions and form low volatile organics that remain, at least partly, in the particulate phase upon droplet evaporation (Blando and Turpin, 2000). The dissolved OH radicals react with organic compounds in the aqueous

- 415 phase by hydrogen abstraction or electron transfer, forming alkyl radicals (R) which, in the presence of dissolved oxygen,
  - 12

further form peroxyl radicals (RO<sub>2</sub>). The OH oxidation of organic compounds in the aqueous phase can lead either to fragmentation or, to the formation of oxidized organic species, resulting overall in CO2. However, the recombination of organic radicals can also be a favorable pathway when the water evaporates, and the aqueous solution becomes more concentrated. Box-model simulations have shown that the cloud processing of polar products from isoprene oxidation can be an important

- 420 contributor to secondary organic aerosol (SOA) production (Lim et al., 2005). Indeed, laboratory measurements show that the aqueous-phase photooxidation of C2 and C3 carbonyl compounds (Perri et al., 2009, 2010), such as glyoxal (Carlton et al., 2007, 2009), methylglyoxal (Altieri et al., 2008), glycolaldehyde, pyruvic acid (Carlton et al., 2006), and acetic acid (Tan et al., 2012) leads to the production of low volatility DCAs, which are commonly found in atmospheric aerosols and clouds (Sorooshian et al., 2006).
- 425 In EC-Earth3-Iron, gas-phase species can be reversibly transferred to the aqueous phase and oxidized by radicals and radical anions. The partitioning of 15 organic species that exist in both phases are considered in the aqueous-phase mechanism, namely methyl-peroxy radical (CH3O2), methyl hydroperoxide (CH3O2H), formaldehyde (HCHO), methanol (CH3OH), formic acid (HCOOH), acetaldehvde (CH3CHO), glvcolaldehvde (GLYAL: HOCH2CHO), glvcoal (GLY: CH(O)CH(O)), ethanol (CH<sub>3</sub>CH<sub>2</sub>OH), acetic acid (CH<sub>3</sub>COOH), methylglyoxal (MGLY; CH<sub>3</sub>C(O)CHO), hydroxyacetone (HYAC; CH<sub>3</sub>C(O)CH<sub>2</sub>OH),
- 430 pyruvic acid (PRV; CH<sub>3</sub>C(O)COOH), GLX, and oxalic acid (H<sub>2</sub>C<sub>2</sub>O<sub>4</sub>). The aqueous phase oxidation is taking place by the OH and NO3 radicals, as well as the CO3<sup>-</sup> radical anion. OH is either produced by photolytic reactions of dissolved compounds or via a direct transfer from the gas phase into the solution, as well as by Fenton reaction (Deguillaume et al., 2010). NO3 radicals are transferred from the gas phase, while the CO3- radical anion is produced mainly via the oxidation of hydrated CO2. In general, the aqueous phase oxidation largely proceeds via OH radicals, followed by NO3 radicals under dark conditions, while 435 the CO<sub>3</sub><sup>-</sup> radical has an overall small impact on the oxidizing capacity of the solution.

Upon their transfer to the solution, aldehydes are considered to be in equilibrium with the corresponding diols. The hydrated aldehydes are oxidized via H-atom abstraction with radicals (OH, NO<sub>3</sub>) or radical anions (CO<sub>3</sub><sup>-</sup>), followed by the elimination

oxidized via an H-atom abstraction; the resulting α-hydroxy-alkyl radicals, however, are not explicitly resolved, but the direct 440 formation of aldehydes (e.g., formaldehyde and acetaldehyde) is considered via the respective peroxyl radical reactions with molecular oxygen to yield HO2. Moreover, the glycolic acid (HOCH2COOH) production via glycolaldehyde oxidation is not also explicitly described in the aqueous-phase scheme, and only the direct production of GLX is considered (Lin et al., 2012; Myriokefalitakis et al., 2011). This assumption is expected to have a negligible impact on the overall chemical mechanism since the glycolic acid is rapidly oxidized into glyoxylic acid with its net in-cloud production being rather small (Liu et al.,

of HO2 in reaction with O2, leading overall to the formation of organic acids. Alcohols, such as CH3OH and C2H3OH, are also

445 2012).

> After cloud evaporation, OXL and SO4<sup>2-</sup> are considered to reside entirely in the particulate phase of the model. This approximation may, nevertheless, result in an overestimate of OXL (pKa1 = 1.23; (COO')2, pKa2 = 4.19) concentrations, since

low levels of gas-phase oxalic acid have been also observed in the atmosphere under favorable conditions (e.g., Baboukas et al., 2000; Martinelango et al., 2007). Note that other products, such as pyruvate, glyoxylate<sub>a</sub> as well as the oligomers from GLY and MGLY, are also considered to reside in the particulate phase upon cloud evaporation (Lim et al., 2005; Lin et al., 2012; Liu et al., 2012) and are thus added directedly to the SOA pool of the model. However, in contrast to OXL and the low-volatility oligomers, the pyruvic and glyoxylic acids are allowed to be partially transferred back to the gas phase of the model when the cloud droplets evaporate.

For the present work, the aqueous reaction rate coefficients are taken (where available) from the available literature of the
CAPRAM schemes and supplemented with reaction rates from laboratory and modeling studies (i.e., Carlton et al., 2007; Deguillaume et al., 2009; Lim et al., 2005; Sedlak and Hoigné, 1993). For the sulfur chemistry, the aqueous reaction rates are taken from Seinfeld and Pandis (2006). In the case of missing experimental data for temperature dependencies, the rate constants for T = 298 K are only applied in chemistry calculations. O<sub>3</sub>, H<sub>2</sub>O<sub>2</sub>, NO<sub>3</sub>, HONO/NO<sub>2</sub><sup>-</sup>, HNO<sub>3</sub>/NO<sub>3</sub><sup>-</sup>, and CH<sub>3</sub>O<sub>2</sub>H, along with Fe<sup>3+</sup>, [Fe(SO4)]<sup>+</sup> and [Fe(OXL)<sub>2</sub>]<sup>-</sup> are photolyzed in the aqueous phase. Aqueous photolysis frequencies (where available) are taken from the gas-phase chemistry. For Fe-species (e.g., Fe<sup>3+</sup>, [Fe(SO4)]<sup>+</sup>, [Fe(OXL)<sub>2</sub>]<sup>-</sup>), their maximum (i.e., noontime at 51° N) photolysis frequencies, as proposed by Ervens et al. (2003), are scaled based on the gas-phase H<sub>2</sub>O<sub>2</sub> photolysis rates. A list of all aqueous and photochemical reactions included in the chemical scheme of this study is presented in Table S2, with the respective equilibrium reactions shown in Table S3.

#### 2.2.4 The iron solubilization scheme

465 A three-stage kinetic approach (Shi et al., 2011) is applied to describe the solubilization of the Fe-containing dust mineral pools (Ito and Shi, 2016), representing: 1) a rapid dissolution of ferrihydrite on the surface of minerals (i.e., fast pool), 2) an intermediate stage dissolution of nano-sized Fe oxides from the surface of minerals (i.e., intermediate pool), and 3) the Fe release from heterogeneous inclusion of nano-Fe grains in the internal mixture of various Fe-containing minerals, such as aluminosilicates, hematite, and goethite (i.e., slow pool). A separate Fe pool for combustion aerosols (Ito, 2015) is also considered in the model.

470 considered in the model.

The dissolved Fe in the model is produced via dissolution processes in aerosol water and cloud droplets depending on the acidity levels of the solution (i.e., proton-promoted dissolution scheme), the OXL concentration (i.e., ligand-promoted dissolution scheme), and irradiation (photo-reductive dissolution scheme), following Ito (2015) and Ito and Shi (2016). The Fe release from different types of minerals depends, thus, on the solution acidity (pH) and the temperature (T), as well as on the degree of solution saturation. In more detail, the dissolution rates for each of the three dissolution processes considered can be empirically described (e.g., Ito, 2015; Ito and Shi, 2016; Lasaga et al., 1994) as:

14

$$RFe_i = K_i(pH, T) \cdot \alpha(H^+)^{m_i} \cdot f_i \cdot g_i$$

(2)

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480 where  $K_i$  (moles Fe gi<sup>-1</sup> s<sup>-1</sup>) is the Fe release rate due to the dissolution process *i*,  $\alpha(H^+)$  is the H<sup>+</sup> activity of the solution, and  $m_i$  is the empirical reaction order for protons derived from experimental data the functions  $f_i$  and  $g_i$  represent the suppression of the different dissolution rates due to the solution saturation state as:

$$f_i = 1 - (a_{Fe^{3+}} \cdot a_{H^+}^{-n_i})/K_{eq_i}$$
$$g_i = 0.17 \cdot ln(\frac{a_{OXL}}{a_{H^+}}) + 0.63$$

$$q_i = 0.17 \cdot ln(\frac{a_{OXL}}{a_{E_0}^{3+}}) + 0.4$$

485 where,  $\alpha_{H+}$ ,  $\alpha_{Fe+3}$ , and  $\alpha_{OXL}$  stand for the solution's activities of 1) protons, 2) ferric cations, and 3) OXL, respectively, as calculated each time step in the model, and Keqi (mol<sup>2</sup> kg<sup>-2</sup>) is the equilibrium constant. The activation energy that accounts for the temperature dependence is derived as a function of acidity based on soil measurements (Bibi et al., 2014; Ito and Shi, 2016), i.e.:

$$E_{pH} = -1.56 \ 10^3 \cdot pH + 1.08 \ 10^4$$

490 Overall, the net Fe dissolution rate results from the sum of the three rates. All parameters used for the calculation of dissolution rates for this work are presented in Table S4.

# 2.3 The chemistry solver

All concentrations of gas, aqueous, and aerosol species evolve dynamically in the model. The ordinary differential equations that govern the production and destruction terms due to chemical reaction and interphase mass transfer in the model are:

$$\frac{dG}{dt} = R_{G} - LWCk_{mt}G + \frac{K_{mt}}{HRT}A$$

$$\frac{dA}{dt} = R_{A} + LWCk_{mt}G - \frac{K_{mt}}{HRT}A$$
(6)

where,

G = Gas-phase concentrations (molecules cm<sup>-3</sup> of air)

A = Aqueous-phase concentrations (molecules cm<sup>-3</sup> of air)

500  $R_G$  = Gas-phase reaction terms (molecules cm<sup>-3</sup> of air s<sup>-1</sup>)

 $R_A =$  Aqueous-phase reaction terms (molecules cm<sup>-3</sup> of air s<sup>-1</sup>)

LWC = Liquid water content ( $cm^3$  of water  $cm^{-3}$  of air)

kmt = Mass transfer coefficient (s<sup>-1</sup>)

H = Henry's Law coefficient (moles L<sup>-1</sup> atm<sup>-1</sup>)

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<b>Deleted:</b> The net Fe dissolution rate results from the sum o three rates	f the

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(5)

(3) (<u>4</u>)

#### R = Ideal gas constant (L atm mol<sup>-1</sup> K<sup>-1</sup>)

T = Temperature (K)

The mass transfer between the gas and aqueous phases (Lelieveld and Crutzen, 1991; Schwartz, 1986) is applied only for those species that exist in both phases and is represented in the mechanism by two separate reactions, i.e., one reaction for transfer from the gas to the aqueous phase and one for the transfer from the aqueous to the gas phase. All Henry's law solubility constants (H) used in this work are taken from Sander (2015) and are presented in Table S5.

The mass transfer coefficient (k<sub>mt</sub>) for a species is calculated as:

$$k_{mt} = \left(\frac{r^2}{3D_g} + \frac{4r}{3\nu\alpha}\right)^{-1}$$

where *r* is the effective droplet or aqueous aerosol radius (m),  $D_g$  is the gas-phase diffusion coefficient (m<sup>2</sup> s<sup>-1</sup>), *v* the mean molecular speed (m s<sup>-1</sup>), and *α* the mass accommodation coefficient (dimensionless).  $v = \sqrt{\left(\frac{8R_gT}{\pi M_w}\right)}$  The cloud droplet effective radius may vary between ~3.6 - 16.5 µm for remote clouds, 1 - 15 µm for continental clouds, and ~1 - 25 µm for polluted clouds (Herrmann, 2003). For this work, the effective radius of cloud droplets (ranging between 4-30 µm in the model) is calculated online based on the cloud liquid water content and the cloud droplet number concentration (van Noije et al., 2021).

530 The effective radii (i.e., the ratio of the third to the second wet aerosol moments) for the accumulation and coarse deliquescence particles, are based on the respective M7 calculations. According to Eq. (8), the gas transfer to small droplets is faster, owing to the larger surface-to-volume ratio of smaller droplets. However, sensitivity model simulations using different droplet radii showed that varying droplet sizes result only in small changes in the chemical production of aqueous-phase species (Lelieveld and Crutzen, 1991; Liu et al., 2012; Myriokefalitakis et al., 2011).

535 The mean molecular speed of a gaseous species is calculated as:

$$v = \sqrt{\left(\frac{8R_gT}{\pi M_w}\right)}$$

where,  $M_W$  is the respective molecular weight (kg mol<sup>-1</sup>) and  $R_g$  is the ideal gas constant (J mol<sup>-1</sup>  $K^{-1}$ ) (Herrmann et al., 2000). The D<sub>g</sub> and  $\alpha$  used for this study are also presented in Table S5.

KPP version 2.2.3 (Damian et al., 2002; Sandu and Sander, 2006) was used to generate the Fortran 90 code for the numerical integration of the aqueous-phase chemical mechanism. For this, a separate model driver was developed to arrange the respective couplings to the TM5-MP I/O requirements (e.g., species that partition in the aqueous-phase, the reaction and dissolution rates, and the photolysis coefficients). The Rosenbrock solver is used in this work as the numerical integrator, since it is found to be rather robust and capable of integrating very stiff sets of equations (Sander et al., 2019). However, as for the case of the gas-phase mechanism's coupling (Myriokefalitakis et al., 2020b), minor changes were needed to be applied in the

16

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(8)

(9)

**Moved down [2]:**  $D_g$  and  $\alpha$  used for this study are also presented in Table S5. The mean molecular speed of a gaseous species is calculated as: ¶

 $\upsilon = \left[\overline{\left(\frac{8R_gT}{\pi M_w}\right)} \rightarrow (9)\right]$ 

where,  $M_W$  is the respective molecular weight (kg mol<sup>-1</sup>) and  $R_g$  is the ideal gas constant (J mol<sup>-1</sup>  $K^{-1}$ ) (Herrmann et al., 2000).

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original KPP code. For instance, the aqueous and photolysis reactions are not calculated inside KPP, but directly provided
 through calculations in the aqueous chemistry driver. In contrast, for the Fe dissolution scheme, the suppressions of the mineral dissolution rates due to the solution saturation are calculated online by KPP (see Eq. 3 and Eq. 4).

#### 2.4 Simulations

We performed a range of present-day simulations, including experiments using EC-Earth3-Iron atmosphere-only runs (hereafter referred to as EC-Earth), and TM5-MP standalone driven by ERA-Interim (Dee et al., 2011) reanalysis fields

- (hereafter referred to as ERA-Interim), covering the period 2000-2014. For the EC-Earth simulation, TM5-MP is coupled to the IFS atmospheric dynamics. We used prescribed sea surface temperature and sea-ice concentration fields from a set of input files through the AMIP interface (Taylor et al., 2000). Thus, for the atmosphere and chemistry modules, our setup follows the EC-Earth3-AerChem standard configuration in CMIP6 experiments. The IFS horizontal resolution is T255 (i.e., a spacing of roughly 80 km), 91 layers are used in the vertical direction up to 0.01 hPa, and a time step of 45 min is applied. Respectively, 565 TM5-MP (both for the online and offline configurations) has a horizontal resolution of 3° in longitude by 2° in latitude and 34
- layers in the vertical direction up to 0.1 hPa (~60 km).

The ERA-Interim setup allows constraining the model with the assimilated observed atmospheric circulation data and is therefore used for budget analysis and comparison with other estimates from the literature. ERA-Interim is further used to explore uncertainties regarding the aqueous-phase chemistry scheme. Specifically, an additional simulation is performed to

- 570 identify the potential importance of glyoxal-derived oligomers and high molecular weight species in the aqueous phase (Carlton et al., 2007) on the OXL production rates and the respective ambient concentrations. In this sensitivity simulation, (hereafter referred to as ERA-Interim(sens)), the OXL formation via high molecular weight species formation from glyoxal oxidation is neglected. Comparisons between the corresponding 15-year climatologies from the EC-Earth and ERA-Interim simulations are used to identify uncertainties in the aqueous-phase production terms of OXL, the iron-dissolution rates, and
- 575 finally the atmospheric concentrations and deposition rates of Fe-containing aerosols due to the applied meteorology (i.e., online vs. offline). Note that the same emission datasets are used both in the ERA-Interim driven and the EC-Earthexperiments, thus only natural primary sources depending on meteorology may differ (see Sect. 2.1). A summary of the simulations is listed in Table 1.

# 2.5 Observations

- 580 A general evaluation of the modeled Aerosol Optical Depth (AOD) at 550 nm allows for characterizing EC-Earth3-Iron's ability to reproduce the aerosol fields. The Aerosol Robotic Network (AERONET) version 3 (Giles et al., 2019) level 2.0 direct sun retrievals at a monthly basis are used to calculate annual mean AOD values for the 2000-2014 period. The model's coarse horizontal resolution hinders, however, the representation of high-altitude locations, thus, following Huneeus et al. (2011), we exclude sites above 1000 m asl., leaving 738 locations with information available during the simulated period. In addition, we
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perform a specific evaluation of mineral dust, which constitutes a key modulator of the outcome of our new developments as a source of Fe and Ca. To that end, we apply two additional filters to the AERONET data mentioned above, also following

- 590 Huneeus et al. (2011), to identify dust-dominated sites. First, we exclude those sites where the monthly mean Angstrom exponent is above 0.4 more than 2 months in the selected period. To further discriminate dust from sea salt, a minimum threshold of 0.2 for AOD at 550 nm is considered (i.e., if more than half of the retrieved AOD is above that threshold, the site is considered as dust-dominated). This filtering allows identifying a subset of stations potentially dominated by dust aerosols, however, it cannot ensure that three is no influence of other aerosol types in the monthly retrievals. Therefore, the evaluation
- 595 of AOD at 550 nm at those sites is taken as a proxy for the dust optical depth, acknowledging that other aerosols may also be present.

Pure dust measurements of surface concentration and deposition complement our evaluation of the model. The modeled annual mean surface dust concentration for <u>the years</u> 2000-2014 is compared to climatological observations from the Rosenstiel School of Marine and Atmospheric Science (RSMAS) of the University of Miami (Arimoto et al., 1995; Prospero, 1996, 1999;

- 600 Prospero et al., 1989) and the African Aerosol Multidisciplinary Analysis (AMMA) international program (Marticorena et al., 2010) observations. The 23 available sites cover locations close to sources (e.g., the AMMA stations over the Sahelian dust transect), in transport regions (e.g., stations from RSMAS on the Atlantic), and remote regions (e.g., RSMAS sites close to Antarctica). The modeled dust deposition fluxes are compared to the compilation of observations for the modern climate in Albani et al. (2014) including measurements at 110 locations, the mass fraction for particles with diameter lower than 10 μm 605 is used to keep the observed mass fluxes within the range of the modeled sizes.
- is used to keep the observed mass nuxes within the range of the modeled sizes

The simulated OXL and SO<sub>4</sub><sup>2</sup> concentrations are compared against measurements for representative sites, such as the Eastern Mediterranean (Finokalia, Greece; Koulouri et al., 2008), central Europe (Puy de Dome, France; Legrand et al., 2007), and the North Atlantic Ocean (Azores, Portugal; Legrand et al., 2007). Simulated monthly mean surface concentrations of OXL are also compared against a range of observations (n=143) from remote sites around the world as compiled in Myriokefalitakis et

- 610 al. (2011). Moreover, SO<sub>4</sub><sup>2-</sup> monthly mean surface concentrations over Europe and the USA are also compared against observations (n=3828) obtained from the European Monitoring and Evaluation Programme (EMEP; <u>http://www.emep.int;</u> last access 11/06/2021) and the Interagency Monitoring of Protected Visual Environments (IMPROVE; <u>http://vista.cira.colostate.edu/improve/;</u> last access 11/06/2021), respectively, as compiled in Daskalakis et al. (2016). The simulated Fe-containing aerosol concentrations are evaluated against cruise measurements covering a period from late 1999
- 615 up to early 2015, as compiled by Myriokefalitakis et al. (2018) and Ito et al. (2019a), and include daily observations for fine, coarse, and the total suspended particles.

Statistical parameters are here used to demonstrate the model's ability to represents atmospheric observations. These are the correlation coefficient (R) that reflects the strength of the linear relationship between model results and observations (i.e., the ability of the model to simulate the observed variability), the normalized mean bias (nMB), and the normalized root mean

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square error (nRMSE) as a measure of the mean deviation of the model from the observations due to random and systematic errors. The equations used for the statistical analysis of model results are provided in the Supplement (Eq. S1–S3) and the

625 locations (and regions) of the various observations used for evaluating the model for this work are presented in Fig. 1. Overall, a summary of statistics used for the evaluation of model simulations with the observations is presented in the Appendix (see Tables A1 and A2).

# 2.6 Model performance

The coupling of the aqueous-phase chemistry scheme along with the description of the atmospheric iron cycle for this work
 increases the model runtime. EC-Earth3-Iron uses here 109 transported and 33 non-transported tracers which are significantly larger numbers than in the EC-Earth3-AerChem configuration (i.e., 69 transported and 21 non-transported tracers). Note, however, that EC-Earth3-Iron used for this work employs the MOGUNTIA gas-phase chemistry scheme configuration, in contrast to the modified Carbon Bond Mechanism 2005 (mCB05) configuration (Huijnen et al., 2010; Williams et al., 2013, 2017) used in EC-Earth3-AerChem, which is overall found to be computationally more expensive by ~27 % alone
 (Myriokefalitakis et al., 2020b). In the Marenostrum4 supercomputer architecture (2x Intel Xeon Platinum 8160 24C at 2.1 GHz), the EC-Earth3-AerChem configuration (van Noije et al., 2021) simulates 1.85 years per day simulation time (SYPD) with 187 CPUs, while to reach to a comparable performance (i.e., 1.41 SYPD) with the EC-Earth3-Iron configurations, 432 CPUs are required. This means, overall, that the EC-Earth3-Iron corresponds to 7353 computation hours per year (CHPY), which is roughly 3 times larger than the standard EC-Earth3-AerChem.

#### 640 3 Results

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# 3.1 Budget calculations

The chemical production and destruction terms of OXL and its precursors, along with the Fe-containing aerosols' dissolution rates from combustion (FeC) and mineral dust (FeD), their emissions, and their removal terms from the atmosphere are presented for EC-Earth and ERA-Interim model configurations in this section. Additionally, we discuss differences compared to sensitivity simulations. Due to the common formation pathways of  $SO_4^{2-}$  and OXL in the atmosphere, the  $SO_4^{2-}$  budget calculations are also presented and discussed. All calculations are presented as a mean ( $\pm$  standard error) for the years 2000-2014.

#### 3.1.1 Oxalate

The annual net chemistry production of OXL (Table 2a) in EC-Earth is 12.615 ± 0.064 Tg yr<sup>1</sup>, which is lower than in ERA-650 Interim (18.116,± 0.071 Tg yr<sup>1</sup>). The difference is explained by a higher oxidizing capacity in ERA-Interim than in EC-Earth. ERA-Interim calculates higher OH concentrations in the tropical and subtropical troposphere (Fig. S1b). In contrast, zonal

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mean OH levels in EC-Earth are slightly higher in the extratropics, causing a more efficient oxidation of the OXL precursors, such as GLY (Fig. S1d), GLYAL (Fig. S1f), MGLY(Fig. S1h), and CH<sub>3</sub>COOH (Fig. S1j) at higher latitudes, especially in the

- 655 SH. Note that, van Noije et al. (2014) also showed that in large parts of the troposphere, the simulated oxidizing capacity in the previous version of EC-Earth (EC-Earth v2.4) was lower compared to a respective ERA-Interim configuration, due to the simulated lower temperatures (cold biases) and specific humidities. However, since SSTs and sea-ice concentrations are prescribed in our EC-Earth atmosphere-only simulations, the long-term means of tropospheric temperatures and water vapor are not expected to differ significantly to ERA-Interim close to the surface levels, as also indicated by the low differences in the OH levels of the two simulations at low altitudes (Fig. S1b). 660

The production term of OH in EC-Earth is ~5 % lower than in ERA-Interim, due to a lower amount of water vapor available to react with O(<sup>1</sup>D). In addition, a ~6 % lower OH production through the H<sub>2</sub>O<sub>2</sub> photolysis is simulated in EC-Earth. Note, that H<sub>2</sub>O<sub>2</sub> is an important driver of the aqueous-phase oxidizing capacity in the model, with about 80,% of the OH radicals in the liquid phase being produced by photolysis of the dissolved H2O2. In more detail, the lower atmospheric abundance of the gasphase  $H_2O_2$  in EC-Earth (~11 %) leads to smaller  $H_2O_2$  uptake in the aqueous phase (~13 %), and thus to a slower oxidation of OXL precursors due to the respective lower dissolved OH radical production (~19 %). Overall, the total OH production is ~7 % lower in EC-Earth, which corresponds to a ~18 % lower aqueous-phase OH production, resulting in a ~30 % lower OXL

net chemistry production compared to ERA-Interim.

