# Interactive comment on "Analytical solutions for mantle flow in cylindrical and spherical shells" by Stephan C. Kramer et al. 

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Thank you for your very thorough, careful and constructive review of our study. We are pleased to know you enjoyed reading the paper. In response to your main point of criticism, we agree that the paper should include some more (recent) references and a discussion of other work containing analytical Stokes solutions. Most relevant here is HMB20, your very nice paper with Andre Horbach and Hans-Peter Bunge, which came out last year and we had not yet seen. The two papers (i.e. ours and yours) use similar techniques to derive non-trivial free slip and no-slip analytical Stokes solutions in spherical shells. In our paper, we derive the analytical Stokes solution for a simple (radially monomial) forcing combined with spherical harmonics. This is achieved by decompos-
poloidal part. An inhomogeneous solution is combined with four solutions to the homogeneous equation, the coefficients of which are determined by imposing boundary conditions. In contrast, HMB20 describes a procedure starting with the selection of four arbitrarily chosen, linearly independent, solutions for the poloidal function, after which a linear combination of these is determined through the boundary conditions. Finally the necessary right-hand-side term is then constructed by substitution of the corresponding velocity field into the Stokes equations. In a sense, this approach is more akin to the method of manufactured solutions (MMS), but overcomes its usual problem of not satisfying desired boundary conditions. Both approaches are fairly flexible in terms of selecting solutions with desired physical properties, e.g. selecting high order polynomial to obtain strong gradients near the top boundary. In addition, our paper provides similar solutions in cylindrical shell domains, and a different set of solutions, both in 2D and 3D, based on an infinitely thin density anomaly. For the latter, for which the spherical solutions have been used previously as benchmarks [Zhong'08 and others], we discuss a particular issue with the continuous Galerkin finite element method. We believe the two papers are therefore highly complementary.
In the revised version of the manuscript we have given more attention to this and other references you mention, by referring to them in the introduction and by adding a discussion at the end of the manuscript that compares ours with previously published analytical spherical Stokes solutions. In our response to Cedric Thieulot we already indicated that we believe the analytical solution in Thi17 (isoviscous case) does not satisfy no-slip or free slip boundary conditions and therefore does not overlap with ours. The same appears to be true for the solutions in BMP16. PLP+14, which indeed also deserves a mention, deals with Cartesian domains only.

## Specific Comments

- Equation (65) for computing the relative error: You are working with two different

domains, $\Omega$ being the physical problem domain, and $\Omega_{h}$ the computational domain for the finite element method. I assume that the computation of the error happens w.r.t. $\Omega_{h}$. In the isoparametric approach that does not make a significant difference, but for the affine mapping approach, Figure 5, I guess it would be better to make that distinction explicit.
As highlighted, we did not make a proper distinction between the (idealised) physical domain and the numerical domain. This has been corrected by introducing the notation $\Omega$ and $\Omega_{h}$ respectively, as suggested.
- You treat the case of both boundaries being either zero-slip or free-slip. However, aren't many simulations run as a 'mixed case', i.e. with Dirichlet boundary conditions for velocity (from plate reconstructions) on top and free-slip conditions at the CMB? In that sense, would it make sense to add (maybe in an appendix) also the coefficients for such a mixed case?
We agree that solutions with mixed boundary conditions, free slip at the bottom and a prescribed, non-zero Dirichlet boundary condition for velocity at the top, inspired by kinematically-driven models (from plate motion histories), are very much of interest as well. The derivation of another such case would be fairly straightforward in principle, but there is a non-trivial amount of practical work involved. The procedure would be similar to the zero-slip and free-slip cases presented, albeit with some adjustments to the conditions imposed, where the solution can again be derived automatically using sympy or sage. What is far more involved is to transform the often unwieldy sympy/sage solutions into something that can be compactly expressed in a manuscript. The required hand-editing of the solutions is the reason why we introduced automated tests that extract the final latex expression of the solutions and checks their correctness. More importantly we think that adding just a single simple case, the most obvious would be a rigid rotation around a single axis, would be of limited value. To properly test the capability to simulate plate-driven models would require benchmarks that in-
clude shear, convergence and divergence in the prescribed velocity. We would be willing to consider, if the interest arises, to include such cases in the assess python package (where readability is of less concern), but we believe adding an appendix with the fully worked out solution would add little to the reader of the
manuscript.