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The total OXL production is 15.5 Tg yr<sup>1</sup> in Lin et al. (2014) and 14.5 Tg yr<sup>1</sup> in Liu et al. (2012), both lower compared to our

- 670 ERA-Interim estimates (20.8 Tg yr<sup>1</sup>; Table 2a), but close to EC-Earth (15.0 Tg yr<sup>1</sup>; Table 2a). The main reason for the lower chemistry production of other published estimates compared to our results is the contribution of the aqueous-phase glyoxal oxidation scheme proposed by Carlton et al. (2007) that is applied in our simulations. The oxidation of the glyoxal-derived high molecular weight products formed mainly in the cloud droplets is calculated to contribute significantly to the global OXL production in our model (Table 2a). This result is in line with Carlton et al. (2007), who indicated that the GLX-pathway may
- 675 not be the primary pathway for oxalic acid formation, but instead the rapid oxidation of GLY-multifunctional products via the OH radicals (i.e., 3.1 10<sup>10</sup> L mol<sup>-1</sup> s<sup>-1</sup>; Table S2). For ERA-Interim(sens) however, where no such reactions are considered, the total OXL chemical production is calculated on average 11.5 Tg yr<sup>-1</sup> (Table 2); i.e., closer to the estimates of Lin et al. (2014) and Liu et al. (2012). On the other hand, our ERA-Interim net chemistry production calculations are close to the estimates of Myriokefalitakis et al. (2011) (i.e., 21.2 Tg yr<sup>1</sup>) when no potential effects of the ionic strength (e.g., Herrmann,
- 680 2003) on OXL precursors are considered, although still lower since no Fe chemistry was considered in that latter study. Indeed, the enhanced aqueous-phase oxidation capacity due to the Fenton reaction increases both the production and the destruction terms of OXL in our model, leading to ~7 % lower net OXL production and, respectively, a lower (~8 %) atmospheric abundance. Nonetheless, our calculations indicate that Fe chemistry impacts on OXL net production drastically, increasing by at least ~50 % the destruction of the dissolved oxalic acid. The potential primary sources  $(0.373 \pm 0.005 \text{ Tg yr}^{-1})$  accounted

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690 for in the model (Table 2a) do not, however, significantly contribute to the simulated OXL atmospheric levels, and only a small fraction of OXL is calculated to be formed in aerosol water ( $\sim$ 6 %) for all simulations in this work.

Focusing further on the atmospheric sinks of OXL, roughly  $\frac{1}{3}$ % in ERA-Interim and  $\frac{1}{6}$ % in EC-Earth of the produced oxalic acid is oxidized into CO<sub>2</sub> in the aqueous phase, mainly via the photolysis of the [Fe(OXL)<sub>2</sub>]<sup>-</sup> complex (~55 %) and via OH radicals (~45 %). The fraction of the total produced OXL that is destroyed in the aqueous-phase is higher than in Liu et al.

- 695 (2012) by ~7 %, where no Fe chemistry was considered, but lower compared to Lin et al. (2014) and Myriokefalitakis et al. (2011), where roughly 30 % of the produced OXL is oxidized into CO<sub>2</sub> in the aqueous phase. Finally, a total average deposition rate of 18.5 Tg yr<sup>-1</sup> is calculated in ERA-Interim, primarily due to wet scavenging (~99 %), resulting in a global atmospheric lifetime of 5.7 days, close to Liu et al. (2012) and Lin et al. (2014), but higher compared to Myriokefalitakis et al. (2011) (~3 days) probably because the more intense OXL production at higher altitudes in our model.
- 700 The major pathways of global OXL production, both in ERA-Interim and EC-Earth, are the oxidation of glyoxal (~74 %), followed by glycolaldehyde (~11 %), methylglyoxal (~8 %), and acetic acid (~7 %). Glyoxylic acid is, nevertheless, an important intermediate species because it is directly converted to OXL in the aqueous phase upon oxidation. Other important findings concerning the chemical budgets are summarized below:
- 1. Glyoxal: About 70 Tg yr<sup>-1</sup> GLY is produced in the gas-phase in ERA-Interim, similar to Lin et al. (2014), while in 705 EC-Earth is calculated 3 % lower. The global gas-phase production of the present work is higher than other global model estimates, e.g., ~56 Tg yr<sup>-1</sup> (Myriokefalitakis et al., 2008), ~40 Tg yr<sup>-1</sup> (Fu et al., 2009, 2008), and ~21 Tg yr<sup>-1</sup> (Liu et al., 2012). This difference can be explained by the more comprehensive isoprene chemistry of the gas-phase scheme used here (Myriokefalitakis et al., 2020b). Indeed, isoprene secondary oxidation products (e.g., epoxides) are significant precursors of GLY in the atmosphere (Knote et al., 2014) and the contribution of isoprene epoxides 710 (IEPOX) from the gas-phase isoprene oxidation is here considered as a pathway of GLY formation. Note that the oxidation of other biogenic hydrocarbons, like terpenes and other reactive organics, may also result in GLY formation, since their chemistry is lumped on the first generation peroxy radicals of isoprene in the model (Myriokefalitakis et al., 2020b). Besides the biogenic hydrocarbon oxidation, the model considers GLY formation due to the oxidation of other organic species (e.g., Warneck, 2003), such as acetylene (4.8 Tg yr<sup>-1</sup>) and aromatics (18.8 Tg yr<sup>-1</sup>). In the gas 715 phase, other hydrocarbons, like ethene, further contribute to the atmospheric production of GLY via their oxidation products, mainly glycolaldehyde (5.4 Tg yr<sup>1</sup>). However, as in many modeling studies, additional primary and/or secondary glyoxal sources might be still missing in our model. Indeed, the elevated glyoxal concentrations over oceans that have been observed from space (e.g., Wittrock et al., 2006) would require at least 20 Tg yr<sup>-1</sup> of extra marine sources to reconcile model simulations with satellite retrievals (Myriokefalitakis et al., 2008). Great 720 uncertainties, however, still exist on these oceanic sources (Alvarado et al., 2020; Sinreich et al., 2010) and therefore the only glyoxal primary sources accounted for in the model are from biofuel combustion and biomass burning

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processes (e.g., Christian et al., 2003; Fu et al., 2008; Hays et al., 2002), resulting overall in <u>about</u> 7 g yr<sup>-1</sup> on average in the model (Myriokefalitakis et al., 2020b). Glyoxal is rapidly destroyed in the atmosphere via photolysis (~70 %), followed by its oxidation in the gas phase (~15 %) and the aqueous phase (~15 %). Roughly 5.4 Tg yr<sup>-1</sup> of glyoxal <u>is</u> produced in the aqueous phase via the dissolved GLYAL oxidation in ERA-Interim, close to the Liu et al. (2012) calculations, but somehow higher compared to EC-Earth. Overall, the net cloud uptake of glyoxal in ERA-Interim is <u>6.3 Tg yr<sup>-1</sup></u>, which is higher than the estimates from Liu et al. (2012) (<u>1.6 Tg yr<sup>-1</sup></u>). As expected, this increase is due to the applied glyoxal oxidation scheme in the aqueous phase of our base simulations. Finally, <u>4.2 Tg yr<sup>-1</sup> of glyoxal</u> are removed from the atmosphere via wet scavenging (~73 %) and dry deposition (~27 %).

- 2. Glycolaldehyde: GLYAL is also a significant species for OXL atmospheric abundance since its oxidation directly produces GLY both in the gas and the aqueous phase. In ERA-Interim, the gas-phase production is 92.5 Tg yr<sup>-1</sup> on global scale, with the primary sources accounting for 5.4 Tg yr<sup>-1</sup> (Myriokefalitakis et al., 2020b) on average. In EC-Earth, the gas-phase production is ~1 % lower. GLYAL is destroyed via gas-phase photolysis (~55 %) and by OH radicals in the gas phase (~35 %) and the aqueous phase (~10 %). Ethene oxidation products contribute ~39 % to GLYAL production, but isoprene chemistry dominates its chemical production in the model. The only source of GLYAL in the aqueous phase is nevertheless the transfer from the gas phase. The dissolved GLYAL is oxidized to produce GLY (~60 %) and GLX (~40 %), overall resulting in a net aqueous uptake of §.3 Tg yr<sup>-1</sup> in ERA-Interim, close to the estimates of Liu et al. (2012), but almost 40 % higher than in Lin et al. (2014). This higher uptake of GLYAL in the aqueous uptake of GLYAL is calculated ~24 % lower compared to ERA-Interim.
- 3. Methylglyoxal: The global annual mean gas-phase production of MGLY in ERA-Interim is 237 Tg yr<sup>1</sup> on average, with the primary sources accounting for 4.6 Tg yr<sup>1</sup>. The gas-phase production is higher than the 160 169 Tg yr<sup>1</sup> reported by other modeling studies (Fu et al., 2008; Lin et al., 2014; Liu et al., 2012) owing to the contribution of oxidation products considered in the gas-phase isoprene chemistry scheme (Myriokefalitakis et al., 2020b). Roughly 56 % of MGLY is produced via the gas-phase oxidation of HYAC with OH radicals, which is lower than the estimated ~75 % in Fu et al. (2008). The remaining MGLY production is due to isoprene oxidation products, i.e., ~10 % from IEPOX oxidation, and ~7 % from methyl vinyl ketone (MVK) and methacrolein (MACR) oxidation. In the aqueous phase, MGLY is produced via the dissolved HYAC oxidation (13.0 Tg yr<sup>1</sup>) and then further oxidized by OH radicals (11.6 Tg yr<sup>1</sup>) into pyruvic acid (PRV), methylglyoxal oligomers (MGLYOLI), and to a lesser extent into GLX. Note that the calculated contribution of dissolved HYAC to the aqueous-phase production of MGLY is higher compared to the nearly negligible rates in Liu et al. (2012) because of the higher gas-phase production of HYAC in our model. MGLY is chemically destroyed in the model mainly by gas-phase photolysis (~60 %), the OH radicals in the gas phase (~55 %).

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4. Pyruvic and acetic acids: The chemical production of PRV is 14.7 Tg yr<sup>-1</sup> in ERA-Interim and 16.7 Tg yr<sup>-1</sup> in EC-Earth. PRV is mainly produced by terpene oxidation via O<sub>3</sub> (~51 %) in the gas phase followed by methyl vinyl ketone (MVK) oxidation ( $\sim$ 5 %). In the aqueous phase, PRV is solely produced from MGLY oxidation (6.5 Tg yr<sup>1</sup>) and subsequently oxidized to CH3COOH. PRV is mainly removed via photolysis in the gas phase and via oxidation by OH radicals in the aqueous phase (~30 %). However, more than half of the produced PRV in the aqueous phase directly contributes to the SOA mass of the model upon cloud evaporation. The gas-phase production of acetic acid is 44.3 Tg yr<sup>-1</sup>, with the primary sources accounting for approximately 2<u>3.9</u> Tg yr<sup>-1</sup>. In the aqueous phase, roughly 3 Tg yr<sup>1</sup> of CH<sub>3</sub>COOH is produced via PRV oxidation. Note that the net uptake of CH<sub>3</sub>COOH (~0.7 Tg yr<sup>1</sup>) is calculated in the model similar to the Lin et al. (2014) estimates, but smaller than the 6.7 Tg yr<sup>-1</sup> calculated by Liu et al. (2012).

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5. Glyoxylic acid: The GLX production rate is 7.1 Tg yr<sup>-1</sup> in EC-Earth and it is ~30 % lower in ERA-Interim. About 55 % of the produced GLX is directly oxidized to oxalic acid in the aqueous phase and ~25 % is added directly to the SOA pool. Upon cloud evaporation, part of the produced GLX is also transferred in the gas phase, where it is either 790 oxidized by OH radicals (~60 %), photolyzed (~33 %), or deposited (~7 %). Due to the destruction of GLX in the gas phase, its total production is lower (~60 %) compared to the production estimates in Lin et al. (2014) and Liu et al. (2012). For the EC-Earth and ERA-Interim, most of the produced GLX in the aqueous phase is derived from the oxidation of GLYAL (~48 %), followed by the oxidation of CH<sub>3</sub>COOH (30 %), GLY and its oligomeric products (~ 15 %), and MGLY. The relative contributions in our calculations differ from the estimates in Lin et al. (2014), where GLX is primarily produced by GLY oxidation (~77 %) followed by GLYAL (~14 %), MGLY (~1 %), and acetic acid (~8 %). These differences are also caused by the direct contribution of the GLY oxidation products to the OXL formation. On the other hand, in the ERA-Interim(sens) simulation the calculated fractions agree well with other published estimates, where GLY overall dominates (~60 %) the GLX production in the aqueous phase.

Figure 2a presents the annual mean (average 2000-2014) net chemistry production rates of OXL in EC-Earth, and the respective 800 absolute differences compared to ERA-interim (Fig. 2c). The maximum OXL production rates are calculated around the tropics and in the southern hemisphere, where both biogenic emissions (mainly isoprene) and the liquid cloud water are substantially enhanced (Fig. S2a). The Amazon region appears as the largest source of OXL, along with central Africa and Southeast Asia. At higher latitudes (> 45° N) the lower cloud liquid water content and vegetation cover lead to a lower OXL production over Asia and North America. However, over highly populated regions in the northern hemisphere, such as in Europe, the US, and 805 China, enhanced OXL production rates are calculated due to its anthropogenic precursors. Furthermore, a significant source of OXL is calculated downwind of land areas, such as the South Pacific, and the tropical Atlantic Ocean, due to the long-range transport of OXL precursors.

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The illustrated differences in OXL production between ERA-interim and EC-Earth (Fig. 2c) are caused due to the adopted atmospheric dynamics (i.e., online calculated versus offline), as both simulations use identical prescribed anthropogenic and biogenic emissions (see van Noije et al., 2021). For instance, EC-Earth calculates higher cloud water concentrations at ~800-600 hPa around the tropics (30° S - 30° N) compared to ERA-Interim. In contrast, lower concentrations are derived aloft, with ERA-Interim presenting enhanced cloud water concentrations at ~400 hPa (Fig. S2b). Moreover, due to the lower OH concentrations in the tropical and subtropical troposphere (Fig. S1b), EC-Earth gives lower OXL production rates, especially over intense biogenic emission areas. Overall, the difference in the oxidizing capacity of the atmosphere between the two
configurations significantly impacts the aqueous phase OXL production efficiency in the model.

#### 3.1.2 Sulfate

Sulfate (SO<sub>4</sub><sup>2-</sup>) is the main inorganic aerosol species produced in the aqueous phase, and similar to OXL its production in the model mainly occurs in cloud droplets. In addition, these two species largely reside in the aerosol accumulation mode of the model (roughly 99 % for SO<sub>4</sub><sup>-2</sup> and 97 % for OXL). SO<sub>4</sub><sup>2-</sup> is a key species for determining atmospheric acidity, and therefore
we also present here the sulfate budget in conjunction with that of OXL. Sulfate is produced both in cloud droplets and in aerosol water, with the production in aerosol water having however a negligible contribution on a global scale. In contrast to OXL, for which no gas-phase production is considered, the gas-phase oxidation of SO<sub>2</sub> via OH radicals contributes to the total SO<sub>4</sub><sup>2-</sup> concentrations with about 12.0 Tg S yr<sup>-1</sup> (Table 2b). Our global estimate of the gaseous sulfuric acid (H<sub>2</sub>SO<sub>4</sub>) production is higher than in EC-Earth v2.4 (~7.8 Tg S yr<sup>-1</sup> averaged for the years 2000–2009; van Noije et al., 2014), but slightly lower

than in the EC-Earth3-AerChem AMIP simulations (van Noije et al., 2021) used for the CMIP6 experiments (available in <u>https://esg-dn1.nsc.liu.se/search/cmip6-liu/;</u> last access 11/06/2021) where ~12.9 Tg S yr<sup>-1</sup> of H<sub>2</sub>SO<sub>4</sub> are produced (averaged for the years 2000-2014). These differences can be directly attributed to the OH radical production rates in the gas phase between the new and the previous <u>chemistry</u> versions of the atmospheric model, as have been discussed in Myriokefalitakis et al. (2020b). Despite the generally lower gas-phase OH radical levels (Fig. S1b), the slightly higher (~8 %) global H<sub>2</sub>SO<sub>4</sub> gas-

835 phase production rate in EC-Earth than in ERA-Interim (Table 2b), can be attributed to the higher (~6%) DMS emissions in EC-Earth (Fig. S4b), that contribute to the atmospheric SO<sub>2</sub> levels over the ocean.

The aqueous-phase SO<sub>4</sub><sup>2-</sup>chemistry production from the oxidation of dissolved SO<sub>2</sub> is 39.8 Tg S yr<sup>-1</sup> in EC-Earth (Table 2b), which is higher than in EC-Earth v2.4 (29.3 Tg S yr<sup>-1</sup>; van Noije et al., 2014) and EC-Earth3-AerChem (32.5 Tg S yr<sup>-1</sup>). The higher SO<sub>4</sub><sup>2-</sup> chemical production is mainly due to the higher SO<sub>2</sub> aqueous-phase oxidation rates by H<sub>2</sub>O<sub>2</sub>. In more detail, our calculations show that ~84 % of the global SO<sub>4</sub><sup>2-</sup> production in EC-Earth is due to the dissolved SO<sub>2</sub> oxidation via H<sub>2</sub>O<sub>2</sub>; ~33.3 Tg S yr<sup>-1</sup> is produced due to H<sub>2</sub>O<sub>2</sub>, higher compared to 23.9 Tg S yr<sup>-1</sup> in van Noije et al. (2014). The dissolved SO<sub>2</sub> oxidation via O<sub>3</sub> (6.4 Tg S yr<sup>-1</sup>) is also higher than in EC-Earth v2.4 (5.4 Tg S yr<sup>-1</sup>). The contribution of CH<sub>3</sub>O<sub>2</sub>H to the SO<sub>4</sub><sup>2-</sup> aqueousphase production is however small (<u>0.05 Tg S yr<sup>-1</sup></u>) in the model, with the HO<sub>2</sub> contribution being practically negligible on the global scale (<u>0.02 Tg S yr<sup>-1</sup></u>) for all simulations performed in this study. A total annual mean deposition rate of <u>52.5</u>, Tg S yr<sup>-1</sup>

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Deleted: ~ Deleted: ~ Deleted: ~ Deleted: ~ is simulated in EC-Earth, with wet scavenging dominating the total deposition rate (~93\_%). Note also that 2.5 % of the sulfur in the SO<sub>2</sub> emissions (~1.6 Tg S yr<sup>-1</sup>) is assumed to be in the form of SO<sub>4</sub><sup>2-</sup> for all simulations, which accounts for its formation in the sub-grid plumes (Aan de Brugh et al., 2011; Huijnen et al., 2010). Overall, a global SO<sub>4</sub><sup>2-</sup> lifetime over deposition of
 ~4.8 days is calculated in EC-Earth, which is lower than in ERA-Interim (~6.6 days), but similar to the EC-Earth v2.4 estimate

(~4.9 days).

Figure 2b also shows the annual mean  $SO_4^{2-}$  net chemistry production rates in EC-Earth. High  $SO_4^{2-}$  production rates are calculated downwind of major anthropogenic  $SO_2$  emission hotspots, such as Central Europe, Eastern US, India, Russia, and Eastern Asia. Furthermore, relatively high production rates due to biomass burning and volcanic eruptions are calculated in

South America, Southern Africa, and Indonesia. Significant SO4<sup>2-</sup> production is calculated over almost all oceanic regions due to the SO<sub>2</sub> production via the gas-phase oxidation of marine DMS emissions (Fig. S4a). Compared to the ERA-Interim simulation, however, the SO4<sup>2-</sup> production rates in EC-Earth are on average slightly higher over land in the tropics and extratropics. This increase can be attributed to combined effects that result in differences in chemical production and deposition rates (Table 2b). Some differences over oceans are nevertheless expected due to the differences in DMS concentrations, since
 BMS emissions are online calculated in the model based on sea surface temperature and wind velocity (Fig. S4b).

In EC-Earth, the total Fe (TFe) soil emissions result in  $59.33 \pm 1.22$  Tg yr<sup>1</sup>, while in ERA-Interim they are  $48.96 \pm 0.95$  Tg

# 3.1.3 Iron

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yr<sup>1</sup>. This difference results from the differences in wind speed between EC-Earth and ERA-Interim. EC-Earth produces higher dust emissions over large parts of the Middle East and Asia compared to ERA-Interim (Fig. S4f), which explains the differences in TFe emissions (TFeC emissions do not differ). However, most of the dissolved Fe from mineral dust in the model originates from atmospheric dissolution processes. In EC-Earth, FeD is primarily dissolved due to aerosol acidity at 0.31 Tg yr<sup>-1</sup>, followed by the ligand-promoted dissolution that additionally produces 0.17 Tg yr<sup>-1</sup> (Table 2d). Fe primarily resides (~98.4 %) in the slow pool of Fe-containing dust aerosols in the model, in particular in the coarse mode, with about 1.0 % being emitted as nano-sized iron oxides (intermediate pool), and ~0.5 % as ferrihydrite (fast pool). Thus, most of the dissolved Fe release originates from the heterogeneous inclusion of nano-Fe grains in the internal mixture of various Fe-containing minerals such as aluminosilicates, hematite, and goethite (~66 %), followed by nano-sized iron oxides (~24 %) and to a lesser extent by ferrihydrite (~10 %). Note, however, that the Fe release from aluminosilicates, hematite, and goethite particles is a slower process compared to the other soil classes considered in the model, as dictated by the three-stage approach applied for this

880 study (Table S4).

Fe emissions from combustion processes are estimated at  $2\frac{518}{518} \pm 0.105$  Tg yr<sup>-1</sup> in both simulations, with 0.012 Tg yr<sup>-1</sup> being weitted as dissolved from the primary oil combustion processes. Roughly 0.274  $\pm$  0.010 Tg yr<sup>-1</sup> are released through Fe

dissolution from combustion aerosols in EC-Earth, in good agreement with ERA-Interim ( $0.285 \pm 0.011$  Tg yr<sup>-1</sup>). The acid-900 promoted dissolution contributes ~17 % and photo-reductive processes ~16 % to the Fe release from combustion particles, thus most Fe release comes from the ligand-promoted dissolution. This result is in line with laboratory studies (e.g., Chen and Grassian, 2013), where the contribution of oxalate-promoted dissolution is several times larger than the proton-promoted pathway under highly acidic dark conditions. According to our calculations, the relative contribution of atmospheric processing to the combustion aerosol Fe solubilization (~11 %) is significantly higher compared to that of crystalline dust minerals (~1 %), in agreement with laboratory (e.g., Chen et al., 2012; Fu et al., 2012) and modeling (e.g., Ito, 2015; Ito and Shi, 2016)

studies.

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The annual mean dissolution rates of FeC and FeD in EC-Earth are presented in Fig. 3. For combustion aerosols, the maximum dissolution rates occur downwind of biomass burning sources and highly populated regions, such as South America and Central Africa, the Middle East, India, and China. High dissolution rates are more likely to coincide with high OXL concentrations

- 910 (Ito, 2015). Indeed, the model calculates important dissolution rates near regions where the OXL production rates are enhanced (Fig. 2a), such as over the Amazon basin and Central Africa, as well as downwind of these regions as the combustion aerosols are transported to the open ocean, in agreement with observations (e.g., Sholkovitz et al., 2012). For the mineral dust aerosols, most of the FeD dissolution fluxes occur downwind of the major dust source regions (e.g., the Sahara and the Gobi Desert), where the atmospheric transport of anthropogenic pollutants, such as SOx and NOx, enhances atmospheric acidity; e.g., the
- 915 Fe release from the dust minerals due to proton-promoted dissolution processes is enhanced over the Middle East. Significant dissolution rates are also simulated over the Atlantic Ocean at the outflow of the Sahara, as well as at the outflow of Asian desert regions to the Pacific Ocean. High rates due to the contribution of the organic ligand-promoted dissolution processes are calculated downwind of central Africa and the equatorial Atlantic Ocean, where the oxidation of biogenic hydrocarbons in the presence of cloudiness leads to enhanced OXL aqueous-phase formation rates. On the contrary, the efficiency of ligand-920 promoted dissolution is substantially suppressed near dust source regions due to the low OXL availability (Fig. 2a).

Figure 3 also presents the absolute differences between the ERA-Interim and EC-Earth annual mean Fe dissolution rates. ERA-Interim has significantly lower dissolution rates in the tropics (e.g., Central Africa) and around the equator, both for FeC (Fig. 3c) and FeD (Fig. 3d). This decrease is attributed both to the differences in atmospheric dynamics between the two model configurations, as well as to the suppression of the organic-dissolution processes with lower OXL production. Indeed, Fig. 2c

925 shows that in ERA-Interim, OXL production rates increase in the tropics, impacting Fe-dissolution rates. In contrast, FeC dissolution rates increase in ERA-Interim over the Arabian Peninsula, India, and Eastern Asia, due to fluctuations in OXL production and aerosol acidity. EC-Earth also shows lower FeD dissolution rates over the Northern Pacific in the outflow of Asia. These differences are due to a higher aerosol acidity (i.e., up to ~1 pH unit; Fig. S3d,f) in ERA-Interim due to changes in the buffering capacity of dust promoted by the higher calcite emissions in EC-Earth (Fig. S4f). This is the case especially

for coarse dust aerosols where the majority of the Fe resides. EC-Earth also shows differences (positive or negative) with ERA-930

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Interim in the Fe dissolution rates over oceanic regions (Fig. 3e,d), likely due to differences in SO4<sup>2-</sup> production over oceans from marine DMS emissions (Fig. S4b), and to the impact of sea salt emissions (Fig. S4d) upon the buffering capacity of the

940 solution. All in all, the total DFe atmospheric source in EC-Earth, accounting for both primary emissions and atmospheric processing, is 0.806 ± 0.014 Tg yr<sup>1</sup> for the present day, well in the range of estimates presented in the model intercomparison study  $(0.7 \pm 0.3 \text{ Tg yr}^{-1})$  in Myriokefalitakis et al. (2018).

#### 3.2 Evaluation of new model features against observations

All developments described in this work have been implemented over the EC-Earth3-AerChem model version, which has 945 proved to simulate the atmospheric aerosol cycles comparably well to other global models and to reproduce satisfactory the optical properties (Gliß et al., 2021). Thus, we do not expect substantial changes in EC-Earth3-Iron ability to represent the aerosol cycle, nor their optical properties, compared to EC-Earth3-AerChem. However, owing to the significant differences in the gas-phase and aqueous chemistry between versions, we provide an overall assessment of the Aerosol Optical Depth (AOD). In addition, as one of the novelties of this work is to consider explicitly how dust composition affects the atmospheric iron burden and alters acidity (e.g., through calcite), and therefore a comparison of dust fields with *in-situ* observations is also

provided. Finally, simulations of specific species key to our developments, such as oxalate, sulfate, and total and soluble iron

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are also evaluated.

#### 3.2.1 AOD, dust concentration, and deposition

- The annual mean AOD at 550 nm modeled in EC-Earth for 2000-2014 compares favorably with AERONETv3 direct-sun level 955 2.0 data (Fig. 4a). Overall, the model presents an nMB of -9.% and an nRMSE of 46.%, considering information from 738 AERONET sites. The regional analysis suggests a slightly better behavior in northern hemisphere regions (e.g., North America, Europe, East Asia) dominated by anthropogenic aerosols (normalized errors and biases, below 45 % and  $\pm 10$  %, respectively). The largest deviations from observed AOD occur in the Southern Hemisphere (e.g., South Africa, Australia, and Oceania) or remote regions. Over dust-dominated regions (e.g., North Africa, West Asia, and the Middle East) the model also behaves well 960 (with normalized errors and biases below 45 % and  $\pm 10$  %, respectively). Selecting specifically dust-dominated sites for the comparison (Fig. 4b) and following the criteria explained in Sect. 2.5, EC-Earth slightly overestimates the retrieved AOD at
- 550 nm over North Africa (nMB=21,%) and shows underestimations over sources in West Asia and the Middle East, as well as in transport regions such as Central America. In general, EC-Earth's ability to reproduce the annual mean AOD at 550 nm holds for dusty sites, with a normalized mean bias of 3 %, and a normalized root mean square error of 37, % over 38 sites. 965 Overall, the average optical depth for dust at 550 nm (annual mean over the 2000-2014) yields a value of  $0.032 \pm 0.005$ , which

falls well in the range of observationally based estimates, based on *in-situ* measurements, satellites and global models (i.e.,  $0.030 \pm 0.005$ ; Ridley et al., 2016).

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The comparison of model outputs with climatologies of dust surface concentration from the RSMA and the AMMA campaign 975 (Fig. 4c) yields slightly poorer results, with an nMB of 19, % and an nRMSE of 81, %, as an average of the 23 sites available. EC-Earth best reproduces dust surface concentrations over source regions, such as North Africa (nMB=21.9 %, nRMSE=37.7 %), shows underestimations in transport areas (e.g., Central America: nMB=-37,%, nRMSE=39,%) and poorly represents the surface concentration in remote regions (e.g., South Pacific and Southern Ocean, with nMB up to -98, % and nRMSE up to 113.%). This issue is also reflected in the evaluation of the dust deposition field (Fig. 4d), with positive and negative biases 980 over source regions but a clear underestimation of the deposited mass on transport and remote regions (e.g., the Southern Ocean). These discrepancies point towards an overestimation of dust deposition. For instance, EC-Earth3-Iron may share the difficulties of many global models in representing the long-range transport of dust, in particular coarse particles downwind of dust source regions (e.g., Adebiyi and Kok, 2020). As minerals in dust constitute the primary source of TFe to the atmosphere, the overestimation close to the source regions and the underestimation in remote regions will also influence the representation 985 of the DFe in the model.

# 3.2.2 Oxalate

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The averaged OXL surface concentrations in EC-Earth for the boreal winter (December, January, and February; DJF) and summer (June, July, and August; JJA) are presented in Fig. 5. OXL surface concentrations are distributed roughly between 60° S and 60° N, mainly in regions where intensive VOC emissions from anthropogenic and biogenic sources coexist with cloud water. The highest OXL concentrations are calculated over tropical Africa, the Amazon Basin, eastern Asia, the eastern United

- 990 States, and Europe, clearly showing the strong impact of OXL precursors (e.g., glyoxal) and the availability of cloud water. In the northern hemisphere, OXL concentrations are generally calculated higher in summer and lower in winter, indicating a strong impact of temperature and photochemistry on the production rate of oxalic acid in the aqueous phase. During DJF, the model calculates lower OXL concentrations over mid- and high-latitude regions, such as East Asia, Central Europe, and
- 995 Northern US. Over these highly populated regions, the aerosol water content is enhanced, following the increased SO4<sup>-2</sup> production due to anthropogenic activities, and the aqueous-phase OXL production in deliquesce particles also contributes to OXL atmospheric concentrations. Furthermore, high OXL concentrations are calculated in the tropics for both seasons, due to the photochemical activity and the intense sources of biogenic VOCs in these regions.
- Figure 5 further presents the differences of OXL concentrations between EC-Earth and ERA-Interim. In general, OH levels in 1000 ERA-Interim are higher, which causes a more efficient oxidation of OXL precursors for both seasons. Moreover, ERA-Interim shows higher concentrations around the intertropical convergence zone (ITCZ) due to differences in meteorology between the two simulations, as discussed above. During boreal winter, some differences are observed over the subtropics of the Northern Hemisphere. Although OXL concentrations are very low over these latitudes, the relatively strong increase in liquid water that serves as a medium for OXL production, both for clouds (Fig. S2b) in higher altitudes (~400 hPa), and at the surface for

1005 deliquescent particles (Fig. S2d,e) in the ERA-Interim simulation. In the vertical, OXL concentrations are distributed in the

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model from the surface to ~400 hPa with a maximum at around 900 hPa. The zonal mean differences, however, indicate strong increases in the southern hemisphere (30° S - 0°) during boreal winter (Fig. 5e). Compared to EC-Earth, ERA-Interim calculates higher OXL concentrations in the upper troposphere for both seasons (Fig. 5e,f), mostly due to more efficient transport of OXL answers by deer converting into the transport of a structure resonance.

1020 OXL precursors by deep convection into the tropical and extratropical upper troposphere. In the lower and middle troposphere, higher concentrations are calculated in ERA-Interim depending on the location and the season. The concentrations in EC-Earth are also lower than in ERA-Interim in the NH extratropics during boreal summer due to a lower chemical production (Fig. 5f).

Figure 6 presents the comparison of the different model simulations performed for this work with OXL surface observations. OXL concentrations show a strong seasonal dependence, with maxima during the warm season due to the intense

- 1025 photochemical activity combined with the higher precursor's abundance. Over the Mediterranean, and specifically the eastern part which is characterized by the long-range transport of air pollution and from surrounding urban centers (Kanakidou et al., 2011), the model underestimates the observed concentrations during winter at the Finokalia station in all simulations (Fig. 6a), either due to missing OXL primary and secondary sources, or a too strong removal. During summer, ERA-Interim satisfactorily simulates the observed OXL levels, also representing the observed trend, which indicates that the model reproduces the mixing
- 1030 and aging of the air masses in the region under favorable meteorological conditions and intense solar radiation. EC-Earth calculates lower OXL concentrations than ERA-Interim, due to the lower oxidizing capacity, underestimating thus the observed concentrations for all seasons. On the other hand, ERA-Interim(sens) tends to underestimate the observations for all seasons more than the other simulations, further indicating the important role of the secondary sources to OXL atmospheric concentrations in the region. At the Puy de Dome site (Fig. 6c), which is located at 1450m asl, ERA-Interim underestimates
- 1035 the observed OXL concentrations, although simulates them more realistically compared to EC-Earth, especially during summer (Fig. 2c). The seasonal variation in the area can be explained by the stronger upward transport of air masses during summer (Legrand et al., 2007), increasing thus the OXL production in the region. However, the model fails to represent the observed OXL levels, possibly due to missing sources. The importance of other production pathways, not related to the aqueous-phase GLX oxidation, is demonstrated in the comparison of the observed OXL levels with the ERA-Interim(sens) simulation. Again,
- 040 ERA-Interim(sens) deviates stronger from the measured values than other simulations. Nevertheless, this indicates that other species may further contribute to OXL production, such as the decay of longer diacids (e.g., azelaic, and malonic acids) (Legrand et al., 2007) that are currently not included in the model. Another reason may be the impact of the enhanced cloud LWC in the region, implying a more intense cloud processing compared to other surface sites, and thus, <u>a</u> faster oxidation of oxalic acid into CO<sub>2</sub> (Ervens et al., 2004) in the model. Finally, at the Azores (Fig. 6e), a site that is characterized by a marine
- environment, the model tends to underestimate the observed OXL concentrations most of the time, with ERA-Interim presenting again a better skill than other simulations. EC-Earth underestimates the observed concentrations more than ERA-Interim, especially during summertime, and ERA-Interim(sens) simulates the lowest OXL concentrations. The observed OXL levels in the region, however, can be explained either by the transport of pollutants from the continents or the photochemical production in the region. Thus, the illustrated differences against the observations between EC-Earth and ERA-Interim can be

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attributed to differences in the oxidizing capacity and in simulated transport of EC-Earth, such as the vertical mixing in the 1055 troposphere (e.g., van Noije et al., 2014), that has a further impact on OXL precursors like glyoxal. Furthermore, since the long-range transport is found relatively constant in summer and winter in the region (Legrand et al., 2007), other species of marine origin, such as the unsaturated fatty acids (e.g., linoleic and oleic acids) may also contribute as precursors to the OXL production, especially during summer, but are not included in the model too. All in all, we acknowledge that other formation pathways of OXL, primary or secondary, may exist in the atmosphere (e.g., Baboukas et al., 2000), for example, higher DCAs (such as malonic, succinic, glutaric, and adipic acids) may act as precursors for smaller dicarboxylic acids like OXL, both in the gas- (e.g., Kawamura and Ikushima, 1993) and the aqueous phase (e.g., Ervens et al., 2004; Lim et al., 2005; Sorooshian et al., 2006) and could further contribute to atmospheric OXL concentrations.

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In Fig. 6g, OXL observations reported in the literature are compared with monthly mean simulations. Due to the relatively low resolution of the global model (i.e., 3° x 2° in longitude by latitude), the spatial variability of urban emissions cannot be well 065 resolved. Therefore, urban stations are omitted, and the comparison is limited to locations representative of background concentrations, All simulations tend to underestimate OXL observations, with lower biases in ERA-Interim (i.e., nMB = 46 %, nRSME = 117 %). As expected, for almost all sites the ERA-Interim simulation calculates the highest OXL concentrations and the ERA-Interim(sens) the lowest (i.e., nMB = 74 %, nRSME = 125 %). The latter indicates that additional production pathways need to be considered in modeling studies to capture the observed OXL concentrations. EC-Earth underestimates 070 the observed concentrations more than ERA-Interim (i.e., nMB = 40%, nRSME = 10%), but less than ERA-Interim(sens),

highlighting also the importance of the atmospheric oxidating capacity and the atmospheric dynamics in the OXL production. Overall, our analysis indicates that the model either misses OXL sources (primary and secondary) or overestimates OXL sinks, especially during winter. Thus, under relatively lower temperatures and irradiation, the model may neither represent well the fast secondary OXL production in wood-burning plumes nor the secondary production through species produced by the 1075 oxidation of emitted from vehicles and other anthropogenic activities, such as ethane and aromatic hydrocarbons.

#### 3.2.3 Sulfate

The averaged SO42- surface concentrations as calculated in EC-Earth for the boreal winter and summer are presented in Fig. 7. During both seasons, high SO42-concentrations are simulated near or downwind major anthropogenic emission hotspots, where the vast majority of the surface SO<sub>2</sub> emissions from anthropogenic origin occur (e.g., Tsai et al., 2010). Enhanced SO<sub>4</sub><sup>2-</sup> surface 1080 concentrations are also calculated downwind of biomass burning and volcanic eruptions, showing overall the impact of SO2 primary sources and the abundance of cloud water over these latitudes. Over the remote oceans, however, DMS oxidation may significantly contribute to the SO4<sup>2-</sup> surface concentrations, as also the sulfur emissions over major shipping routes. In Fig. 7, the differences between EC-Earth and EC-Earth-AerChem in the averaged SO42-surface and zonal mean concentrations are also presented, both for boreal winter (Fig. 7c,e) and summer (Fig. 7d,f). Considering, however, that the EC-Earth version

1085 developed for this work is based on the EC-Earth-AerChem model version, the illustrated differences are solely due to the

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- than in EC-Earth (see Sect. 3.1.2), leading overall to a lower conversion of SO<sub>2</sub> to sulfate in the aqueous phase. Moreover, in our simulations SO<sub>4</sub><sup>2-</sup> is also produced in deliquesced particles, which partly contribute to the SO<sub>4</sub><sup>2-</sup> atmospheric concentrations, especially over highly populated regions during the boreal winter when sulfur emissions due to anthropogenic activities are enhanced. On the contrary, during boreal summer some differences are illustrated (Fig. 7d) mostly over the Middle East where
- 1105 the EC-Earth-AerChem simulation results in lower SO4<sup>2-</sup> concentrations, and over Eastern Asia, where some higher concentrations in EC-Earth-AerChem occur. These negative/positive differences are attributed to differences in oxidizing capacity between the two models, both in the gas and the aqueous phase. Finally, the zonal mean differences indicate higher concentrations in EC-Earth in the lower troposphere of the northern hemisphere during boreal winter (Fig.7e), as for the SO4<sup>2-</sup> surface concentrations. Again, for boreal summer, no important differences are presented (Fig.7f).
- Figure 6 further presents the model comparison with SO4<sup>2-</sup> surface observations. Sulfate concentrations maximize<u>under</u> intense photochemical activity and high SO<sub>2</sub> atmospheric levels; generally suggesting <u>also</u> a faster formation rate compared to oxalic acid (Ervens et al., 2004; Legrand et al., 2007). At the Finokalia station in the Mediterranean (Fig. 6b), the model overestimates the observed SO4<sup>2-</sup> concentrations during boreal winter and summer in ERA-Interim, probably due to too high SO<sub>2</sub> background concentrations <u>because of a too strong</u> long-range transport from surrounding regions. During winter, EC-Earth better
- 1115 reproduces the observations, probably due to the lower oxidizing capacity compared to ERA-Interim, but in late spring and early summer, the active photochemistry in the region leads again to an overestimation of the observed concentrations. The EC-Earth-AerChem simulation leads to generally lower concentrations compared to our EC-Earth simulation, tending to somehow underestimate the observed concentrations, except for autumn. At Puy de Dome (Fig. 6d), ERA-Interim, which has a higher cloud liquid water content aloft and a more intense oxidizing capacity compared to EC-Earth, overestimates the
- 120 observed SO4<sup>2-</sup> concentrations in almost in all seasons. In contrast, EC-Earth better simulates the measured\_SO4<sup>2-</sup> concentrations, although it seems to underestimate the observations during summer. At that site, the EC-Earth-AerChem calculations agree well with EC-Earth, although concentrations are again slightly lower. At the Azores (Fig. 6f), ERA-Interim also simulates the observed concentrations well, following the observed annual cycle. In contrast, both EC-Earth ad EC-Earth-AerChem underestimate the SO4<sup>2-</sup> observations, especially during spring and summer. Note that the SO4<sup>2-</sup> production in
- 1125 this marine site is attributed to the SO<sub>2</sub> atmospheric levels both from air masses advected from industrialized regions and the local production due to the oxidation of marine DMS emissions. Thus, the differences between ERA-Interim and the other EC-Earth simulations presented in this study may indicate slower aging of the polluted air masses transported in the region. Finally, a comparison of the model's monthly mean predictions with a compilation of SO<sub>4</sub><sup>2-</sup> observations (n=3828) around the globe (Daskalakis et al., 2016) is presented in Fig. 6h. EC-Earth tends to overestimate the available SO<sub>4</sub><sup>2-</sup> observations,
- 130 presenting positive biases (i.e., nMB = 16,%, nRSME = 5%), that are slightly lower than in ERA-Interim (i.e., nMB = 23%, nRSME = 57%). In contrast, EC-Earth-AerChem tends to slightly underestimate the observed concentrations (i.e., nMB = 5.2%)

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%, nRSME = 57 %), showing a slightly lower correlation coefficient (R = 0.70) than EC-Earth (i.e., R = ~0.76) and ERA-Interim (i.e., R = ~0.75).

# 3.2.4 Dissolved iron

Figures 8 and 9 present the averaged dissolved FeC (DFeC) and FeD (DFeD) surface concentrations, respectively, for DJF and JJA. EC-Earth calculates an annual global DFeC atmospheric burden of ~0.002 Tg, while ERA-Interim calculates slightly higher concentrations (~0.003 Tg) due to the more intense ligand promoted dissolution rates (Table 2c). Elevated DFeC
concentrations during boreal winter (Fig. 8a) are calculated over central Africa, eastern Asia, and India, where significant DFe concentrations (~0.01-0.1 µg m<sup>-3</sup>) are associated with biomass burning and anthropogenic combustion emissions. During boreal summer (Fig. 8b), the maximum DFeC concentrations are calculated in the Northern latitudes, in particular over the Mediterranean Basin, the Middle East, the Western US, and China. The increase in the surface dissolved Fe concentrations over these regions, ranging from ~0.01-0.1 µg m<sup>-3</sup>, clearly highlights the anthropogenic contribution due to the enhanced solubilization of Fe when the combustion aerosols are mixed with acidic and organic pollutants during atmospheric transport (Fig. 3b). Due to the intense biomass burning in the Southern Hemisphere (i.e., South America, Central Africa, and Indonesia), the enhanced OXL production rates over such regions (Fig. 2a) lead to higher dissolved Fe concentrations. Figures 8a and 8b demonstrate, overall, that the geographic pattern of the DFeC concentrations may change from boreal winter to summer,

1160 origin.

Mineral dust emissions mostly occur in the mid-latitudes of the Northern Hemisphere (Figs. 9a, b), relatively close to where the vast majority of the population exists and the anthropogenic emissions of acidic compounds dominate. For both seasons, high DFe concentrations from dust occur over the midlatitudes of the Northern Hemisphere, where the major dust sources are located. However, the equatorial maximum during boreal winter tends to shift to the North during boreal summer following

following the biomass burning activity and the atmospheric processing of Fe-containing combustion aerosols of anthropogenic

- 1165 the migration of the ITCZ (Fig. 9b). DFe from mineral dust aerosols maximize over the major dust source regions, with surface concentrations of roughly 0.1-1 µg<sub>4</sub>m<sup>-3</sup> (Fig. 9a, b), overall dominating the Fe burden. The outflow from these regions transports DFe over the global oceans, where secondary maxima of ~0.01-0.1 µg<sub>4</sub>m<sup>-3</sup> are calculated, mainly over the Northern Hemisphere in the tropical Atlantic Ocean. The dissolved Fe associated with Saharan dust is, nevertheless, attributed to the long-range transport and the atmospheric processing that converts the insoluble Fe minerals to soluble forms.
- 1170 The differences between the EC-Earth and ERA-Interim are illustrated in the averaged DFeC (Fig. 8c,d) and DFeD (Fig. 9c,d) surface concentrations. The differences in DFeC between the EC-Earth and ERA-Interim are well correlated with those of OXL concentrations (Fig. 5 c,d), indicating the strong impact of ligand-promoted dissolution on the DFeC atmospheric load (Table 2c). Note, however, that the most important relative differences between the two simulations are calculated over regions with low dissolved FeC concentrations, and consequently the total burden does not change significantly (Table 2c). The

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- 1180 differences in the DFe concentrations associated with mineral dust aerosols follow the general anomaly pattern of the two model configurations due to differences in transport that lead overall to higher (~15 % globally) dust emissions in EC-Earth (Fig. S4f). We note, however, that the annual mean dust emission in EC-Earth3-Iron amounts 1257 ± 26 Tg yr<sup>-1</sup> (2000-2014 averaged), which falls in the lower range of the AEROCOM phase III models (Gliß et al., 2021) and is also at the low end of the range estimated by Kok et al. (2021) from inverse modelling (~1200-2900 Tg yr<sup>-1</sup>) for dust aerosol with a geometric /
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185 diameter  $\leq 10 \mu m$ .

Figure 10 presents a comparison of the different model simulations with cruise observations of dissolved Fe concentrations. The spatial distributions of the DFe observations for the accumulation, the coarse, and the total suspended aerosols are shown in Fig. 10a-c, respectively. The median (mean  $\pm$  standard deviation) DFe concentration in the accumulation mode amounts to 0.96 (3.68  $\pm$  6.44) ng m<sup>-3</sup> in the observations, while in EC-Earth and ERA-Interim is 3.67 (5.79  $\pm$  5.17) ng m<sup>-3</sup> and 3.12 (5.25,

- ±4,54) ng m<sup>-3</sup>, respectively. The respective observed concentration of DFe in the coarse mode is 0.85 (5.03 ± 10.90) ng m<sup>-3</sup>, while in EC-Earth and ERA-Interim the calculated values are around 1.74 (6.18,±7.29) ng m<sup>-3</sup> and 1.04 (5.33,±7.30, ng m<sup>-3</sup>). Finally, the concentration of DFe in total suspended particles (tsp) is 1.93 (7.83 ± 19.03) ng m<sup>-3</sup> in the observations, as well as 4.37 (12.51 ± 19.09) ng m<sup>-3</sup> and 4.39 (11.63,±17.12) ng m<sup>-3</sup> in EC-Earth and ERA-Interim, respectively. The correlation coefficients between the median values of the DFe cruise observations and the model results for EC-Earth and ERA-Interim
  are calculated as 0.49 (nMB = 38 %, nRMSE = 164 %) and 0.58 (nMB = 23 %, nRMSE = 145 %) for the accumulation
- aerosols, while for the coarse mode are calculated, respectively,  $0.4G_{4}(nMB = -9.\%, nRMSE = ~193.\%)$  and 0.59.(nMB = -2.5.%),  $0.4G_{4}(nMB = -9.\%)$ ,  $0.4G_{4}(nMB = -9.\%)$ ,  $0.4G_{4}(nMB = -17.5\%)$  for EC-Earth and ERA-Interim simulations (Fig. S5).

The spatial distributions of the absolute differences between the simulated DFe concentrations in EC-Earth and the observations are also presented in Fig.10 d-e. The model shows a general overestimation of the observed DFe concentrations around the tropics (up to ~10 ng m<sup>-3</sup>), but an underestimation at the mid-to-higher latitudes (up to ~5 ng m<sup>-3</sup>). It is, however, unclear why the model is unable to capture the observed concentrations over such regions. Reasons could be the misrepresentation of dust coarse emissions, missing anthropogenic primary Fe emissions, weak secondary sources of the

- dissolved Fe (especially in the southern latitudes), or even systematic errors in the transport of coarse particles. Over the Pacific, EC-Earth better predicts the average DFe concentrations. For completeness, Fig. S5 also presents a comparison of the
  simulated and observed TFe concentrations, showing that the model better captures the TFe concentrations in the accumulation mode (Fig. S6g) than in coarse mode (Fig. S6h). Considering, however, that the TFe is mostly dominated by primary sources, the calculated differences to the observed concentrations (Fig. S6e-f) downwind continental sources should mainly depict errors in the emission parameterizations, or a misrepresentation in the mineralogical composition of the larger Fe-containing soil particles.
- 1210 The overestimation around the tropics and the northern latitudes is further illustrated by a comparison of the model predictions with observations as a function of latitude (binned at 2°) (Fig. 10g-i). Although the dissolved Fe in the accumulation mode

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for both simulations (Fig. 10h and Fig. 10i), along with an overestimation around the tropics as for the accumulation DFe

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aerosols. All in all, the differences between the two model configurations are relatively small.

# 4 Discussion

The ocean is a critical component of the Earth's climate system and Fe plays a key role in the efficiency of the biological carbon pump. For this reason, accurate estimates of the bioavailable Fe inputs to the ocean are a prerequisite for climate

- simulations. Our work attempts at properly simulating the effects of atmospheric multiphase processes on the chemical sources and sinks of the Fe-containing aerosols in an Earth System Model. Indeed, the atmospheric processing of Fe-containing aerosols is rather important for the geographical pattern of DFe deposition fluxes into the ocean, especially in remote regions away from land sources. In agreement with other studies (e.g., Hamilton et al., 2022; Ito et al., 2019; Mahowald et al., 2005; Myriokefalitakis et al., 2018; Scanza et al., 2018), we find that mineral dust is the principal source of atmospheric Fe in EC-
- Earth (95 ± 1 %), with most of the remaining sources attributable to biomass burning. Focusing on the bioavailable fraction for marine biota, OXL aqueous-phase production is shown to be an important driver of aerosol DFe release, contributing ~44 % to the aerosol Fe dissolution, along with the atmospheric acidity that accounts for ~45 % of total DFe secondary sources in the model. Therefore, the realistic representation of the atmospheric OXL concentrations is a prerequisite to properly simulate the atmospheric mineral Fe dissolution processes.
- 290 Present-day simulations indicate that 61<u>816 ± 1.295</u> Tg yr<sup>-1</sup> of Fe in EC-Earth is deposited to the Earth's surface, which is towards the low end (40 140 Tg yr<sup>-1</sup>) of the model intercomparison study by Myriokefalitakis et al. (2018). The amount of total Fe deposited to the global ocean is calculated to be 12.937 ± 0.308 Tg yr<sup>-1</sup> in EC-Earth, which is about 50 % lower than recent estimates by Hamilton et al. (2019), owing to the significantly larger (almost double) mineral dust emission flux in that study. However, the large variability in global models can be partly attributed to the different mineral dust size ranges
- 1295 considered in the models. Indeed, since most of the Fe mass is associated with coarse dust aerosols (~91 % in this work), models that additionally account for super-coarse mineral dust emission sources (i.e., > 10  $\mu$ m in diameter), eventually calculate a higher TFe source and thus increased TFe global deposition rates. In our EC-Earth simulations, roughly 0.878 ± 0.015 Tg yr<sup>-1</sup> of DFe are calculated to be deposited globally (Table 2), in the range of estimates presented in Myriokefalitakis et al. (2018) (0.8 ± 0.2 Tg yr<sup>-1</sup>). Focusing on the marine environment, about 40 % (0.376 ± 0.005 Tg yr<sup>-1</sup>) of the simulated DFe
- is deposited into the global ocean, indicating that a large fraction of Fe atmospheric inputs to the global ocean results from the dissolution of atmospheric aerosols. Our <u>results</u> are close to the high-end of other global estimates  $(0.173 0.419 \text{ Tg yr}^{-1})$  as presented in the model intercomparison study of Myriokefalitakis et al. (2018), slightly higher than the respective DFe

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deposition fluxes in Ito et al. (2021) <u>(0.271 Tg yr<sup>-1</sup>)</u>, but somehow lower compared to Hamilton et al. (2019) estimates (~0.5
 320 Tg yr<sup>-1</sup>) although the significant differences in the dust emission fluxes. Thus, even though we do not consider a super-coarse mode of dust in our simulations, the DFe deposition rates over the remote ocean are not severely impacted (Myriokefalitakis et al., 2018) by the size of the emitted minerals, but instead by the atmospheric processing during long-range transport. Nevertheless, reaching a firm conclusion in that respect will need further work.

- The Fe-containing dust aerosols dominate (~70.%) the total deposition fluxes over the ocean in the model, although combustion sources are calculated to have a significant impact <u>on the Fe inputs to</u> remote oceanic regions, such as the Pacific and the Southern Oceans <u>in agreement with other studies (e.g.</u>, Hamilton et al., 2020), The average Fe solubility at the deposition of combustion aerosols is ~19\_% (Fig. 11a), much higher compared to the solubility of mineral dust aerosols (~2\_%; Fig. 11b), clearly indicating the importance of atmospheric processing on the potential bioavailable inputs to the global ocean. We further
- note that although here a relatively high Fe solubility is applied for ship oil combustion emission, a sensitivity simulation (not
   shown) using an iron solubility for ship oil emissions of 47.5 %, as proposed by Rathod et al<sub>x</sub>(2020), leads overall to an only slight decrease (up to ~2 %) in iron solubility in deposition, mainly in the northern Atlantic Ocean, and does not substantially affect our results. Indeed, our simulations show high Fe solubilities far from continental regions (Fig. S6g-i), such as the tropical Pacific and Atlantic Oceans (Fig. 11c), due to aerosol aging and lower Fe concentrations. Overall, the maximum DFe deposition fluxes occur in EC-Earth downwind of the main desert source regions, with high deposition rates being also
- 1335 simulated in the outflow of tropical biomass burning regions (such as South America, Africa, and Indonesia), as well as over highly populated regions due to the Fe released from anthropogenic combustion processes in the presence of polluted air masses (such as in India and China).

#### 5 Summary and conclusions

This work documents the implementation of a detailed multiphase chemistry scheme in the EC-Earth3 Earth system model,
aiming to provide consistent estimates of the atmospheric concentrations of the Fe-containing aerosols, along with the species that modulate its atmospheric processing, i.e., OXL and SO4<sup>2-</sup>. For this, a comprehensive description of the atmospheric Fe cycle is included in the model, accounting for 1) an explicit soil mineralogy, 2) the contribution of combustion emissions, and 3) an atmospheric dissolution scheme that accounts for atmospheric acidity, ambient levels of OXL, and photoinduced processes. The multiphase chemistry scheme simulates the aqueous-phase processes of the troposphere for inorganic and organic compounds, along with the Fenton reaction. The KPP software is used in the model to integrate the aqueous-phase and the dissolution equations, which adds flexibility to the code. Overall, simulations of tropospheric chemistry and aerosols for present-day conditions (2000-2014) have been realized, and budget calculations for OXL, SO4<sup>2-</sup>, and the DFe-containing aerosols have been presented.

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- 355 Model simulations have been performed both as a coupled system with IFS, as well as driven by offline meteorological fields from the ERA-Interim reanalysis. Budget analysis has shown that glyoxal is the main precursor of OXL in the atmosphere (~74 %), and that the potential primary sources in the model (0.373  $\pm$  0.005 Tg yr<sup>-1</sup>) have a negligible impact on OXL concentrations. Non-traditional, laboratory-derived, aqueous-phase production pathways of OXL via glyoxal oxidation are also accounted for in our simulations. We have shown that when such pathways are omitted in the simulations, the calculated 360 global OXL atmospheric concentrations are substantially lowered (~43 %). For the ERA-Interim set-up, the OXL net chemical
- production is calculated as  $18.116 \pm 0.071$  Tg yr<sup>1</sup>, resulting in an atmospheric burden of 0.330 Tg on average. For the onlinecoupled system, however, the OXL net chemical production is ~30 % lower, mainly attributed to a lower atmospheric OH abundance in EC-Earth due to biases that can be generally found in climate-chemistry models, such as in temperature and humidity especially at the higher altitudes. Overall, the simulated oxidizing capacity along with the contribution of secondary
- 365 sources - other than the GLX oxidation - are shown to have a significant impact on the OXL atmospheric abundance. However, we acknowledge that other formation pathways of OXL, primary or secondary, may exist in the gas- and the aqueous phases of the atmosphere that could further contribute to OXL levels.

The dissolution of dust and combustion aerosols dominates on the dissolved Fe fraction in the model, calculated in EC-Earth at  $0.806 \pm 0.014 \text{ Tgyr}^{-1}$ , in good agreement with the ERA-Interim simulation and well in the range of other model estimates. 1370 Furthermore, a broad evaluation of the EC-Earth proves the models' ability to represent AOD, particularly over regions that are dominated by anthropogenic aerosol, but also over selected dusty sites, in line with previous EC-Earth evaluations. However, dust is underestimated over remote areas. The model also underestimates the *in-situ* OXL measurements, especially during winter, indicating nevertheless that additional sources (primary and/or secondary) are needed. Model comparisons with cruise measurements demonstrate a strong link between atmospheric DFe concentrations and atmospheric composition. The

1375 model seems to better capture the observations in the accumulation mode than in the coarse mode. This is attributed to differences either in the atmospheric processing between the accumulation and coarse particles (i.e., aerosol water content and acidity levels), the misrepresentation in the aerosol sizes (e.g., surface area/volume ratios), or to systematic errors in the mineralogical composition of the emitted Fe-containing soil particles. For this, several developments are planned by the EC-Earth consortium, aiming to improve the representation of the dust aerosols' size distribution, the description of the 1380 mineralogical composition, and the Fe-content in combustion sources, which are expected to reduce the existing uncertainties in the model for a more accurate simulation of the atmospheric Fe cvcle.

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Emphasizing on the biogeochemistry-related implications of this study, EC-Earth calculates a global annual present-day DFe deposition flux of  $0.878 \pm 0.015$  Tg yr<sup>-1</sup>, well within the range of estimates of other global modeling studies. About 40 % of the DFe deposition fluxes are calculated to occur over the global ocean  $(0.376 \pm 0.005 \text{ Tg yr}^{-1})$ , with a strong spatial and temporal variability. The highest annual mean DFe inputs to the global ocean are associated with aerosols of soil origin, especially downwind of the major dust source regions. In addition, Fe-containing combustion aerosols are calculated to have

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a significant contribution downwind of biomass burning source regions and highly populated areas in the Northern Hemisphere. It is further demonstrated that over the open ocean the Fe solubility in deposition for aerosols of combustion origin is about an order of magnitude higher than that of mineral dust origin, indicating that the relative contributions of the primary sources can significantly affect bioavailable aerosol fraction and may thus play an important role in oceanic areas where the phytoplankton growth is limited by Fe supply, such as the Southern Ocean.

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It is widely recognized that a combined approach considering both Fe atmospheric processing and deposition over oceans should ideally be used in Earth System Models for the assessment of the impact of nutrient-containing aerosol deposition on marine productivity. A deeper understanding of the atmospheric Fe cycle is thus needed for a better description of the biogeochemistry implications in the presence of a changing climate. Such type of knowledge, however, should be obtained by

- 1415 extensive model evaluation with observations, especially over the remote regions of the world like the Southern Ocean, where currently the largest discrepancies between models and measurements exist. Therefore, a comprehensive calculation of the Fe physicochemical transformations is necessary to predict the strength of DFe inputs to the ocean, despite the complexity of the related atmospheric multiphase processes. The present study <u>aims to</u> complement, thus, the marine biogeochemistry component of EC-Earth by a fully coupled calculation scheme for atmospheric dissolved Fe fluxes into the global ocean. That
- 420 <u>new</u> EC-Earth model version is expected to eventually allow for a better representation of the marine biogeochemistry perturbations in past and future climates and air quality.

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## Appendix A

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Table A1. Summary of statistics for all points and per region (as depicted in Fig. 1) for the evaluation of a) the modeled annual mean AOD at 550 nm against AERONET version 3 level 2.0 retrievals for all available stations covering the 2000-2014 period, b) the same but for selected dust-dominated AERONET sites (characterized as described in Sect. 2.5), c) the modeled annual mean dust surface concentration for 2000-2014 compared to climatological mean values from RSMA sites and AMMA campaign, and d) the modeled annual dust deposition flux averaged for the period 2000-2014 against observations as compiled in Albani et al. (2014) from several sources. The number of stations (n), the Pearsons' correlation coefficients (R) between the simulated and measured monthly mean concentrations, the normalized mean bias (nMB), and the normalized root mean square errors (nRMSE) are indicated for the EC-Earth simulation.

	n	R	nMB (%)	nRSME (%)
a) AOD(550 nm)				
N. America	208		-8.8	40.9
C. America	41		-2.8	33.1
S. America	41		-13.3	54.2
Europe	159		-9.9	26.9
N. Africa	27		21.6	44.6
S. Africa	15		-35.0	61.8
W. Asia	49		-21.5	36.6
E. Asia	162		-8.9	38.4
Australian Ocean	26		-21.7	110.8
Remote oceans	10		22.9	54.2
All points	738	0.8	-9.1	45.7
b) AOD(550 nm) - dust-domin				
C. America	2		-26.5	49.7
N. Africa	18		21.5	43.2
S. Africa	1		-15.4	15.4
W. Asia	12		-22.5	29.5
E. Asia	5		-8.7	14.5
All sites	38	0.76	2.7	37.2
c) Dust surface concentrations				
N. America	1		-74.2	74.2
C. America	2		-37.5	38.7
Europe	2		-90.8	116.6
N. Africa	4		21.9	37.7
E. Asia	2		71.8	82.2
Australian Ocean	3		-42.0	81.5
S. Pacific Ocean	3		-60.3	68.5
N. Pacific Ocean	4		89.9	132.7
Southern Ocean	2		-98.5	112.7
All points	23	0.99	19.3	81.2
d) Dust deposition rates				
N. America	7		-96.8	130.6
C. America	3		-59.6	97.6
S. America	1		-99.6	99.6
Europe	14		-94.5	137.1
N. Africa	23		-56.4	142.2
S. Africa	4		-91.2	109.3
W. Asia	5		-94.7	164.4
E. Asia	14		-25.2	94
Australian Ocean	9		-98.3	201.1
S. Pacific Ocean	2		-92.0	93.2
N. Pacific Ocean	13		90.5	148.9
Southern Ocean	15		-91.9	205.0
All points	110	0.75	-61.2	204.0

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Table A2. Summary of statistics for the evaluation of the simulated concentrations (as depicted in Fig. 6) of a) oxalate, b) sulfate, aswell as for c,d,e) the dissolved, f,g,h) the total Fe-containing aerosols, and i,j,k) the derived aerosol solubility (SFe=%DFe/TFe) forc,f,i) the accumulation mode, d,g,j) the coarse mode, and e,h,k) the total suspended particles (tsp), respectively. The number of1440stations (n), the Pearsons' correlation coefficients (R) between the simulated and the measured concentrations, the normalized meanbias (nMB) and the normalized root mean square errors (nRMSE) are indicated for the EC-Earth and ERA-Interim simulation.The results for the sensitivity simulation EC-Earth(sens) for oxalate evaluation, as well as, the EC-Earth(AerChem-AMIP) for thesulfate evaluation are also shown.

	n	R	nMB (%)	nRSME (%)
a) Oxalate				
EC-Earth		0.48	-64.18	116.97
ERA-Interim	143	0.45	-45.96	110.31
ERA-Interim(sens)		0.44	-73.66	124.89
b) Sulfate				
EC-Earth		0.76	15.89	55.20
ERA-Interim	3828	0.75	22.83	57.19
ERA-Interim(AerChem-AMIP)		0.70	-4.69	57.05
c) Dissolved Fe (accumulation mode	2)			
EC-Earth	438	0.49	38.01	164.05
ERA-Interim	438	0.58	23.43	145.53
d) Dissolved Fe (coarse mode)				
EC-Earth	439	0.46	-8.70	193.51
ERA-Interim	439	0.59	-25.46	177.45
e) Dissolved Fe (tsp)				
EC-Earth	955	0.27	30.23	274.96
ERA-Interim	955	0.29	20.59	265.68
f) Total Fe (accumulation mode)				
EC-Earth	92	0.59	-23.34	166.29
ERA-Interim	92	0.71	-46.55	175.23
g) Total Fe (coarse mode)				
EC-Earth	83	0.58	4.41	136.43
ERA-Interim	85	0.54	-21.13	141.46
h) Total Fe (tsp)				
EC-Earth	796	0.44	29.39	322.75
ERA-Interim	/90	0.30	28.24	414.24
i) Fe Solubility (accumulation mode	2)			
EC-Earth	92	0.19	19.57	110.92
ERA-Interim	92	0.24	33.07	117.36
j) Fe Solubility (coarse mode)				
EC-Earth	07	0.12	0.61	143.51
ERA-Interim	83	0.17	4.94	133.74
k) Fe Solubility (tsp)				
EC-Earth	402	0.21	596.47	880.99
ERA-Interim	483	0.23	607.00	921.44

#### Code availability

The EC-Earth3-Iron code is available from the EC-Earth development portal (https://dev.ec-earth.org/) for members of the consortium. EC-Earth3-Iron adds new features to the EC-Earth-AerChem version 3.3.2.1, which includes the IFS cycle 36r4, the NEMO-LIM3 release 3.6, the TM5-MP 3.0, and makes use of the OASIS3-MCT version 3.0 coupler. An AMIP reader is

450 also available to use ocean prescribed data. Model codes developed at ECMWF, including the atmosphere model IFS, are intellectual property of ECMWF and its member states. Permission to access the EC-Earth3-Iron source code can be requested from the EC-Earth community via the EC-Earth website (http://www.ec-earth.org/) and may be granted if a corresponding software license agreement is signed with ECMWF. The corresponding repository tag is 3.3.2.1-Fe. Currently, only European users can be granted access, due to license limitations of the model.

## 1455 Data availability

The Aerosol Robotic Network retrievals of optical depth were downloaded through the AERONET data download tool (available at: https://aeronet.gsfc.nasa.gov/; last access 28/03/2020). The derived climatologies used for model evaluation in Fig. 4 have been permanently stored in the zenodo repository, accessible through https://doi.org/10.5281/zenodo.5776347, Access to the observations of dust concentration from the AMMA campaign can be requested through the INDAAF website (https://indaaf.obs-mip.fr/; last access 01/05/2021). Other observational datasets used for evaluation are available in the referenced articles. ERA-Interim data is available through the ECMWF data download site (https://apps.ecmwf.int/datasets/data/interim-full-daily/levtype=sfc/; last access 02/12/2019), and the gridded emission datasets, as well as ocean surface conditions used in this work, are available through the input4MIPs download tool (https://esgf-node.llnl.gov/projects/input4mips/; last access 02/12/2019). The model outputs relevant for this study are permanently stored in the zenodo repository, accessible through: https://doi.org/10.5281/zenodo.5752596,

#### Author contribution

SM developed the aqueous-phase chemistry and the iron-dissolution schemes, designed the experiments, and performed the ERA-Interim simulations. EBM, MGA, and CPGP developed the mineralogy applied to dust emissions, co-designed the experiments, and performed the EC-Earth simulations. AI provided the Fe-containing combustion aerosol emissions. EA contributed to the emission parameterizations. AN provided the ISORROPIA II code. Other co-authors contributed to the development of specific aspects of the model or parts of the code that are shared with the EC-Earth3 community. SM wrote the paper, with the contribution of MGA, and inputs from all co-authors.

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#### **Competing interests**

The authors declare that they have no conflict of interest.

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510 IRD and the Observatoire Midi-Pyrénées and the Observatoire des Sciences de l'Univers EFLUVE. Model development was

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## **Table and Figures**

# $2050 \qquad {\rm Table 1. Overview of the simulations performed for this study.}$

Simulation	Description
	Aqueous-phase chemistry scheme for simulating OXL production and Fe dissolution, coupled to the MOGUNTIA gas-phase
EC-Earth	chemistry scheme. Meteorology calculated online by IFS and observed sea surface temperature and sea ice concentration
	boundary conditions (AMIP-CMIP6) are applied.
ERA-Interim	As for EC-Earth simulation but driven by meteorological data from the ECMWF reanalysis ERA-Interim.
ERA-Interim(sens)	As for ERA-Interim simulation but neglecting the contribution of glyoxal high molecular weight species on GLX and OXL
EKA-Interini(sens)	formation.



	EC-Earth	Era-Interim	Era-Interim(sens)
a) Oxalate			
Emissions (Tg yr <sup>-1</sup> )		0.373	
Chemistry Production (Tg yr <sup>-1</sup> )			
GLYOLI + OH	10.597	14.764	-
MGLY + OH	1.079	1.415	1.552
$GLX + OH/NO_3$	3.369	4.618	9.962
Chemistry Loss (Tg yr <sup>-1</sup> )			
$OXL + OH/NO_3$	1.019	1.189	0.822
$[Fe(OXL)_2]^* + hv$	1.412	1.475	0.680
Deposition (Tg yr <sup>-1</sup> )			
Dry Deposition	0.134	0.176	0.097
Wet Scavenging	12.850	18.313	10.286
Atmospheric Burden (Tg)	0.219	0.330	0.189
Lifetime (days)	5.175	5.691	5.810
b) Sulfate	5.175	5.091	5.810
Emissions (Tg S yr <sup>-1</sup> )		1.593	
H <sub>2</sub> SO <sub>4</sub> Chemistry Production (Tg S yr <sup>-1</sup> )		1.595	
$SO_2 + OH$	11.976	11.088	
S(VI) Chemistry Production (Tg S yr <sup>-1</sup> )	11.9/0	11.000	
S(VI) Chemistry Production (1g S yr <sup>-</sup> ) $S(IV) + H_2O_2$	32.902	35.812	
	32.902 5.927	35.812 4.760	
$S(IV) + O_3$			
$S(IV) + HO_2$	0.004	0.004	
$S(IV) + CH_3O_2H$	0.051	0.049	
Deposition (Tg S yr <sup>-1</sup> )			
Dry Deposition	3.079	2.912	
Wet Scavenging	49.368	50.394	
Atmospheric Burden (Tg S)	0.692	0.961	
Lifetime (days)	4.816	6.579	
e) Dissolved FeC			
Emissions (Tg yr <sup>-1</sup> )		0.012	
Dissolution (Tg yr <sup>-1</sup> )			
FeC + H+	0.047	0.049	0.049
FeC + OXL	0.182	0.188	0.183
FeC + hv	0.045	0.047	0.046
Deposition (Tg yr <sup>-1</sup> )			
Dry Deposition	0.081	0.080	0.077
Wet Scavenging	0.206	0.217	0.212
Atmospheric Burden (Tg)	0.002	0.003	0.003
Lifetime (days)	2.970	4.163	4.203
l) Dissolved FeD			
Emissions (Tg yr <sup>-1</sup> )	0.059	0.049	0.049
Dissolution (Tg yr <sup>-1</sup> )			
$FeD + H^+$	0.315	0.311	0.311
FeD + OXL	0.170	0.168	0.164
FeD + hv	0.047	0.049	0.048
Deposition (Tg yr <sup>-1</sup> )			
Dry Deposition	0.140	0.132	0.130
Wet Scavenging	0.452	0.446	0.441
Atmospheric Burden (Tg)	0.006	0.010	0.010
Lifetime (days)	3.837	6.236	6.264
	-1021	5.250	0.201

 Table 2. Global budgets, atmospheric burdens, and lifetimes, averaged for the period 2000-2014, of a) oxalate (OXL), b) sulfate as well as, for the dissolved Fe-containing aerosols from c) combustion processes (FeC) and d) mineral dust (FeD) for the EC-Earth,

 2055
 ERA-Interim, and ERA-Interim(sens) simulations.

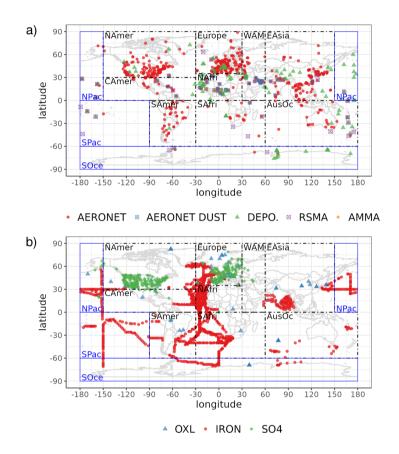
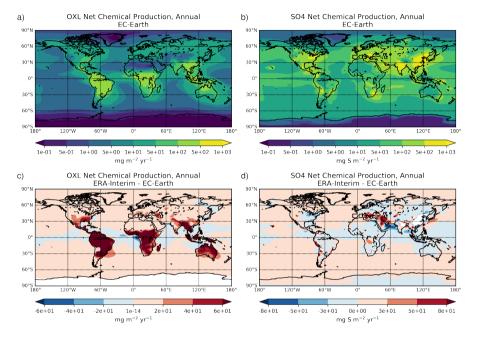


Figure 1: Site location map of observations a) for AOD (AERONET; red dots, AERONET-DUST; blue squares), dust surface 2060 concentration (RSMA; purple squares, and AMMA; orange diamonds) and dust deposition rates (several sources compiled in Albani et al. (2014); green triangles) and b) for surface oxalate (OXL; blue triangles), cruise aerosol Fe concentrations (red circles), and sulfate (green diamonds).



2065 Figure 2: Annual mean net chemical production rates for a) oxalate (mg m<sup>-2</sup> yr<sup>-1</sup>) and d) sulfate (mg S m<sup>-2</sup> yr<sup>-1</sup>) as calculated for the EC-Earth simulation averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (c,d).



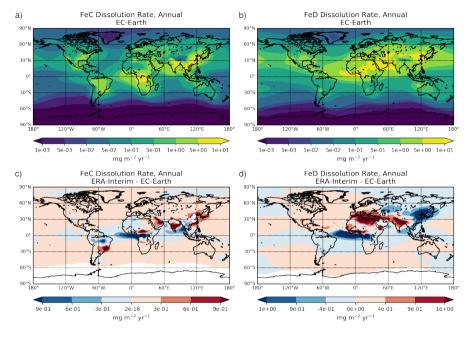


Figure 3: Annual mean dissolution rates (mg m<sup>-2</sup> yr<sup>-1</sup>) of combustion (a) and mineral dust (b) aerosols as calculated for the EC-Earth 2070 simulation averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (c,d).



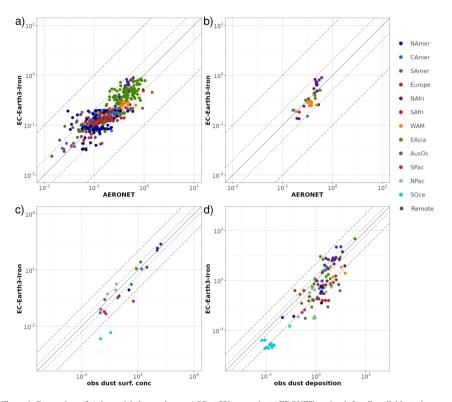


Figure 4: Comparison of a) the modeled annual mean AOD at 550 nm against AERONET retrievals for all available stations covering the 2000-2014 period, b) same but for selected dusty AERONET sites, c) the modeled annual mean dust surface concentration for 2000-2014 compared to climatological mean values from RSMA sites and AMMA campaign, and d) modeled annual dust deposition flux averaged for the period 2000-2014 against observations compiled in Albani et al. (2014) from several sources.

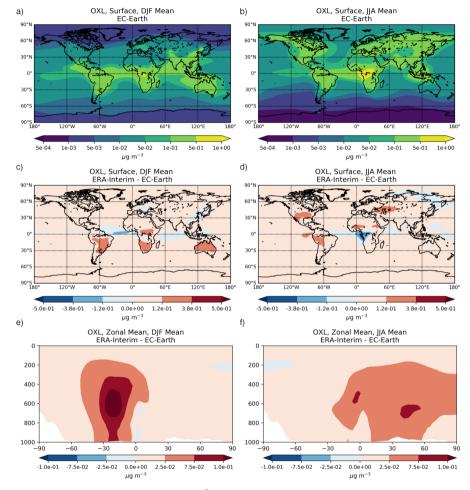


Figure 5: Oxalate (OXL) surface concentrations (µg m<sup>3</sup>) for the boreal winter (DJF; a) and boreal summer (JJA; b) as simulated for the EC-Earth simulation averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim 2080 simulation for surface (c,d) and zonal mean (e,f).

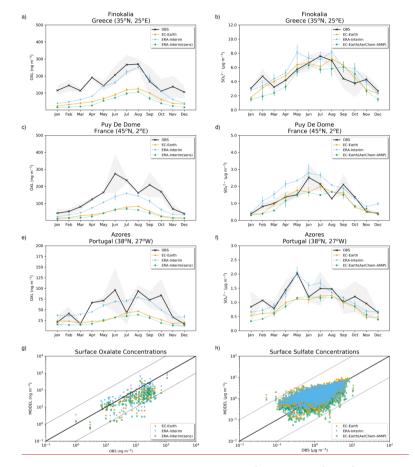


Figure 6: Comparison of daily mean observations (black line) of OXL (ng m<sup>-3</sup>; left) and nss-SO<sub>4</sub><sup>-2</sup> (µg m<sup>-3</sup>; right) with the EC-Earth (orange line/circles) and the ERA-Interim (light blue line/triangles) simulations for a,b) Finokalia (Greece) for the period July 2004
 July 2006 (Koulouri et al., 2008), for c,d) Puy de Dome (France) and e,f) Azores (Portugal), for the period September 2002 September 2004 (Legrand et al., 2007), along with d,f) scatterplot comparisons for observations around the globe; the solid line represents the 1: 1 correspondence and the dashed lines show the 10: 1 and 1: 10 relationships, respectively. For completeness, the comparisons for the sensitivity simulation ERA-Interim(sens) for OXL and the EC-Earth(AerChem-AMIP) (green line/squares) for sulfate are also presented. Gray shaded areas represent the standard deviation of the observations and the color-coded error bars represent the model's standard error of the multi-annual mean for the individual observational period.

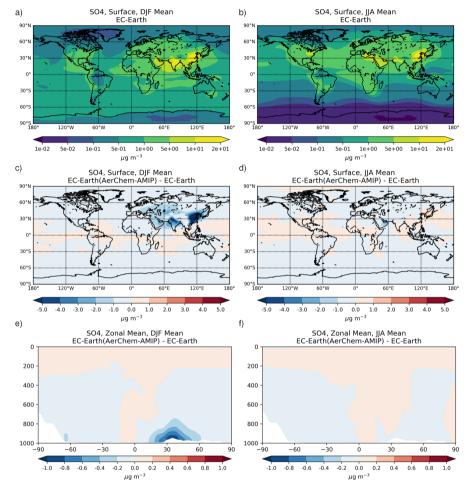
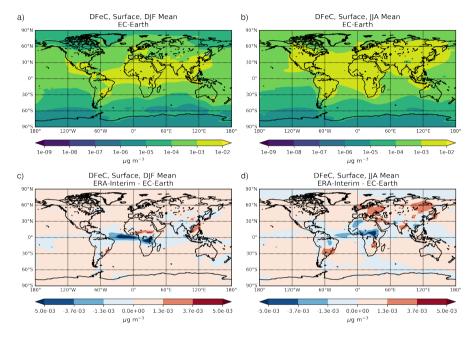
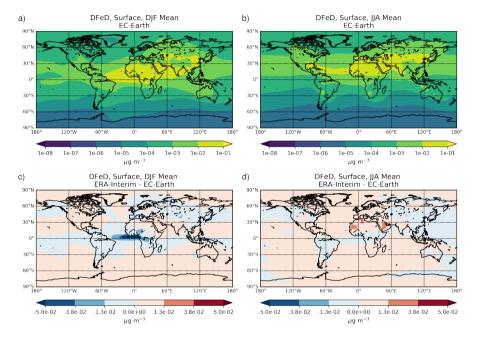




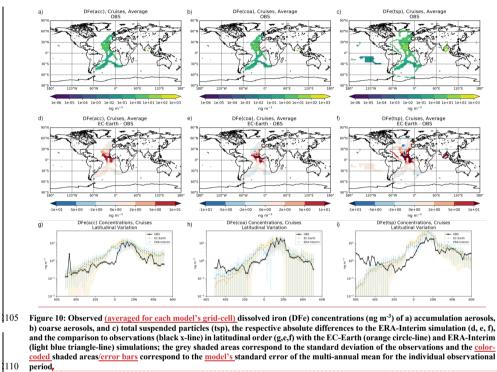
Figure 7: Sulfate (SO4) surface concentrations ( $\mu$ g m<sup>-3</sup>) for the boreal winter (DJF; a) and boreal summer (JJA; b) as simulated for the EC-Earth simulation averaged for the period 2000-2014, and the respective absolute differences to the EC-Earth(AerChem-AMIP) simulation for surface (c,d) and zonal mean (e,f).



2095 Figure 8: Dissolved iron surface concentrations (μg m<sup>-3</sup>) from combustion aerosols (DFeC) for the boreal winter (DJF; a) and summer (JJA; b) seasons for the EC-Earth simulation averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (c,d).



2100 Figure 9: Dissolved iron surface concentrations (μg m<sup>-3</sup>) from mineral dust (DFeD) for the boreal winter (DJF; a) and summer (JJA; b) seasons for the EC-Earth simulation averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (c,d).



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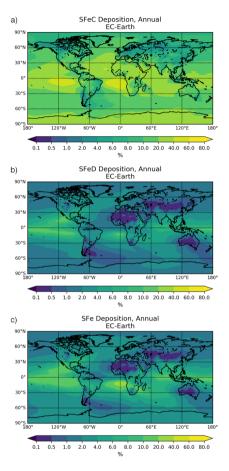


Figure 11: Annual mean Fe-containing aerosol solubility at deposition fluxes (%) as simulated for the EC-Earth simulation averaged for the period 2000-2014, for a) mineral dust aerosols, b) sum of solid fuel combustion, liquid fuel combustion and open biomass burning aerosols, and c) sum of all aerosol sources.



## Supplementary Tables

Table S1: Averaged factors for the vears 2000-2014 used to represent the Fe-containing aerosols in the emitted fine and coarse aerosols of the model, as applied to a) the calculated dust mineral emissions as derived from the updated mineralogy maps originally created by Claquin et al., (1999), b) the CMIP6 anthropogenic emission sectors (Hoesly et al., 2018) as retrieved based on estimates from Ito et al. (2018) for the emitted submicron carbonaceous particulate matter (i.e., sum of OC and BC, and inorganic matters), and c) the CMIP6 biomass burning emissions (van Marle et al., 2017) also based on Ito et al. (2018) estimates. In parentheses, the standard deviation (where available) is also provided.

Fe types	Fine aerosols	Coarse <u>aerosols</u>
a) Dust minerals		4
Illite	0.048	
Smectite	0.164	
Kaolinite	0.007	
Feldspars	0.025	
Iron oxides	<u>0.660</u>	
b) Anthropogenic sectors		4
Energy	$1.861 \cdot 10^{-2} (\pm 6.281 \cdot 10^{-4})$	$3.528 \cdot 10^{-1} (\pm 1.190 \cdot 10^{-2})$
Industrial	$1.459 \cdot 10^{-2} (\pm 1.414 \cdot 10^{-3})$	$2.764 \cdot 10^{-1} (\pm 2.681 \cdot 10^{-2})$
Residential & Commercial	4.3 <u>36 10.6(±3.485 • 10.7)</u>	$1.350 \cdot 10^{-5} (\pm 1.085 \cdot 10^{-6})$
Shipping	$3.751 \cdot 10^{-2} (\pm 9.043 \cdot 10^{-9})$	$1.398 \cdot 10^{-2} (\pm 4.454 \cdot 10^{-9})$
<b>T</b>		<b>T</b>
Waste treatment	1.389 • 10 <sup>-3</sup> (± 9.622 • 10 <sup>-11</sup> )	$3.595 \cdot 10^{-3} (\pm 8.487 \cdot 10^{-10})$
<b>V</b>	<b>T</b>	Χ
c) Biomass Burning	6.292 • 10 <sup>-3</sup>	2.296 • 10 <sup>-2</sup>

Table S2. Aqueous phase chemical mechanism and reaction rate constants. The photolysis frequencies (J) are expressed in  $s^{-1}$ , and the aqueous reactions (K) are expressed in L mol<sup>-1</sup>  $s^{-1}$ .

	Aqueous-phase Reactions			Jmax/K298	E <sub>s</sub> /R	Reference
	Photolysis					
J01	$O_3 + hv$ (+ H <sub>2</sub> O)	$\rightarrow$	$H_2O_2 + O_2$	a		
J02	$H_2O_2 + hv$	$\rightarrow$	2 OH	a		
J03	$NO_3 + hv$	$\rightarrow$	$NO + O_2$	a		
J04	$NO_3 + hv$	$\rightarrow$	$NO_2 + O_3$	a		
J05	$NO_2 + hv (+ H^+)$	$\rightarrow$	NO + OH	a		
J06	$NO_3 + hv (+ H^+)$	$\rightarrow$	$NO_2 + OH$	a		
J07	$CH_{3}O_{2}H + hv (+ O_{2})$	$\rightarrow$	$CH_2(OH)_2 + HO_2 + OH$	a		
J08	$Fe^{+3} + hv (+ H_2O)$	$\rightarrow$	$Fe^{+2} + OH + H^+$	6.41 10 <sup>-6</sup> b		Deguillaume et al. (2004)
J09	$[Fe(OH)]^{2+} + hv$	$\rightarrow$	Fe <sup>2+</sup> + OH	5.63 10 <sup>-3 b</sup>		Deguillaume et al. (2004)
J10	$[Fe(OH)_2]^+ + hv$	$\rightarrow$	Fe2+ + OH + HO-	7.52 10 <sup>-3 b</sup>		Deguillaume et al. (2004)
J11	$[Fe(SO_4)]^+ + hv (+ H_2O)$	$\rightarrow$	Fe2++ SO42+ OH + H+	4.51 10 <sup>-5</sup> b		Deguillaume et al. (2004)
J12	$[Fe(OXL)_2] + hv (+ O_2)$	$\rightarrow$	Fe <sup>+2</sup> + OXL <sup>2-</sup> + O <sub>2</sub> <sup>-</sup> + 2 CO <sub>2</sub>	2.47 10 <sup>-2 b</sup>		Ervens et al. (2003)
	H <sub>x</sub> O <sub>y</sub> chemistry					
K001	$O_3 + OH$	$\rightarrow$	$HO_2 + O_2$	1.1 108		Deguillaume et al. (2010)
K002	$O_3 + HO_2$	$\rightarrow$	OH + 2 O <sub>2</sub>	1.0 104		Deguillaume et al. (2010)
K003	$O_3 + O_2 + H_2O$	$\rightarrow$	$OH + O_2 + HO$	1.5 10 <sup>9</sup>	2200	Deguillaume et al. (2010)
K004	HO <sub>2</sub> + OH	$\rightarrow$	$O_2 + H_2O$	1.0 10 <sup>10</sup>		Ervens et al. (2003)
K005	$O_2^- + OH$	$\rightarrow$	$O_2 + HO^2$	1.1 10 <sup>10</sup>	2120	Ervens et al. (2003)
K006	$HO_2 + HO_2$	$\rightarrow$	$H_2O_2 + O_2$	8.3 10 <sup>5</sup>	2720	Ervens et al. (2003)
K007	$HO_2 + O_2^- (+ H^+)$	$\rightarrow$	$H_2O_2 + O_2$	9.7 10 <sup>7</sup>	1060	Ervens et al. (2003)
K008	OH + OH	$\rightarrow$	H <sub>2</sub> O <sub>2</sub>	3.6 10 <sup>9</sup>	930	Deguillaume et al. (2010)
K009	$H_2O_2 + OH$	$\rightarrow$	$HO_2 + H_2O$	3.0 107	1680	Ervens et al. (2003)
	Nitrogen chemistry					
K010	NO + OH	$\rightarrow$	NO2 <sup>-</sup> + H <sup>+</sup>	2.2 10 <sup>10</sup>	1500	Deguillaume et al. (2010)
K011	NO <sub>2</sub> + NO (+ H <sub>2</sub> O)	$\rightarrow$	2 NO <sup>-</sup> + 2 H <sup>+</sup>	3.0 10 <sup>8</sup>		Deguillaume et al. (2010)
K012	$NO_2 + NO_2 (+ H_2O)$	$\rightarrow$	$HONO + NO_3 + H^+$	8.4 107	-2900	Ervens et al. (2003)
K013	$NO_2 + OH$	$\rightarrow$	NO3 <sup>-</sup> + H <sup>+</sup>	1.2 10 <sup>10</sup>		Ervens et al. (2003)
K014	$NO_3 + HO_2$	$\rightarrow$	$NO_{3}^{-} + O_{2} + H^{+}$	3.0 10 <sup>9</sup>		Ervens et al. (2003)
K015	$NO_3 + O_2$	$\rightarrow$	$NO_3 + O_2$	3.0 10 <sup>9</sup>		Ervens et al. (2003)
K016	$NO_3 + H_2O_2$	$\rightarrow$	$NO_{3}^{-} + HO_{2} + H^{+}$	4.9 106	2000	Deguillaume et al. (2004)
K017	NO <sub>3</sub> + HO <sup>-</sup>	$\rightarrow$	NO3 <sup>-+</sup> OH	9.4 107	2700	Ervens et al. (2003)
K018	$NO_2 + O_3$	$\rightarrow$	$NO_3 + O_2$	5.0 10 <sup>5</sup>	7000	Ervens et al. (2003)
K019	HONO + OH	$\rightarrow$	$NO_2 + H_2O$	1.0 1010		Deguillaume et al. (2010)
K020	$NO_2 + OH$		NO <sub>2</sub> + HO <sup>-</sup>	1.1 1010		Ervens et al. (2003)
K021	NO2 <sup>-</sup> + CO3 <sup>-</sup>	$\rightarrow$	NO <sub>2</sub> +CO <sub>3</sub> <sup>2-</sup>	6.6 10 <sup>5</sup>	850	Ervens et al. (2003)
	Sulfur chemistry					
K022	$SO_2 + O_3 (+ H_2O)$	$\rightarrow$	$HSO_4^{-} + O_2 + H^+$	$2.4 \ 10^4$		Seinfeld and Pandis (2006)
K023	$HSO_3 + O_3$	$\rightarrow$	$HSO_4 + O_2$	3.7 10 <sup>5</sup>	5530	Seinfeld and Pandis (2006)
K024	$SO_3^{2-} + O_3$	$\rightarrow$	$SO_4^{2-} + O_2$	1.5 10 <sup>9</sup>	5280	Seinfeld and Pandis (2006)
K025	$HSO_{3} + H_{2}O_{2}(+ H^{+})$	$\rightarrow$	$H_2SO_4 + H_2O$	s		Seinfeld and Pandis (2006)
K026	$SO_2 + HO_2 (+ H_2O)$	$\rightarrow$	$H_2SO_4 + OH$	1.0 106		Seinfeld and Pandis (2006)
K027	SO <sub>2</sub> + O <sub>2</sub> (+ H <sub>2</sub> O)	$\rightarrow$	HSO3 <sup>-+</sup> OH + HO <sup>-</sup>	1.0 10 <sup>5</sup>		Seinfeld and Pandis (2006)
K028	HSO <sub>3</sub> <sup>-</sup> + CH <sub>3</sub> O <sub>2</sub> H (+ H <sup>+</sup> )	$\rightarrow$	$SO_4^{2-} + CH_3OH + 2H^+$	1.7 107	3160	Seinfeld and Pandis (2006)
	Carbonate chemistry					
K029	HCO3 <sup>-</sup> + OH	$\rightarrow$	$CO_3$ + H <sub>2</sub> O	1.7 107	1900	Ervens et al. (2003)
K039	CO32- + OH	$\rightarrow$	CO3 <sup>-</sup> + HO <sup>-</sup>	3.9 10 <sup>8</sup>	2840	Ervens et al. (2003)
K031	$HCO_3 + NO_3$	$\rightarrow$	CO3 <sup>+</sup> + NO3 <sup>+</sup> + H <sup>+</sup>	4.1 10 <sup>7</sup>		Ervens et al. (2003)

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as well as) on he CMIP6 anthropogenic emissi (Hoesly et al., 2018) as retrieved based on estimates	
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K032						
	CO <sub>3</sub> <sup>2</sup> + NO <sub>3</sub>	$\rightarrow$	CO3 <sup>-</sup> + NO3 <sup>-</sup>	4.1 107		Ervens et al. (2003)
K033	$CO_{3} + O_{3}$	$\rightarrow$	$CO_2 + O_2 + O_2$	1.0 105		Ervens et al. (2003)
K034	$CO_3 + HO_2$	$\rightarrow$	$HCO_3 + O_2$	6.5 10 <sup>8</sup>		Ervens et al. (2003)
K035	CO3 <sup>-</sup> + O2 <sup>-</sup>	$\rightarrow$	$CO_3^{2-} + O_2$	6.5 10 <sup>8</sup>		Ervens et al. (2003)
K036	$CO_{3} + H_{2}O_{2}$	$\rightarrow$	$HCO_3$ + $HO_2$	4.3 10 <sup>5</sup>		Ervens et al. (2003)
K037	$CO_3 + NO_2$	$\rightarrow$	$CO_2 + NO_3$	1.0 109		Ervens et al. (2003)
K038	CO3 <sup>-</sup> + CO3 <sup>-</sup> (+ O2)	$\rightarrow$	2 O <sub>2</sub> ·+ 2 CO <sub>2</sub>	2.2 106		Ervens et al. (2003)
	Organic chemistry					
K039	$CH_3O_2H + OH$	$\rightarrow$	0.8 (CH <sub>3</sub> O <sub>2</sub> + H <sub>2</sub> O) + 0.2 (HCOOH + HO <sub>2</sub> )	3.0 107	1680	Ervens et al. (2003)
K040	$CH_3O_2H + CO_3$	$\rightarrow$	CH <sub>3</sub> O <sub>2</sub> + HCO <sub>3</sub>	4.3 10 <sup>5</sup>		Ervens et al. (2003)
K041	$CH_3O_2 + HO_2$	$\rightarrow$	$CH_3O_2H + O_2$	4.2 10 <sup>5</sup>	3000	Deguillaume et al. (2010)
K042	$CH_{3}O_{2} + O_{2}(+H_{2}O)$	$\rightarrow$	$CH_3O_2H + O_2 + HO^2$	4.8 10 <sup>7</sup>	1600	Deguillaume et al. (2010)
K043	$CH_{3}O_{2} + CH_{3}O_{2}(+H_{2}O)$	$\rightarrow$	$CH_3OH + CH_2(OH)_2 + O_2$	1.7 10 <sup>8</sup>	2200	Ervens et al. (2003)
K044	CH <sub>3</sub> O <sub>2</sub> + CH <sub>3</sub> O <sub>2</sub> (+ 2 O <sub>2</sub> + 2	$\rightarrow$	2 CH <sub>2</sub> (OH) <sub>2</sub> +2 HO <sub>2</sub> +O <sub>2</sub>	3.6 107	2200	Ervens et al. (2003)
	H <sub>2</sub> O)	-				
K045	CH2(OH)2 + OH (+ O2)	$\rightarrow$	$HCOOH + HO_2 + H_2O$	1.0 109	1020	Ervens et al. (2003)
K046	$CH_2(OH)_2 + NO_3(+O_2)$	$\rightarrow$	$HCOOH + HO_2 + NO_3 + H^+$	1.0 106	4500	Ervens et al. (2003)
K047	CH <sub>3</sub> OH + OH (+ O <sub>2</sub> )	$\rightarrow$	$CH_2(OH)_2 + HO_2 + H_2O$	1.0 109	580	Ervens et al. (2003)
K048	$CH_{3}OH + NO_{3}(+O_{2})$	$\rightarrow$	$CH_2(OH)_2 + HO_2 + NO_3$	5.4 10 <sup>5</sup>	4300	Ervens et al. (2003)
K049	$CH_{3}OH + CO_{3}(+O_{2})$	$\rightarrow$	$CH_2(OH)_2 + HO_2 + HCO_3$	2.6 10 <sup>3</sup>		Ervens et al. (2003)
K050	$HCOOH + OH (+ O_2)$	$\rightarrow$	$CO_2 + HO_2 + H_2O$	1.3 10 <sup>8</sup>	1000	Ervens et al. (2003)
K051	$HCOO^{-} + OH (+ O_2)$	$\rightarrow$	$CO_2 + HO_2 + HO$	3.2 10 <sup>9</sup>	1000	Ervens et al. (2003)
K052	$HCOOH + NO_3 (+ O_2)$	$\rightarrow$	$CO_2 + HO_2 + NO_3 + H^+$	3.8 10 <sup>5</sup>	3400	Ervens et al. (2003)
K053	$HCOO + NO_3 + O_2$	$\rightarrow$	$CO_2 + HO_2 + NO_3$	5.1 107	2200	Ervens et al. (2003)
K054	HCOO <sup>-</sup> + CO <sub>3</sub> <sup>-</sup> (+ O <sub>2</sub> )	$\rightarrow$	$2 CO_2 + O_2 + HO^2$	1.4 10 <sup>5</sup>	3300	Ervens et al. (2003)
K055	CH <sub>3</sub> CH(OH) <sub>2</sub> + OH (+ O <sub>2</sub> )	$\rightarrow$	$CH_3COOH + HO_2 + H_2O$	1.2 109		Ervens et al. (2003)
K056	$CH_3CH(OH)_2 + NO_3(+O_2)$	$\rightarrow$	$CH_3COOH + HO_2 + NO_3 + H^+$	1.9 106		Ervens et al. (2003)
K057	$CH_2CH(OH)_2 + CO_3^2 (+ O_2)$	$\rightarrow$	$CH_3COOH + HO_2 + CO_2$	1.0 104		Ervens et al. (2003)
K058	GLYAL + OH (+ O2)	$\rightarrow$	$GLY + HO_2$	1.0 109		Lim et al. (2005)
K059	GLYAL + OH (+ O <sub>2</sub> ) GLYAL + OH (+ O <sub>2</sub> )	$\rightarrow$	GLX + 2 HO <sub>2</sub>	5.0 10 <sup>8</sup>		Lim et al. (2005)
K060	GLYAL + NO <sub>3</sub> (+ O <sub>2</sub> )	$\rightarrow$	$GLY + HO_2 + NO_3^- + H^+$	1.1 107		Herrmann et al. (2005)
K061	$\begin{array}{l} GLYAL + NO_3 \left( + O_2 \right) \\ GLYAL + NO_3 \left( + O_2 \right) \end{array}$	$\rightarrow$	$GLX + 2 HO_2 + NO_3 + H^+ + H_2O$	5.5 106		Herrmann et al. (2005)
K062	$GLY + OH (+ O_2)$	$\rightarrow$	$GLX + HO_2 + H_2O$	1.1 109	1516	Ervens et al. (2003)
K063	GLY + OH	$\rightarrow$	GLYOLIG <sup>c</sup>	3.0 1010		Carlton et al. (2007)
K064	GLYOLIG ° + OH	$\rightarrow$	OXL	3.0 1010		Carlton et al. (2007)
K065	GLYOLIG ° + OH	$\rightarrow$	GLX	1.0 109		Carlton et al. (2007)
K066	GLY + NO <sub>3</sub> (+ O <sub>2</sub> )	$\rightarrow$	$GLX + HO_2 + NO_3^- + H^+$	1.1 106	3368	Ervens et al. (2003)
K067	$CH_3CH_2OH + OH (+ O_2)$	$\rightarrow$	$CH_3CH(OH)_2 + HO_2$	1.9 109		Ervens et al. (2003)
K068	$CH_3CH_2OH + NO_3 (+ O_2)$		$CH_3CH(OH)_2 + HO_2 + NO_3 + H^+$	2.2 106		Ervens et al. (2003)
K069	$CH_3CH_2OH + CO_3^- (+ O_2)$ $CH_3CH_2OH + CO_3^- (+ O_2)$	$\rightarrow$	$CH_3CH(OH)_2 + HO_2 + CO_2$	1.5 10 <sup>4</sup>		Ervens et al. (2003)
K070	$CH_3COOH + OH (+ O_2)$		$0.85 \text{ GLX} + 0.15 \text{ CH}_2(\text{OH})_2 + 0.15 \text{ CO}_2 + \text{HO}_2$	1.5 107	1330	Ervens et al. (2003)
K070 K071	$CH_{3}COO^{-} + OH (+ O_{2})$	$\rightarrow$	0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 0.15 CO <sub>2</sub> + HO <sub>2</sub> 0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 0.15 CO <sub>2</sub> + HO <sub>2</sub>	1.0 10 <sup>8</sup>	1800	Ervens et al. (2003)
K071 K072	$CH_{3}COOH + NO_{3}(+ O_{2})$	$\rightarrow$	0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 0.15 CO <sub>2</sub> + HO <sub>2</sub> 0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 0.15 CO <sub>2</sub> + NO <sub>3</sub> <sup>-</sup> + H <sup>+</sup>	1.4 10 <sup>4</sup>	3800	Ervens et al. (2003) Ervens et al. (2003)
K072	$CH_{3}COO^{-} + NO_{3}(+O_{2})$	→	0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 0.15 CO <sub>2</sub> + NO <sub>3</sub> + H <sup>+</sup>	2.9 106	3800	Ervens et al. (2003)
K073 K074	$CH_{3}COO^{-} + CO_{3}^{-} (+ O_{2})$ $CH_{3}COO^{-} + CO_{3}^{-} (+ O_{2})$	$\rightarrow$	0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 0.15 CO <sub>2</sub> + NO <sub>3</sub> + H 0.85 GLX <sup>-</sup> + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 1.15 CO <sub>2</sub>	5.8 10 <sup>2</sup>	3800	Ervens et al. (2003)
K074 K075	$CH_3COO + CO_3 (+ O_2)$	$\rightarrow$	0.85 GLX + 0.15 CH <sub>2</sub> (OH) <sub>2</sub> + 1.15 CO <sub>2</sub> MGLY + HO <sub>2</sub>	1.2.10 <sup>9</sup>		Herrmann et al. (2005)
K075 K076	HYAC + OH HYAC + NO <sub>3</sub>	$\rightarrow$	$MGLY + HO_2$ $MGLY + NO_3 + H^+$	1.7 106		
K076 K077	MCLV + OU(+ O)	$\rightarrow$		1.7 10 <sup>-</sup> 1.1 10 <sup>9</sup>	1600	Herrmann et al. (2005)
K077 K078	$MGLY + OH (+ O_2)$ MGLY + OH	$\rightarrow$	0.92 PRV + 0.08 GLX + 0.08 CO <sub>2</sub> + HO <sub>2</sub> + H <sub>2</sub> O 0.8 MGLYOLIG ° + 0.2 OXL	1.1 10 <sup>9</sup>	1600	Ervens et al. (2004)
	MGLY + OH				1600	Lin et al. (2014) (Tan et al., 2012)
K079 K080	$MGLY + NO_3 (+ O_2)$	$\rightarrow$	0.92 PRV + 0.08 GLX + 0.08 CO <sub>2</sub> + HO <sub>2</sub> + NO <sub>3</sub> <sup>-</sup> + H <sup>+</sup>	6.3 10 <sup>7</sup> 1.2 10 <sup>8</sup>	2800	Herrmann et al. (2005)
K080 K081	$PRV + OH (+ O_2)$ PRV + OH	$\rightarrow$	$CH_3COOH + HO_2 + CO_2$	7.0 10 <sup>8</sup>	2800	Herrmann et al. (2005)
			$CH_3COO^+ + HO_2 + CO_2$		2285	Herrmann et al. (2005)
K082	$PRV + NO_3(+O_2 + H_2O)$	$\rightarrow$	$CH_3COOH + CO_2 + HO_2 + NO_3 + H^+$	4.8 106		Herrmann et al. (2005)
K083	$PRV + NO_3(+O_2 + H_2O)$	$\rightarrow$	$\mathrm{CH_3COO^{\text{-}}+CO_2+HO_2+NO_3^{\text{-}}+H^{\text{+}}}$	1.9 10 <sup>8</sup>		Herrmann et al. (2005)
K084	$GLX + OH (+ O_2)$	$\rightarrow$	$OXL + HO_2 + H_2O$	3.6 10 <sup>8</sup>	1000	Deguillaume et al. (2009)
K085	GLX + OH (+ O <sub>2</sub> )	$\rightarrow$	$OXL^{\cdot} + HO_2 + H_2O$	2.6 109	4330	Deguillaume et al. (2009)
K086	$GLX + NO_3(+O_2)$	$\rightarrow$	$OXL + HO_2 + NO_3^- + H^+$	3.0 10 <sup>6</sup>		as for glycolic acid from Herrmann et al. (2005)
K087	$GLX^{-} + NO_3(+O_2)$ $OXL^{-} + OH(+O_2)$	$\rightarrow$	$OXL^{+} + HO_2 + NO_3^{-} + H^+$	1.1 10 <sup>8</sup>		as for glycolic acid from Herrmann et al. (2005)
K088	$OXL^+ + OH (+ O_2)$ $OXL^+ + OH (+ O_2)$	$\rightarrow$	$2CO_2 + HO_2 + H_2O$	1.4 106		Ervens et al. (2004) Ervens et al. (2003)
K089			$2CO_2 + O_2 + H_2O$	3.2 107		
	$OAL + OII(+ O_2)$	$\rightarrow$	$2CO_2 + O_2 + HO_2$	5.3 10 <sup>6</sup>		
K090	OXL <sup>2-</sup> + OH (+ O <sub>2</sub> )	$\rightarrow$				Ervens et al. (2003)
K091	$OXL^{2} + OH (+ O_2)$ $OXL + NO_3 (+ O_2)$	$\rightarrow$ $\rightarrow$	$2CO_2 + NO_3 + H^+ + HO_2$	6.8 107		Ervens et al. (2003)
K091 K092	$OXL^{2} + OH (+ O_2)$ $OXL + NO_3 (+ O_2)$ $OXI - + NO_3 (+ O_2)$	$\rightarrow$ $\rightarrow$ $\rightarrow$	$2CO_2 + NO_3 + H^+ + HO_2$ $2CO_2 + NO_3 + H^+ + O_2$	6.8 10 <sup>7</sup>		Ervens et al. (2003) Ervens et al. (2003)
K091	$\begin{array}{l} OXL^{2*} + OH (+ O_2) \\ OXL + NO_3 (+ O_2) \\ OXL^* + NO_3 (+ O_2) \\ OXL^{2*} + NO_3 (+ O_2) \end{array}$	$\rightarrow$ $\rightarrow$	$2CO_2 + NO_3 + H^+ + HO_2$			Ervens et al. (2003)
K091 K092 K093	$\begin{array}{l} OXL^{2*} + OH \ (+ \ O_2) \\ OXL + NO_3 \ (+ \ O_2) \\ OXL^{-} + NO_3 \ (+ \ O_2) \\ OXL^{2*} + NO_3 \ (+ \ O_2) \\ \textbf{Iron chemistry} \end{array}$	$$ $$ $$ $$	$\begin{array}{l} 2CO_2 + NO_3^{-} + H^+ + HO_2 \\ 2CO_2 + NO_3^{-} + H^+ + O_2^{-} \\ 2CO_2 + NO_3^{-} + O_2^{-} \end{array}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup>		Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094	$\begin{array}{l} OXL^{2*} + OH \ (+ \ O_2) \\ OXL + NO_3 \ (+ \ O_2) \\ OXL^{-} NO_3 \ (+ \ O_2) \\ OXL^{2*} + NO_3 \ (+ \ O_2) \\ \hline \textbf{Iron chemistry} \\ Fe^{3+} + O_2 \end{array}$	$$ $$ $$ $$ $$ $$	$\begin{array}{l} 2CO_2 + NO_3 + H^+ + HO_2 \\ 2CO_2 + NO_3 + H^+ + O_2 \\ 2CO_2 + NO_3 + H^+ + O_2 \\ 2CO_2 + NO_3 + O_2 \end{array}$ Fe <sup>2s</sup> + O <sub>2</sub>	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup>		Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095	$\begin{array}{l} OXL^{2*} + OH (+ O_2) \\ OXL + NO_3 (+ O_2) \\ OXL^* + NO_3 (+ O_2) \\ OXL^2 + NO_3 (+ O_2) \\ \hline \textbf{Iron chemistry} \\ Fe^{3*} + O_2 \\ \hline \left[Fe(OH)]^{2*} + HO_2 \end{array}$	$\begin{array}{c} \uparrow \\ \uparrow $	$\frac{2CO_2 + NO_3 + H^2 + HO_2}{2CO_2 + NO_3 + H^2 + O_2}$ $\frac{2CO_2 + NO_3 + H^2 + O_2}{2CO_2 + NO_3 + O_2}$ $Fe^{23} + O_2$ $Fe^{23} + O_2 + HO$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.3 10 <sup>5</sup>		Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096	$\begin{array}{l} OXL^{2*} + OH (+ O_2) \\ OXL + NO_3 (+ O_2) \\ OXL^* + NO_3 (+ O_2) \\ OXL^2 + NO_3 (+ O_2) \\ \hline \textbf{Iron chemistry} \\ Fe^{3*} + O_2 \\ \hline \left[Fe(OH)]^{2*} + HO_2 \end{array}$	$\begin{array}{c} \uparrow \uparrow \uparrow \uparrow \uparrow \uparrow \\ \uparrow \uparrow \uparrow \uparrow \uparrow \uparrow \uparrow \uparrow \end{array}$	$\begin{array}{c} 2CO_{1} + NO_{7} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{7} + H^{+} + O_{7} \\ 2CO_{2} + NO_{7} + H^{+} + O_{7} \\ 2CO_{2} + NO_{7} + O_{7} \\ Fe^{2s} + O_{2} \\ Fe^{2s} + O_{2} + HO \\ Fe^{2s} + O_{2} + HO \end{array}$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.3 \ 10^5 \\ 1.5 \ 10^8 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097	$\begin{array}{l} OXL^{2-} + OH (+ O_2) \\ OXL + NO_3 (+ O_2) \\ OXL + NO_3 (+ O_2) \\ OXL^{2-} + NO_3 (+ O_2) \\ Iron chemistry \\ Fe^{2+} + O_2 \\ [Fe(OH)]^{2+} + HO_2 \\ [Fe(OH)]^{2+} + O_2 \\ Fe^{2+} + SO_2^{2-} \end{array}$	1 1 1 1 1 1 1 1	$\begin{array}{c} 2CO_2 + NO_2 + H^* + HO_2 \\ 2CO_2 + NO_3 + H^* + O_2 \\ 2CO_2 + NO_3 + H^* + O_2 \\ 2CO_2 + NO_3 + O_2 \\ Fe^{2*} + O_2 \\ Fe^{2*} + O_2 + HO \\ Fe^{2*} + O_2 + HO \\ Fe^{2*} + O_2 + HO \\ [Fe(SO_4)]^* \end{array}$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.3 \ 10^5 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Deguillamme et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098	$\begin{array}{l} & OXL^{+} + OH (+ O_2) \\ & OXL + NO_1 (+ O_2) \\ & OXL + NO_1 (+ O_2) \\ & OXL^{+} + NO_1 (+ O_2) \\ & Tron chemistry \\ Fe^{i_1} + O_i^{-} \\ & [Fe(OH)]^{i_2} + HO_2 \\ & [Fe(OH)]^{i_2} + O_2^{-} \\ & Fe^{i_1} * SO_i^{-} \\ & [Fe(SO_i)]^{i_1} \end{array}$	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	$\begin{array}{c} 2CO_{+} \times NO_{7} + H^{+} + HO_{2} \\ 2CO_{+} \times NO_{7} + H^{+} + O_{7} \\ 2CO_{+} \times NO_{7} + O_{7} \\ ECO_{+} \times NO_{7} + O_{2} \\ Fe^{2} + O_{2} + HO \end{array}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.3 10 <sup>5</sup> 1.5 10 <sup>8</sup> 3.2 10 <sup>3</sup> 2.7 10 <sup>1</sup>		Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098 K099	$\begin{array}{l} & OXL^{2*} + OH (+ O_2) \\ & OXL + NO_1 (+ O_2) \\ & OXL^{2*} + NO_1 (+ O_2) \\ & OXL^{2*} + NO_1 (+ O_2) \\ & Iron chemistry \\ Fe^{2*} + O_2 \\ & [Fe(OH)]^{2*} + HO_2 \\ & [Fe(OH)]^{2*} + O_2 \\ & Fe^{2**} + O_2^{2*} \\ & [Fe(SO_4)]^{2*} \\ & [Fe(SO_4)]^{2*} + HO_2 \end{array}$	* * * * * * * * * * * *	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + O_{2} \\ Fe^{2+} + O_{2} \\ Fe^{2+} + O_{2} + HO \\ Fe^{2+} + O_{2} + HO \\ Fe^{2+} SO_{4}^{2+} - O_{2} + H^{-} \end{array}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.5 10 <sup>8</sup> 3.2 10 <sup>3</sup> 2.7 10 <sup>1</sup> 1.0 10 <sup>5</sup>		Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100	$\begin{array}{l} & OXL^{2} + OH (+ O_2) \\ & OXL + NO_1 (+ O_2) \\ & OXL + NO_1 (+ O_2) \\ & OXL^{2} + NO_1 (+ O_2) \\ & Iron chemistry \\ & Fe^{2+} O_2^{-} \\ & [Fe(OH)]^{2+} + HO_2 \\ & [Fe(OH)]^{2+} + O_2^{-} \\ & Fe^{1+} SO_2^{-2} \\ & [Fe(SO_1)^{2+} + O_2 \\ & [Fe(SO_1)^{2+} + O_2 \\ & [Fe(SO_2)^{2+} + O_2 \\ & \end{bmatrix} \end{array}$	* * * * * * * * * * * *	$\begin{array}{c} 2CO_{+} + NO_{7} + H^{+} + HO_{2} \\ 2CO_{+} + NO_{7} + H^{+} + O_{2} \\ 2CO_{2} + NO_{7} + H^{+} + O_{7} \\ 2CO_{2} + NO_{7} + O_{7} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + SO_{4}^{-2} + HO \\ Fe^{2} + SO_{4}^{-2} + O_{7} + H^{+} \\ Fe^{2} + SO_{4}^{-2} + O_{7} + H^{-} \\ Fe^{2} + SO_{4}^{-2} + O_{7} \\ \end{array}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.3 10 <sup>5</sup> 1.5 10 <sup>8</sup> 3.2 10 <sup>3</sup> 2.7 10 <sup>1</sup> 1.0 10 <sup>5</sup> 1.5 10 <sup>8</sup>	1100	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K101	$\begin{array}{l} & \text{OXL}^{2} + \text{OH} \left( + \text{O}_{2} \right) \\ & \text{OXL} + \text{NO}_{1} \left( + \text{O}_{2} \right) \\ & \text{OXL} + \text{NO}_{1} \left( + \text{O}_{2} \right) \\ & \text{OXL}^{2} + \text{NO}_{2} \left( + \text{O}_{2} \right) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_{2} \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_{2} \\ & \text{[Fe(OH)]}^{2+} + \text{O}_{2} \\ & \text{[Fe(OSI)]}^{2+} + \text{O}_{2} \\ & \text{[Fe(SOI)]}^{2+} + \text{HO}_{2} \\ & \text{[Fe(SOI)]}^{2+} + \text{HO}_{2} \\ & \text{[Fe(SOI)]}^{2+} + \text{HO}_{2} \\ & \text{[Fe(SOI)]}^{2+} + \text{O}_{2} \end{array}$	* * * * * * * * * * * * * *	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ Pc^{2} + O_{2} + Pc^{2} + O_{2} + HcO \\ Fc^{2} + SO_{4}^{2} + O_{2} + H^{-} \\ Fc^{2} + SO^{2} + O_{2} + H^{-} \\ Fc^{2} + SO^{2} + O_{2} + H^{-} \\ Fc^{2} + SO^{2} + O_{2} \\ FcroH^{2} \\ \end{array}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.5 10 <sup>8</sup> 3.2 10 <sup>3</sup> 2.7 10 <sup>1</sup> 1.0 10 <sup>5</sup> 1.5 10 <sup>8</sup> 4.3 10 <sup>8</sup>	1100	Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K101 K102	$\begin{array}{l} & \text{OXL}^{2} + \text{OH} (+ \text{O}_{2}) \\ & \text{OXL} + \text{NO}_{1} (+ \text{O}_{2}) \\ & \text{OXL} + \text{NO}_{1} (+ \text{O}_{2}) \\ & \text{OXL}^{2} + \text{NO}_{1} (+ \text{O}_{2}) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} \text{O}_{2}^{-} \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_{2} \\ & \text{[Fe(OH)]}^{2+} + \text{O}_{2}^{-} \\ & \text{Fe}^{2+} \text{i} \text{O}_{2}^{-} \\ & \text{[Fe(SO_{2})]}^{+} + \text{HO}_{2} \\ & \text{[Fe(SO_{2})]}^{+} + \text{HO}_{2} \\ & \text{[Fe(SO_{2})]}^{+} + \text{O}_{2}^{-} \\ & \text{Fe}^{2+} \text{HO}_{2} \\ & \text{Fe}^{2+} \text{HO}_{2} (+ \text{H}^{-}) \end{array}$	* * * * * * * * * * * * * *	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ Pc^{2} + O_{2} + Pc^{2} + O_{2} + HcO \\ Fc^{2} + SO_{4}^{2} + O_{2} + H^{-} \\ Fc^{2} + SO^{2} + O_{2} + H^{-} \\ Fc^{2} + SO^{2} + O_{2} + H^{-} \\ Fc^{2} + SO^{2} + O_{2} \\ FcroH^{2} \\ \end{array}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.3 10 <sup>5</sup> 1.5 10 <sup>8</sup> 3.2 10 <sup>3</sup> 2.7 10 <sup>1</sup> 1.0 10 <sup>5</sup> 1.5 10 <sup>8</sup> 4.3 10 <sup>8</sup> 1.2 10 <sup>6</sup>	1100 5050	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K101 K102 K103	$\begin{array}{l} & \text{OXL}^{2} + \text{OH} \left( + \text{Oz} \right) \\ & \text{OXL} + \text{NO}_{1} \left( + \text{Oz} \right) \\ & \text{OXL} + \text{NO}_{1} \left( + \text{Oz} \right) \\ & \text{OXL}^{2} + \text{NO}_{1} \left( + \text{Oz} \right) \\ & \text{For charistry} \\ & \text{Fe}^{2+} + \text{Oz} \\ & \text{Fe}^{2+} + \text{Oz} \\ & \text{Fe}^{2+} + \text{Oz}^{2} \\ & \text{Fe}^{2+} + \text{Oz} \\ & \text{Fe}^{2$	* * * * * * * * * * * * * * * *	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + O_{3} \\ Fe^{2} + O_{2} + HO_{3} \\ Fe^{2} + O_{2} + HO_{4} \\ Fe^{2} + O_{2} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{2} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{2} + H^{+} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{4} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \end{array}$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ \hline 1.5 \ 10^8 \\ 1.3 \ 10^5 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.0 \ 10^5 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K095 K096 K097 K098 K099 K100 K101 K102 K103 K104	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH}  (+ \ 0_{2}) \\ {\rm OXL} + {\rm NO}_{1}  (+ \ 0_{2}) \\ {\rm OXL} + {\rm NO}_{1}  (+ \ 0_{2}) \\ {\rm OXL}^{2} + {\rm NO}_{1}  (+ \ 0_{2}) \\ {\rm Iron \ chemistry} \\ {\rm Fe}^{2+}  0_{2} \\ {\rm Fe}^{2+}  0_{2} \\ {\rm Fe}^{(+)}  {\rm SO}_{2}^{2-} \\ {\rm Fe}^{(+)}  {\rm HO}_{2} \\ {\rm HO}_{$	<u> </u>	$\begin{array}{c} 2CO_{+} \times NO_{2} + H^{+} + HO_{2} \\ 2CO_{+} \times NO_{2} + H^{+} + O_{2} \\ 2CO_{+} \times NO_{2} + H^{+} + O_{2} \\ 2CO_{+} \times NO_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ \hline 1.5 \ 10^8 \\ 1.3 \ 10^5 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.0 \ 10^5 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K101 K102 K103 K104 K105	$\begin{array}{l} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{France hemistry} \\ & \text{Fa}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{OA}_2 \\ & \text{Fe}^{2+} + \text{Fe}^{2+} \\ & \text{Fe}^{2+} + \text{OA}_2 \\ & F$	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + O_{3} \\ 2CO_{2} + NO_{3} + H^{+} O_{3} \\ Fe^{2} + O_{2} + HO_{3} \\ Fe^{2} + O_{3} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{3} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{4} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{4} + H^{+} \\ Fe^{2} + SO_{4}^{2} + O_{4} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + O_{4} + HO_{5} \\ Fe^{2} + O_{4} + O_{5} \\ Fe^{2} + O_{4} \\ Fe^{2} + O_{4} \\ Fe^{2} + O_{5} \\ Fe^{2} + O_{6} \\ Fe^{2} +$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ \hline 1.5 \ 10^8 \\ 1.3 \ 10^5 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.0 \ 10^5 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \\ 8.2 \ 10^5 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K101 K102 K103 K104 K105 K106	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH}  (+ \ 0.2) \\ {\rm OXL} + {\rm NO}_1  (+ \ 0.2) \\ {\rm OXL} + {\rm NO}_1  (+ \ 0.2) \\ {\rm OXL}^{2} + {\rm NO}_1  (+ \ 0.2) \\ \hline {\rm Iron chemistry} \\ {\rm Fe}^{2+}  0.2, \\ {\rm [Fe}({\rm OH})]^{2+}  + {\rm HO}_2 \\ {\rm [Fe}({\rm OH})]^{2+}  + {\rm HO}_2 \\ [{\rm Fe}({\rm SO}_1)^{2+}  + {\rm HO}_2 \\ [{\rm Fe}({\rm SO}_1)^{2+}  + {\rm O}_2 \\ [{\rm Fe}({\rm SO}_1)^{2+}  + {\rm O}_2 \\ {\rm [Fe}({\rm SO}_1)^{2+}  + {\rm O}_2 \\ {\rm Fe}^{2+}  +  {\rm HO}_2  + {\rm HO}_1 \\ {\rm Fe}^{2+}  +  {\rm HO}_2  + {\rm HO}_2 \\ {\rm Fe}^{2+}  +  {\rm HO}_2 \\ {\rm Fe}^{2+}  +  {\rm HO}_2 \\ {\rm Fe}^{2+}  +  {\rm HO}_2 \\ {\rm Fe}^{2+}  +  {\rm HO}_2 \end{array}$	* * * * * * * * * * * * * * * * * * * *	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + O_{3} \\ 2CO_{2} + NO_{3} + H^{+} O_{3} \\ Fe^{2} + O_{2} + HO_{3} \\ Fe^{2} + O_{3} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{3} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{4} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{4} + H^{+} \\ Fe^{2} + SO_{4}^{2} + O_{4} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + O_{4} + HO_{5} \\ Fe^{2} + O_{4} + O_{5} \\ Fe^{2} + O_{4} \\ Fe^{2} + O_{4} \\ Fe^{2} + O_{5} \\ Fe^{2} + O_{6} \\ Fe^{2} +$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.0 \ 10^5 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \\ 8.2 \ 10^5 \\ 3.1 \ 10^4 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K096 K097 K098 K097 K098 K100 K101 K102 K103 K104 K105 K106 K107	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH} (+ {\rm Oz}) \\ {\rm OXL} + {\rm NO}_1 (+ {\rm Oz}) \\ {\rm OXL} + {\rm NO}_1 (+ {\rm Oz}) \\ {\rm OXL}^{2} + {\rm NO}_1 (+ {\rm Oz}) \\ {\rm Iron \ chemistry} \\ {\rm Fe}^{2i+} ({\rm Oz}) \\ {\rm Fe}({\rm OH})^{2i} + {\rm HO}_2 \\ {\rm [Fe}({\rm OH})^{2i} + {\rm Oz} \\ {\rm [Fe}({\rm OS})^{2i} + {\rm Oz} \\ {\rm [Fe}({\rm SO}_2)^{2i} + {\rm Oz} \\ {\rm [Fe}({\rm SO}_2)^{2i} + {\rm HO}_2 \\ {\rm [Fe}({\rm SO}_1)^{2i} + {\rm HO}_2 \\ {\rm [Fe}^{2i} + {\rm OO}_1 \\ {\rm Fe}^{2i} + {\rm OO}_1 \\ {\rm Fe}^{2i} + {\rm OO}_1 \\ {\rm Fe}^{2i} + {\rm NO}_2 \\ {\rm Fe}^{2i} + {\rm NO}_2 \end{array} \end{array}$	* * * * * * * * * * * * * * * * * * * *	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{-1} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} \\ Pc^{2} + O_{2} + HcO \\ Fc^{2} + O_{2} + HcO \\ Fc^{2} + O_{2} + HcO \\ Fc^{2} + SO_{4}^{-1} + HcO \\ Fc^{2} + SO_{4}^{-2} + O_{2} + H^{-1} \\ Fc^{2} + SO_{4}^{-2} + O_{2} \\ Fc^{2} + SO_{4}^{-2} + O_{2} \\ Fc^{2} + HcO \\ Fc^{2} + HO \\ Fc^{2} + NO_{7} \\ Fc^{2} + NO_{7} \end{array}$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.0 \ 10^5 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \\ 8.2 \ 10^5 \\ 3.1 \ 10^4 \\ 8.0 \ 10^6 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Deguillamme et al. (2004) Deguillamme et al. (2004) Deguillamme et al. (2004) Ervens et al. (2003) Ervens et al. (2004)
K091 K092 K093 K095 K096 K097 K097 K097 K097 K100 K100 K100 K102 K103 K106 K105 K106 K107 K108	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH} (+\ 0.2)\\ {\rm OXL} + {\rm NO}_{1} (+\ 0.2)\\ {\rm OXL} + {\rm NO}_{1} (+\ 0.2)\\ {\rm OXL}^{2} + {\rm NO}_{1} (+\ 0.2)\\ \hline {\rm Iron chemistry}\\ {\rm Fe}^{2+} 0.2,\\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_{2}\\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_{2}\\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_{2}\\ {\rm [Fe(SO_{2})]}^{+} + {\rm HO}_{2}\\ {\rm [Fe^{2+} + {\rm HO}_{2}, {\rm HO}_{3}]\\ {\rm [Fe^{2+} + {\rm HO}_{3}]}\\ {\rm [Fe^{2+} + {\rm HO}_{3}]\\ {\rm [Fe^{2+} + {\rm HO}_{3}]\\ {\rm [Fe^{2+} + {\rm HO}_{3}]}\\ {\rm [Fe^{2+} + {\rm HO}_{3}]}\\ {\rm [Fe^{2+} + {\rm HO}_{3}]}\\ {\rm [Fe^{2+} + {\rm HO}_{3}]\\ {\rm [Fe^{2+} + {\rm HO}_{3}]}\\ {\rm [Fe^{2+} $	* * * * * * * * * * * * * * * * * * * *	$\begin{array}{c} 2CO_{+} \times NO_{+} + H^{+} + HO_{2} \\ 2CO_{+} \times NO_{+} + H^{+} + O_{2} \\ 2CO_{+} \times NO_{+} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + SO_{+}^{2} + O_{+} + H' \\ Fe^{2} + SO_{+}^{2} + O_{+} + H' \\ Fe^{2} + SO_{+}^{2} + O_{+} + H' \\ Fe^{2} + SO_{+}^{2} + O_{+} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{+} \\ Fe^{2} + NO_{-} \\ Fe^{2} + NO_{-} \\ Fe^{2} + NO_{-} \\ Fe^{2} + NO_{-} \\ Fe^{2} + O_{-}^{2} \end{array}$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \\ 8.2 \ 10^5 \\ 3.1 \ 10^4 \\ 8.0 \ 10^6 \\ 2.7 \ 10^7 \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Ervens et al. (2004) Ervens et al. (2004) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K100 K101 K102 K103 K104 K104 K107 K108 K109	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH}  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL}^{2} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm Fon \ chemistry} \\ {\rm Fe}^{2i+} {\rm O}_2 \\ {\rm Fe}({\rm OH})^{2i} + {\rm HO}_2 \\ {\rm [Fe}({\rm OH})^{2i} + {\rm O}_2 \\ {\rm Fe}^{2i+} {\rm SO}_2^{2i} \\ {\rm [Fe}({\rm SO}_3)^{2i} + {\rm HO}_2 \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\ {\rm Fe}^{2i} + {\rm OH} \\ {\rm OH} \\$	**** ************	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} \\ Pc^{2} + O_{2} + HcO \\ Fc^{2} + O_{3} + HcO \\ Fc^{2} + O_{3} + HcO \\ Fc^{2} + SO_{4}^{2} + O_{4} + HC \\ Fc^{2} + SO_{4}^{2} + O_{4} + H^{+} \\ Fc^{2} + SO_{4}^{2} + O_{4} + H^{+} \\ Fc^{2} + SO_{4}^{2} + O_{4} \\ Fc^{2} + HcO_{4} \\ Fc^{2} + HcO_{5} \\ Fc^{2} + HcO_{5} \\ Fc^{2} + NO_{7} \\ Fc^{2} + NO_{7} \\ Fc^{2} + CO_{5}^{2} \\ Fc^{2} + CO_{7}^{2} \\ Fc^{2} + CO_{7}^{2}$	6.8 10 <sup>7</sup> 2.2 10 <sup>8</sup> 1.5 10 <sup>8</sup> 1.3 10 <sup>5</sup> 1.5 10 <sup>8</sup> 3.2 10 <sup>3</sup> 2.7 10 <sup>1</sup> 1.0 10 <sup>5</sup> 1.5 10 <sup>8</sup> 4.3 10 <sup>8</sup> 1.2 10 <sup>6</sup> 1.0 10 <sup>7</sup> 5.0 10 <sup>1</sup> 8.2 10 <sup>5</sup> 3.1 10 <sup>4</sup> 8.0 10 <sup>6</sup> 2.7 10 <sup>7</sup> 8.6 10 <sup>5</sup>		Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Ervens et al. (2004) Deguillaume et al. (2004) Ervens et al. (2004) Ervens et al. (2004) Ervens et al. (2004)
K091 K092 K093 K094 K095 K096 K097 K098 K099 K100 K101 K102 K104 K105 K106 K107 K108 K109 K110	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH}  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL}^{2} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm For chemistry} \\ {\rm Fe}^{2+} {\rm O}_2 \\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_2 \\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_2 \\ {\rm [Fe(SO_1)]}^{2+} {\rm HO}_2 \\ {\rm Fe}^{2+} {\rm HO}_2 \\ {\rm HO}^{2+} {\rm HO}^{2+} \\ {\rm HO}^{2+} \\ {\rm HO}^{2+} {\rm HO}^{2+} \\ {\rm HO}^{2+} {\rm HO}^{2+} \\ {$	**** ************	$\begin{array}{c} 2CO_{+} \times NO_{2} + H^{+} + HO_{2} \\ 2CO_{+} \times NO_{2} + H^{+} + O_{2} \\ 2CO_{+} \times NO_{2} + H^{+} O_{2} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + SO_{4}^{+} + O_{2} + HO \\ Fe^{2} + SO_{4}^{+} + O_{2} + Fe^{2} + SO_{4}^{+} + O_{2} \\ Fe^{2} + SO_{4}^{+} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{1} \\ Fe^{2} + NO_{1} \\ Fe^{2} + NO_{1} \\ Fe^{2} + O_{2} \\ Fe^{2} + O_{2} \\ Fe^{2} + O_{1} \\ Fe^{2} + O_{2} \\ Fe^{2} + O_{1} \\ Fe$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.0 \ 10^5 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \\ 8.2 \ 10^5 \\ 3.1 \ 10^4 \\ 8.0 \ 10^6 \\ 2.7 \ 10^7 \\ 8.6 \ 10^5 \\ 1.0 \ 10^7 \\ 8.6 \ 10^5 \\ 1.0 \ 10^7 \\ \end{array}$		Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Ervens et al. (2004) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2005) Ervens et al. (2003)
K091 K092 K092 K094 K095 K096 K097 K100 K101 K104 K103 K104 K104 K106 K107 K108 K109 K100 K101 K101 K101 K111	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH}  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL}^{2} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm Fon chemistry} \\ {\rm Fe}^{2i+} {\rm O}_2 \\ {\rm Fe}^{2i+} {\rm SO}_2^{2i} \\ {\rm Fe}^{2i+} {\rm SO}_2^{2i} \\ {\rm Fe}^{2i+} {\rm SO}_2^{2i} \\ {\rm Fe}^{2i+} {\rm O}_2 \\ {\rm Fe}^{2i+} {\rm O}_2 \\ {\rm Fe}^{2i+} {\rm O}_2 \\ {\rm Fe}^{2i+} {\rm O}_1 \\ {\rm O}_1 \\ {\rm Fe}^{2i+} {\rm O}_1 \\ {\rm O}_1 \\ {\rm Ho}_2 \end{array} \end{array}$	**** ***********	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} O_{4} \\ Fe^{2} + O_{2} + HO_{4} \\ Fe^{2} + O_{2} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{4} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} + O_{5}$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.5 \ 10^8 \\ 3.2 \ 10^3 \\ 2.7 \ 10^1 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.5 \ 10^8 \\ 4.3 \ 10^8 \\ 1.2 \ 10^6 \\ 1.0 \ 10^7 \\ 5.0 \ 10^1 \\ 8.2 \ 10^5 \\ 3.1 \ 10^4 \\ 8.0 \ 10^6 \\ 2.7 \ 10^7 \\ 8.6 \ 10^5 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 \\ 1.0 \ 10^6 \\ 1.0 \ 10^7 $	5050	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K092 K094 K095 K096 K097 K100 K101 K102 K103 K104 K105 K104 K105 K107 K108 K107 K108 K107 K109 K110 K112	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Irron chemistry} \\ & \text{Fe}^{2i+} \text{O}_2 \\ & \text{Fe}^{2i+} \text{O}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{OO}_1 \\ & \text{Fe}^{2i+} \text{OO}_2 \\ & \text{Fe}^{2i+} \text$	****	$\begin{array}{c} 2CO_{+} \times NO_{2} + H^{+} + HO_{2} \\ 2CO_{+} \times NO_{2} + H^{+} + O_{2} \\ 2CO_{+} \times NO_{2} + H^{+} O_{2} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + SO_{4}^{+} + O_{2} + HO \\ Fe^{2} + SO_{4}^{+} + O_{2} + H^{+} \\ Fe^{2} + SO_{4}^{+} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{1} \\ Fe^{2} + NO_{1} \\ Fe^{2} + NO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + O_{1} + HO \\ Fe^{2} + O_{2} + O_{2} \\ Fe^{2} + O_{1} + HO \\ Fe^{2} + O_{2} + O_{2} \\ Fe^{2} + O_{1} + HO \\ Fe^{2} + O_{2} + O_{2} \\ Fe^{2} + O_{1} + HO \\ Fe^{2} + O_{2} + O_{2} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + HO \\ Fe^{$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^8 \\ 1.5 \ 10^8 \\ 3.1 \ 10^5 \\ 1.5 \ 10^8 \\ 2.7 \ 10^1 \\ 1.5 \ 10^8 \\ 1.2 \ 10^9 \\ 1.5 \ 10^8 \\ 1.2 \ 10^6 \\ 1.2 \ 10^6 \\ 1.2 \ 10^6 \\ 3.1 \ 10^4 \\ 8.0 \ 10^6 \\ 2.7 \ 10^7 \\ 8.6 \ 10^5 \\ 1.0 \ 10^7 \\ 2.0 \ 10^6 \\ 9.5 \ 10^3 \end{array}$	5050 2766	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Ervens et al. (2005) Ervens et al. (2006) Ervens et al. (2007) Ervens et al. (2007) Er
K091 K092 K093 K094 K095 K096 K097 K098 K097 K100 K101 K101 K103 K104 K105 K106 K107 K108 K107 K108 K109 K111 K111 K111 K113	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Irron chemistry} \\ & \text{Fe}^{2i+} \text{O}_2 \\ & \text{Fe}^{2i+} \text{O}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{SO}_2 \\ & \text{Fe}^{2i+} \text{OO}_1 \\ & \text{Fe}^{2i+} \text{OO}_2 \\ & \text{Fe}^{2i+} \text$	**** ***************	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{-1} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ Pc^{2} + O_{2} + HO \\ Pc^{2} + O_{2} + HO \\ Pc^{2} + O_{2} + HO \\ Pc^{2} + SO_{4}^{2} + O_{2} + HO \\ Pc^{2} + SO_{4}^{2} + O_{2} + HO \\ Pc^{2} + SO_{4}^{2} + O_{2} \\ Pc^{2} + SO_{4}^{2} + O_{2} \\ Pc^{2} + SO_{4}^{2} + O_{2} \\ Pc^{2} + HO_{2} \\ Pc^{2} + HO_{2} \\ Pc^{2} + HO_{2} \\ Pc^{2} + O_{3} \\ Pc^{2} + O_{4} \\ Pc^{2} + O_{5} \\$	$\begin{array}{c} 6810^{7}\\ 2.210^{8}\\ 1.510^{8}\\ 1.310^{5}\\ 3.210^{3}\\ 2.710^{3}\\ 1.510^{4}\\ 1.510^{6}\\ 1.210^{6}\\ 1.210^{6}\\ 1.210^{6}\\ 1.010^{7}\\ 8.010^{7}\\ 8.010^{7}\\ 8.010^{7}\\ 8.010^{7}\\ 2.710^{7}\\ 8.010^{7}\\ 2.010^{6}\\ 2.010^{6}\\ 2.3410^{4}\\ \end{array}$	5050 2766 4089	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K095 K097 K098 K097 K100 K102 K102 K103 K104 K105 K106 K107 K108 K109 K110 K111 K112 K113 K114	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(SO_1)]}^{2+} + \text{HO}_2 \\ & \text{Fe}^{2+} + HO$	****	$\begin{array}{c} \frac{1}{2}CO_{1}+NO_{2}+H^{+}+HO_{2}\\ 2CO_{2}+NO_{3}+H^{+}+O_{2}\\ 2CO_{2}+NO_{3}+H^{+}O_{2}\\ 2CO_{2}+NO_{3}+O_{2}\\ Fe^{2}+O_{3}+HO\\ Fe^{2}+O_{3}+HO\\ Fe^{2}+SO_{4}^{-2}\\ Fe^{2}+SO_{4}^{-2}\\ Fe^{2}+SO_{4}^{-2}\\ Fe^{2}+SO_{4}^{-2}+O_{2}\\ [Fe(OT)]^{D}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{2}\\ Fe^{2}+HO_{4}\\ Fe^{2}+H$	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 1.310^5\\ 3.210^7\\ 2.710^7\\ 1.010^7\\ 1.510^8\\ 4.310^8\\ 4.310^8\\ 4.310^8\\ 1.010^7\\ 5.010^7\\ 8.610^9\\ 2.710^7\\ 8.610^9\\ 2.710^7\\ 8.610^9\\ 2.710^7\\ 8.610^9\\ 2.710^7\\ 1.010^7\\ 2.010^7\\ 2.310^4\\ 1.010^7\\ 2.310^4\\ 1.010^7\\ 2.310^6\\ 1.010^7\\ 2.310^6\\ 1.010^7\\ 2.310^6\\ 1.010^7\\ 2.310^6\\ 1.010^7\\ 2.310^6\\ 1.010^7\\ 2.310^6\\ 1.010^7\\ $	5050 2766 4089 5052	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097 K098 K100 K100 K102 K104 K104 K104 K106 K106 K107 K108 K106 K107 K110 K111 K111 K111 K113 K115	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(OA)]}^{2+} + \text{O}_2 \\ & \text{[Fe(SOA)]}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{O}_1 \\ & \text{Fe}^{2+} + \text{O}_1 \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{NO}_3 \\ & \text{Fe}^{2+} + \text{NO}_3 \\ & \text{Fe}^{2+} + \text{NO}_3 \\ & \text{Fe}^{2+} + \text{NO}_2 \\ & \text{Fe}^{2+} + \text{NO}_3 \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{O}_3 \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{O}_3 \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{O}_3 \\ & \text{Fe}^{2+}$	**** ***************	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{-1} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + SO_{4}^{2} + O_{4} + HO \\ Fe^{2} + SO_{4}^{2} + O_{4} + H^{-1} \\ Fe^{2} + SO_{4}^{2} + O_{4} \\ Fe^{2} + SO_{4}^{2} + O_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + HO_{5} \\ Fe^{2} + NO_{7} \\ Fe^{2} + NO_{7} \\ Fe^{2} + NO_{7} \\ Fe^{2} + O_{7} \\ Fe^{2} + HO_{7} \\ Fe^{2} + HO_{7} \\ Fe^{2} + O_{7} \\ Fe^{2} + O_{7}$	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 1.510^8\\ 3.210^9\\ 2.710^1\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 3.110^7\\ $	5050 2766 4089	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K097 K098 K099 K100 K102 K102 K104 K105 K104 K105 K104 K105 K107 K108 K109 K110 K111 K112 K113 K114 K115 K116	$\begin{array}{l} {\rm OXL}^{2} + {\rm OH}  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm OXL}^{2} + {\rm NO}_1  (+ {\rm O}_2) \\ {\rm For chemistry} \\ {\rm Fe}^{2+} + {\rm O}_2 \\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_2 \\ {\rm [Fe(OH)]}^{2+} + {\rm HO}_2 \\ {\rm [Fe(SO_1]}^{2+} + {\rm HO}_2 \\ {\rm [Fe(SO_1]}^{2+} + {\rm O}_2 \\ {\rm [Fe(SO_1]}^{2+} + {\rm O}_2 \\ {\rm [Fe(SO_1]}^{2+} + {\rm O}_2 \\ {\rm Fe}^{2+} + {\rm IO}_2 \\ {\rm Fe}^{2+} + {\rm OO}_2 \\ {\rm Fe}^{2+} + {\rm IO}_2 \\ {\rm IO}_2 + {\rm IO}_2 \\ {\rm IO$	****	$\begin{array}{c} 1CO_{1} + NO_{2} + H^{-} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-} \\ Pc^{2} + O_{2} + HO \\ Fc^{2} + O_{3} + HO \\ Fc^{2} + SO_{4}^{-} \\ Fc^{2} + HO_{2} \\ Fc^{2} + NO_{4} \\ Fc^{2} + NO_{4} \\ Fc^{2} + NO_{4} \\ Fc^{2} + O_{4} \\ Fc^$	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 1.310^5\\ 1.510^5\\ 3.210^1\\ 2.710^7\\ 1.010^7\\ 1.010^7\\ 3.110^9\\ 4.310^8\\ 4.310^9\\ 4.310^9\\ 8.210^7\\ 8.010^6\\ 2.710^7\\ 8.610^5\\ 1.010^7\\ 2.010^5\\ 9.510^5\\ 2.3410^4\\ 1.110^6\\ 1.110^6\\ \end{array}$	2766 4089 5052 4150	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097 K098 K100 K101 K102 K103 K104 K104 K106 K107 K108 K106 K107 K110 K111 K111 K111 K115 K117	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_{2}) \\ & \text{OXL} + \text{NO}_{1} (+ \text{O}_{2}) \\ & \text{OXL} + \text{NO}_{1} (+ \text{O}_{2}) \\ & \text{OXL}^{2} + \text{NO}_{1} (+ \text{O}_{2}) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_{2} \\ & \text{Fe}^{2+} + \text{O}_{3} \\ $	**** *******************	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{-1} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} \\ Fe^{2} + O_{2} + HO_{3} \\ Fe^{2} + O_{2} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{2} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{2} + H^{-1} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{3} \\ Fe^{2} + HO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} \\ Fe^{2} + NO_{4} \\ Fe^{2} \\ Fe^{2} \\ Fe^{2} \\ Fe$	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 1.510^8\\ 3.210^9\\ 2.710^1\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.210^9\\ 1.510^9\\ 1.210^7\\ 1.010^7\\ 1.010^7\\ 3.110^6\\ 8.210^7\\ 3.110^9\\ 8.210^7\\ 1.010^9\\ 2.710^9\\ 2.710^9\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.010^5\\ 1.010^9\\ 1.01010^9\\ 1.01010^9\\ 1.01010^9\\ 1.01010^9\\ 1.01010^9\\ 1.01010^9\\ 1.010101010101010101010101010$	2766 4089 5052 4150 5352	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K093 K093 K094 K096 K096 K097 K098 K097 K100 K101 K103 K104 K105 K106 K107 K108 K107 K108 K107 K108 K107 K111 K112 K114 K117 K118	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Irron chemistry} \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(SO_1]}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{IO}_2 \\ & \text{Fe}^{2+} + \text{OO}_2 \\ & \text{Fe}^{2+} + \text{IO}_2 \\ & \text{FO}^{2+} + \text{IO}_2 \\ & \text{FO}^{2+} + \text{IO}_2 \\ & \text{FO}^{2+} + IO$	**** ********************	$\begin{array}{c} 1CO_{1} + NO_{2} + H + HO_{2} \\ 2CO_{2} + NO_{3} + H + HO_{2} \\ 2CO_{2} + NO_{3} + H + O_{3} \\ 2CO_{2} + NO_{3} + O_{2} \\ Fe^{2} + O_{3} \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + SO_{4}^{2} + HO \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + O_{4} \\ FO \\ FO \\ Fe^{2} + O_{4} \\ FO \\ F$	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 1.510^8\\ 1.510^9\\ 2.710^1\\ 2.710^1\\ 2.710^1\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 3.110^4\\ 3.110^4\\ 8.010^9\\ 2.710^7\\ 2.710^7\\ 2.710^7\\ 1.010^7\\ 2.710^7\\ 2.310^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.610^7\\ \end{array}$	2766 4089 5052 4150	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K092 K093 K094 K095 K096 K097 K098 K100 K101 K102 K103 K104 K104 K105 K106 K107 K108 K100 K110 K111 K111 K1113 K115 K116 K117 K118	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} +$	**** **********************	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{+} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} \\ 2CO_{2} + NO_{3} + H^{-} \\ Fe^{2} + O_{2} + HO_{3} \\ Fe^{2} + O_{2} + HO_{4} \\ Fe^{2} + SO_{4}^{2} + O_{2} + H^{+} \\ Fe^{2} + SO_{4}^{2} + O_{2} + H^{-} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{7} \\ Fe^{2} + NO_{7} \\ Fe^{2} + O_{7} \\ Fe^{2} + HO_{7} \\ Fe^{2} + O_{7} \\ Fe^{2} \\ Fe^{2} + O_{7} \\ Fe^{2} + O_{7} \\ Fe^{2} \\ Fe^{2} + O_{7} \\ Fe^{2} \\ Fe^{2} + O_{7} \\ Fe^{2} + O_{7} \\ Fe^{2} \\ Fe^{2} + O_{7} \\ Fe^{2} \\ Fe^{2} + O_{7} \\ Fe^{2} $	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 1.510^8\\ 3.210^9\\ 2.710^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 1.010^7\\ 3.110^7\\ 3.110^9\\ 3.110^9\\ 2.710^9\\ 3.110^9\\ 3.110^9\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^6\\ 1.010^5\\ 3.010^9\\ $	2766 4089 5052 4150 5352	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K093 K094 K096 K096 K097 K098 K099 K100 K101 K103 K104 K105 K106 K107 K108 K107 K108 K107 K108 K109 K110 K111 K112 K114 K117 K118 K118 K117 K120	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Fe}^{2}^{+} + \text{O}_2^{-} \\ & [\text{Fe}(\text{OH})]^{2} + \text{HO}_2 \\ & [\text{Fe}(\text{OH})]^{2} + \text{HO}_2 \\ & [\text{Fe}(\text{OH})]^{2} + \text{HO}_2 \\ & [\text{Fe}(\text{SO}_2)]^{+} \\ & [\text{Fe}^{2}^{+} + \text{HO}_2 \\ & [\text{Fe}^{+} + \text{HO}_2$	**** ************************	$\begin{array}{c} 1CO_{1} + NO_{2} + H + HO_{2} \\ 2CO_{2} + NO_{3} + H + HO_{2} \\ 2CO_{2} + NO_{3} + H + O_{3} \\ 2CO_{2} + NO_{3} + O_{2} \\ Fe^{2} + O_{3} \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + SO_{4}^{2} + HO \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + O_{4} \\ FO \\ Fe^{2} \\ FO \\ F$	$\begin{array}{c} 6810^7\\ 2.210^9\\ 1.510^9\\ 1.510^9\\ 3.210^9\\ 2.710^1\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.510^9\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 2.710^7\\ 2.310^4\\ 8.010^7\\ 2.710^7\\ 2.310^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^6\\ $	2766 4089 5052 4150 5352	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K093 K094 K094 K096 K096 K097 K098 K099 K100 K101 K102 K103 K104 K105 K106 K107 K108 K109 K110 K111 K111 K111 K1115 K116 K117 K118 K117 K118 K119 K121	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} +$	****	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{-1} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + SO_{4}^{2} + O_{2} + H^{-1} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{3} \\ Fe^{2} + O_{4} \\ Fe^{2} + O_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} + NO_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} + NO_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} $	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 3.210^8\\ 3.210^9\\ 1.510^8\\ 3.210^9\\ 1.510^8\\ 1.210^9\\ 1.510^8\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 3.110^7\\ 5.010^7\\ 5.010^9\\ 2.710^9\\ 3.110^9\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^6\\ 1.110^6\\ 3.010^5\\ 3.010^3\\ 3.010^3\\ \end{array}$	2766 4089 5052 4150 5352	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K093 K093 K094 K095 K096 K097 K100 K101 K102 K102 K104 K104 K104 K104 K104 K106 K107 K108 K100 K110 K111 K111 K111 K1115 K1116 K117 K118 K1120 K122	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(OH)]}^{2+} + \text{HO}_2 \\ & \text{[Fe(SO_1]}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{IO}_2 \\ & \text{Fe}^{2+} + \text{O}_2 $	**** ************************	$\begin{array}{c} CO_{1} + NO_{2} + H^{-} + HO_{2} \\ CO_{2} + NO_{3} + H^{-} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ 2CO_{2} + NO_{3} + H^{+} + O_{2} \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + SO_{4}^{-} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{3} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + NO_{4} \\ Fe^{2} + O_{4} \\ FO_{4} \\ Fe^{2} + O_{4} \\ Fe^{2} + Fe^{2} \\ Fe^{2} + O_{4} \\ Fe^$	$\begin{array}{c} 6.8 \ 10^7 \\ 2.2 \ 10^8 \\ 1.5 \ 10^3 \\ 1.5 \ 10^4 \\ 1.5 \ 10^1 $	2766 4089 5052 4150 5352	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003)
K091 K093 K094 K094 K096 K096 K097 K098 K099 K100 K101 K102 K103 K104 K105 K106 K107 K108 K109 K110 K111 K111 K111 K1115 K116 K117 K118 K117 K118 K119 K121	$\begin{split} & \text{OXL}^{2} + \text{OH} (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL} + \text{NO}_1 (+ \text{O}_2) \\ & \text{OXL}^{2} + \text{NO}_1 (+ \text{O}_2) \\ & \text{Iron chemistry} \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{HO}_2 \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} + \text{O}_3 \\ & \text{Fe}^{2+} + \text{O}_2 \\ & \text{Fe}^{2+} $	****	$\begin{array}{c} 2CO_{1} + NO_{2} + H^{-1} + HO_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} + O_{2} \\ 2CO_{2} + NO_{3} + H^{-1} \\ Fe^{2} + O_{2} + HO \\ Fe^{2} + O_{3} + HO \\ Fe^{2} + SO_{4}^{2} + O_{2} + H^{-1} \\ Fe^{2} + SO_{4}^{2} + O_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + HO_{2} \\ Fe^{2} + NO_{3} \\ Fe^{2} + O_{4} \\ Fe^{2} + O_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} + NO_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} + NO_{5} \\ Fe^{2} + O_{5} \\ Fe^{2} $	$\begin{array}{c} 6810^7\\ 2.210^8\\ 1.510^8\\ 3.210^8\\ 3.210^9\\ 1.510^8\\ 3.210^9\\ 1.510^8\\ 1.210^9\\ 1.510^8\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 1.210^9\\ 3.110^7\\ 5.010^7\\ 5.010^9\\ 2.710^9\\ 3.110^9\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^4\\ 1.110^6\\ 1.110^6\\ 3.010^5\\ 3.010^3\\ 3.010^3\\ \end{array}$	2766 4089 5052 4150 5352	Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Ervens et al. (2003) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2004) Deguillaume et al. (2003) Ervens et al. (2003)

K124	[Fe(OXL)]++ HO <sub>2</sub>	$\rightarrow$	$Fe(OXL) + O_2 + H^+$	1.2 105	Sedlak and Hoigné (1993)	-
K125	$[Fe(OXL)]^+ + O_2^-$		$Fe(OXL) + O_2$	1.0 106	Sedlak and Hoigné (1993)	
K125	$Fe(OXL) + H_2O_2$		$[Fe(OXL)]^+ + OH + HO$	3.1 104	Sedlak and Hoigné (1993)	
K120 K127	$Fe^{2+} + OXL^{2-}$		Fe(OXL) + OH + HO	3.67 105	Sedlak and Hoigné (1993)	
K127 K128	Fe(OXL)		$Fe^{2+} + OXL^{2-}$	3.0 10-3	as for K121 and K123	
K128	Fe(OAL)	$\rightarrow$	Fer + OAL-	3.0 10 -	as for K121 and K125	

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using the calculated gas-phase photolysis frequencies. aqueous-phase photolysis rates at noontime scaled on model's H<sub>2</sub>O<sub>2</sub> gas-phase photolysis frequencies. representing large multifunctional compounds (see Carlton et al., 2007). c)

### Table S3: Henry's law solubility constants (H), mass accommodation coefficients (a) and gas phase diffusion coefficients (Dg) used in aqueous-phase chemistry scheme.

Trace gas	H (mol m <sup>-3</sup> Pa <sup>-1</sup> )	-ΔH R <sup>-1</sup> (K)	Reference	α	Reference	$D_{g}(m^{2} s^{-1})$	Reference
O <sub>3</sub>	1.0 10-4	2800	1	0.05	2	1.48 10-5	2
$H_2O_2$	9.1 10 <sup>2</sup>	6600	1	0.11	2	1.46 10-5	2
HO <sub>2</sub>	6.8		1	0.01	2	1.04 10-5	2 2 2
OH	3.8 10-1		1	0.05	2	1.53 10-5	2
NO	1.9 10-5	1600	1	as for NO <sub>2</sub>	2	as for NO <sub>2</sub>	
NO <sub>2</sub>	9.9 10-5		1	0.0015	2	1.92 10-5	2
NO <sub>3</sub>	3.8 10-4		1	0.004	2	1.00 10-5	
HONO	4.8 10-1	4800	1	0.5	2	1.30 10-5	2
HNO <sub>3</sub>	8.8 10 <sup>2</sup>		1	0.054	2	1.32 10-5	2
SO <sub>2</sub>	1.3 10-2	2900	1	0.035	2	1.28 10-5	2
CO <sub>2</sub>	3.3 10-4	2400	1	0.0002	2	1.55 10-5	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2
CH <sub>3</sub> O <sub>2</sub>	1.5 10-1	3700	1	as for CH <sub>3</sub> O <sub>2</sub> H	2	as for CH <sub>3</sub> O <sub>2</sub> H	2
CH <sub>3</sub> O <sub>2</sub> H	2.9	5200	1	0.0038	2	1.31 10.5	2
HCHO	3.2 10-1	6800	1	0.02	2	1.64 10 <sup>-5</sup>	2
CH3OH	2.0	5600	1	0.015	2 2	1.16 10 <sup>-5</sup>	2
HCOOH	8.8 10 <sup>1</sup>	6100	1	0.012	2	1.53 10-5	2
GLYAL	4.1 10 <sup>2</sup>	4600	1	as for GLY		as for GLY	
GLY	4.1 10 <sup>3</sup>	7500	1	0.023	3	1.15 10-5	3
MGLY	3.4 10 <sup>1</sup>	7500	1	as for GLY		as for GLY	
HYAC	7.7 10 <sup>1</sup>		1	0.0176	4	9.50 10.7	4
CH3CHO	1.3 10-1	5900	1	0.03	2	1.22 10-5	2
CH <sub>3</sub> CH <sub>2</sub> OH	1.9	6400	1	0.0082	2	9.50 10-6	2 2
CH <sub>3</sub> COOH	4.0 10 <sup>1</sup>	6200	1	0.019	3	1.24 10-5	2
PRV	3.1 10 <sup>3</sup>	5100	1	as for CH <sub>3</sub> COOH		as for CH3COOH	
GLX	1.1 10 <sup>2</sup>	4800	1	as for CH <sub>3</sub> COOH		as for CH <sub>3</sub> COOH	
H <sub>2</sub> OXL	3.1 10 <sup>4</sup>	7300	1	as for CH3COOH		as for CH3COOH	

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# Sander (2015) and references therein. Herrmann et al. (2000) and references therein. Lim et al. (2005) and references therein. Ervens et al. (2003) and references therein.

### Table S4. Dissolution scheme for iron-containing combustion and mineral dust aerosols.

	Iron pool	Scheme g	K298 <sup>a</sup>	m <sup>b</sup>	Keq	n <sup>d</sup>	Reference
D01	Combustion Fe <sup>e</sup>	$H^+$	5.24 10-8	0.36			Ito (2015)
D02		OXL	3.85 10-6	1			Ito (2015)
D03		hv≤	4.10 10-6	1			Ito (2015)
D04	Ferrihydrite	$H^+$	7.13 10-5	1.1	1550	3	Ito and Shi (2016)
D05	-	OXL	4.61 10 <sup>-8</sup>	0.069	1550	3	Ito and Shi (2016)
D06		$hv^{f}$	4.61 10 <sup>-8</sup>	0.069			Ito and Shi (2016)
D07	Nano-Fe oxides	$H^+$	1.43 10 <sup>-4</sup>	1.6	42	2.75	Ito and Shi (2016)
D08		OXL	1.28 10-8	0.069	1550	3	Ito and Shi (2016)
D09		$hv^{f}$	1.28 10-8	0.069			Ito and Shi (2016)
D10	Aluminosilicates	$H^+$	5.85 10-8	0.76	3.3	2.85	Ito and Shi (2016)
D11		OXL	1.68 10-9	0.056	1500	3	Ito and Shi (2016)
D12		hv	$1.68 \ 10^{-9}$	0.056			Ito and Shi (2016)

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b)
 c)
 d)
 e)
 f)
 g)

 $\begin{array}{c} & \text{OXL} & 1.68\ 10^{\circ} & 0.036 & 1.000 & 5 & 100\ \text{Ito and Shi} (2016) \\ \hline \text{hv}^2 & 1.68\ 10^{\circ} & 0.056\ \text{resseld in mol Fe}\ g^{-1}\ s^{-1}\ \text{and defined as:} \\ & \mathbb{K} = K_{228} \cdot exp\left[E_{pn} \cdot (1/298 - 1/T)\right] \\ & \text{where, } \mathbb{E}_{pl1} = -1.56\ 10^{\circ}\ \text{pH} + 1.08\ 10^{\circ}\ (Bibi et al., 2014;\ Ito and Shi, 2016) \\ & \text{m is the reaction order (Ito, 2015;\ Ito and Shi, 2016) \\ & \text{K} = K_{228} \cdot exp\left[E_{pn} \cdot (1/298 - 1/T)\right] \\ & \text{where, } \mathbb{E}_{pl1} = -1.56\ 10^{\circ}\ \text{pH} + 1.08\ 10^{\circ}\ (Bioneville et al., 2004;\ Ito and Shi, 2016) \\ & \text{K} = K_{228} \cdot exp\left[E_{pn} \cdot (1/298 - 1/T)\right] \\ & \text{K} = \text{K}_{228} \cdot exp\left[E_{pn} \cdot (1/298 - 1/T)\right] \\ & \text{The unit for combustion acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for combustion acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for combustion acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for combustion acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for combustion acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for combustion acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for combustion acrosol acrosol is converted from moles } \mathbb{R}^{\circ} \cdot 1 \text{ to and Shi}, 2016) \\ & \text{The unit for call-specific advalue-promoted and solution becomes superstaticted with respect to Fe(III), but any suppression due to OXL is neglected (i.e., 0 \leq f \leq 1 \text{ and } g = 1). For combustion acrosols, both the OXL-promoted and photo-induced dissolution rates are also considered to be suppressed by the formation of amorphous Fe(OH)_{j}(s) (Ito, 2015). \\ & \text{The original strange is a superssed by the formation of amorphous Fe(OH)_{j}(s) (Ito, 2015). \\ & \text{The original strange is a superssele by the formation of amorphous Fe(OH)_{j}(s) (Ito, 2015). \\ & \text{The original strange is a superssele by the formation of amorphous Fe(OH)_{j}(s) (Ito, 2015). \\ & \text{The original str$ 

Table S5: Aqueous phase equilibrium constants  $\left(E_k\right)$  used in aqueous-phase chemistry scheme.

Equilibrium Reaction			Ek298 (mol L-1)	-ΔH R <sup>-1</sup> (K)	Reference
H <sub>2</sub> O	$\leftrightarrow$	$HO^{-} + H^{+}$	1.0 10-14	-6710	1
H <sub>2</sub> O <sub>2</sub>	$\leftrightarrow$	$HO_{2}^{-} + H^{+}$	2.2 10-12	-3730	1
HO <sub>2</sub>	$\leftrightarrow$	$O_{2^{*}} + H^{+}$	3.5 10-5		1
CO <sub>2</sub> (+ H <sub>2</sub> O)	$\leftrightarrow$	H <sub>2</sub> CO <sub>4</sub>	7.7 10-7	-750	2
H <sub>2</sub> CO <sub>4</sub>	$\leftrightarrow$	$HCO_{3}^{-} + H^{+}$	2.0 10-4		2
HCO3 <sup>-</sup>	$\leftrightarrow$	$CO_3^{2*} + H^+$	4.69 10-11	-1820	2
NH4OH	$\leftrightarrow$	NH4 <sup>+</sup> + HO <sup>-</sup>	1.7 10-5	-450	1
SO <sub>2</sub> ·H <sub>2</sub> O	$\leftrightarrow$	$HSO_3^- + H^+$	1.3 10-2	1960	1
HSO3 <sup>-</sup>	$\leftrightarrow$	$SO_4^{2*} + H^+$	6.6 10-8	1500	1
HONO	$\leftrightarrow$	$NO_{2} + H^{+}$	5.1 10-4	-1260	1
HNO <sub>3</sub>	$\leftrightarrow$	$NO_{3}^{-} + H^{+}$	15.4	8700	1
HCOOH	$\leftrightarrow$	$HCOO^{-} + H^{+}$	1.77 10-4	-12	3
CH <sub>3</sub> COOH	$\leftrightarrow$	$CH_3COO^{-} + H^{+}$	1.75 10-5	-46	3
PRV	$\leftrightarrow$	$PRV^+ + H^+$	3.2 10-3		4
GLX	$\leftrightarrow$	$GLX^{*} + H^{+}$	3.47 10-4	-267	4
H <sub>2</sub> OXL	$\leftrightarrow$	$HOXL^{-} + H^{+}$	5.6 10-2	-453	4
HOXL	$\leftrightarrow$	$OXL^- + H^+$	5.42 10-5	-805	4
Fe <sup>3+</sup> (+ H <sub>2</sub> O)	$\leftrightarrow$	[Fe(OH)] <sup>2+</sup> + H <sup>+</sup>	1.1 10-4		3
[Fe(OH)] <sup>2+</sup> (+ H <sub>2</sub> O)	$\leftrightarrow$	$[Fe(OH)_2]^+ + H^+$	1.4 10-7		3

Seinfeld and Pandis (2006) and references there Herrmann et al. (2000) and references therein. Ervens et al. (2003) and references therein. Lim et al. (2005) and references therein.

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# Supplementary Equations

Mathematical formulas for correlation coefficient (R; Eq. S1), normalized mean bias (nMB; Eq. S2), and the normalized root mean square error (nRMSE; Eq. S3), used for the statistical analysis of model comparison against observations;  $O_i$  and  $P_i$  stand for observations and predictions, respectively. N is the number of pairs (observations, predictions) that are compared.

$$R = \begin{bmatrix} \frac{1}{N} \sum_{i=1}^{N} (O_i - \overline{O})(P_i - \overline{P}) \\ \sigma_o \sigma_P \end{bmatrix}$$
(Eq. S1)  
$$NMB = \frac{\sum_{i=1}^{N} (M_i - O_i)}{\sum_{i=1}^{N} O_i} \times 100$$
(Eq. S2)  
$$nRMSE = \sqrt{\frac{\prod_{i=1}^{N} \sum_{i=1}^{N} (P_i - O_i)^2}{\sum_{i=1}^{N} O_i}}$$
(Eq. S3)

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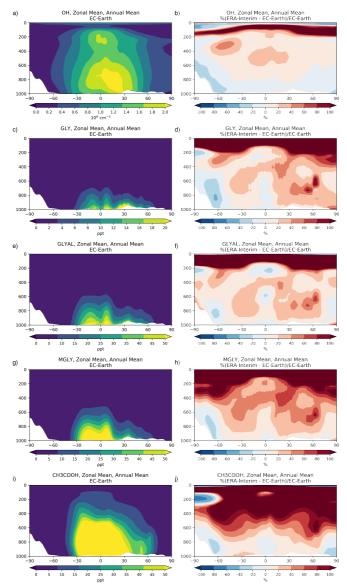


Figure S1: Annual mean zonal mean concentrations of a) OH radicals ( $10^6$  molec. cm<sup>-3</sup>), c) glyoxal (ppt), e) glycolaldehyde (ppt), g) methylglyoxal (ppt), and i) acetic acid (ppt), as simulated for the EC-Earth simulation, averaged for the period 2000-2014, and the absolute differences to the ERA-Interim simulation (b,d,f,h,j).

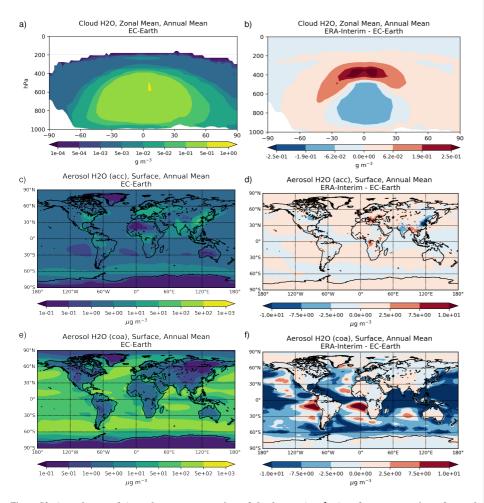


Figure S2: Annual mean of a) zonal mean concentrations of cloud water (g m<sup>-3</sup>), c) surface concentrations of aerosol water associated with accumulation aerosols ( $\mu$ g m<sup>-3</sup>), and e) surface concentrations of aerosol water associated with coarse aerosols ( $\mu$ g m<sup>-3</sup>), as simulated for the EC-Earth simulation, averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (b,d,f).

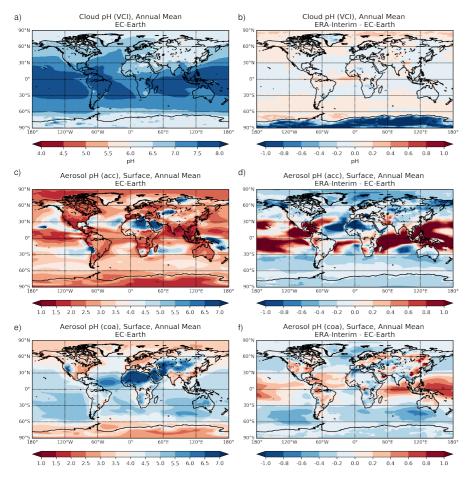


Figure S3: Annual mean a) vertical-column-integrated (VCI) liquid-water-weighted cloud water pH, c) accumulation aerosol pH at the surface, e) and coarse aerosol pH at the surface, as calculated for the EC-Earth simulation, averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (b,d,f).

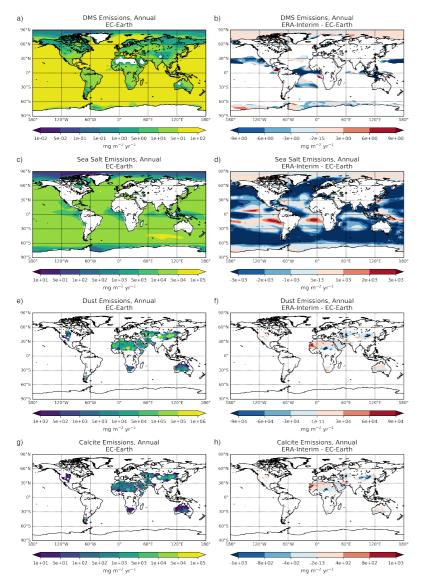


Figure S4: Annual mean emission fluxes (mg m<sup>-2</sup> yr<sup>-1</sup>) of a) DMS, c) sea salt, e) mineral dust, and g) calcite (in mg Ca m<sup>-2</sup> yr<sup>-1</sup>), as calculated for the EC-Earth simulation, averaged for the period 2000-2014, and the respective absolute differences to the ERA-Interim simulation (b,d,f,h).

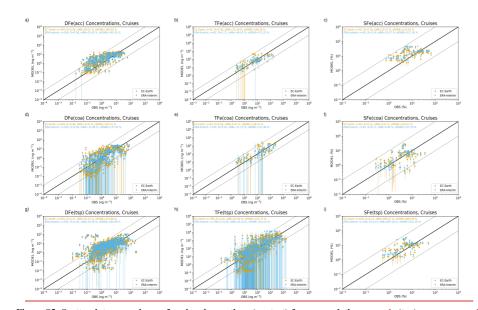


Figure S5: Scatterplot comparisons of cruise observations (see text) for accumulation <u>aerosols</u> (top), coarse <u>aerosols</u> (middle), and total suspended matter (bottom) of dissolved iron (DFe; adde), total iron (TFe; bdel) aerosols (ng m<sup>3</sup>) and the derived aerosol solubility (SFe=%DFe/TFe; c,f,i) of EC-Earth (<u>orange</u> circles) and ERA-Interim (<u>light blue</u> triangles) simulations; the solid line represents the 1 : 1 correspondence and the dashed lines show the 10 : 1 and 1 : 10 relationships, respectively. The <u>color-coded</u> error bars represent the <u>model's</u> standard error of the multi-annual mean for the individual observational period. Summary statistics (<u>color-coded</u> for all points are also included.

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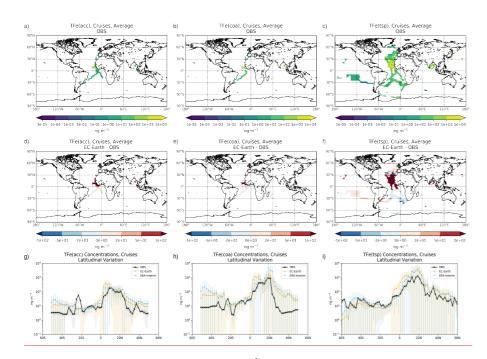


Figure S6: Observed total iron (TFe) concentrations (ng m<sup>-3</sup>) of a) accumulation aerosols, b) coarse aerosols, and c) total suspended particles (tsp), the respective absolute differences to the ERA-Interim simulation (d.e.f.), and the comparison to observations (black z-line) in latitudinal order (g.e.f.) with the EC-Earth (<u>orange circle-line</u>) and ERA-Interim <u>dight blue</u> triangle-line) simulations; the grey shaded areas correspond to the standard deviation of the observations and the <u>color-coded error bars/shaded areas</u> correspond to the <u>model's</u> standard error of the multi-annual mean for the individual observational period.

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