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- p. 6, line 9 and equation (21): Shouldn't this read $R_{-}<r^{\prime}<R_{+}$instead of $R_{-} \leq r^{\prime} \leq R_{+}$and the piecewise $\psi$ be formulated for the two parts $R_{-}<r^{\prime}$ and $r^{\prime}<R_{+}$only? Just to be mathematically more precise?
This has been corrected as suggested.
- Equations (25) and (26): To me the transition from (25) to (26) seemed mathematically quite involved, e.g. if one was to evaluate the integral in the right-hand side of (25) one would get a zero. Maybe you could add some additional details on this transition?

The transition between (25) and (26) was indeed not entirely mathematically sound. This was the consequence of a failed attempt to simplify the argument. We have now corrected the derivation of (26) and inserted two additional steps.

- p. 7, line 16: You are making use of the Mie representation of the velocity field.The necessary condition for this to have the form (27) is that the velocity field is solenoidal. For $R^{3}$ this is equivalent to $u$ being divergence-free. However, in the case of the target domain, the thick spherical shell, the two properties are not the same. You get that $u$ is solenoidal from $u$ being divergence-free and the fact the you have no outflow in the boundary conditions you consider. You might want to reformulate the sentence in this respect.
We have clarified that the condition for Mie's representation is a solenoidal velocity which indeed follows from divergence freeness in combination with the no normal flow conditions in our domains.
- If I have not missed it, you do not explicitly specify how $r^{\prime}$ was chosen in you numerical tests. I assume that this is an interface between layers of the mesh already for the coarses mesh resolution used, isn't it? Could you please add that detail from completeness.

Out of curiosity, would you expect to observe oscillations in the convergence behaviour, if $r^{\prime}$ was not a layer boundary on the meshes? I remember that in dipole modelling (geoelectricity and EEG simulation) people sometimes resort to special discretisation approaches for the $\delta$-function, e.g. St. Venant's principle, to avoid such issues.
The value that was used for $r^{\prime}$ (and in fact $R_{-}$and $R_{+}$) was indeed missing in the paper. They were: $R_{-}=1.22, r^{\prime}=1.72, R_{+}=2.22$ following a standard non-dimensionalisation in which the Mantle depth is one. In other words, $r^{\prime}$ was always halfway between top and bottom, and indeed always coinciding with a grid level. This allows us to treat the contribution from the infinitely thin forcing as a surface integral as explained in section 3.3. Other authors [e.g. Zhong et al. '08] have approximated the delta function with a finite element basis function with a support over the two adjacent cells and a grid-dependent amplitude that goes to infinity as $h \rightarrow 0$. In local experiments I have seen similar convergence issues with this kind of forcing, and I suspect the same would be true if the location were not grid-aligned. It might be that other ways of smoothing the forcing would improve the convergence. Fundamentally though, convergence at the ideal order would not be guaranteed through the classical finite element analysis.

- Sec. 3.4: The problem matrix representing the discretised Stokes system is singular independent of the type of boundary conditions due to the pressure only being determined up to an additive constant, isn't it? The free-slip conditions enhance the kernel of the matrix significantly leading to higher numerical effort (performing step (64) in each iteration seems required, while pressure needs only be adapted following (62) once in the end); you might want to make that clearer


## in line 15.

The combined velocity pressure saddle point system is indeed singular in all cases (see also line 6, page 14) The projections in equations (62) and (63) are big L2-projections and are both only performed once after the entire solving process has finished. To clarify lines 14-15 on page 14: 'It should be noted that the same modes...'; This sentence should really start a new paragraph and 'these modes' refer to both the pressure and velocity modes described just before that. The corresponding (little) I2-projections are performed in the iterative solution process performed by PETSc, which consists of an outer (FGMRES) iteration solving the Schur complement equation in which the required matrix vector multiplication with the inverse of the velocity block induces an inner (CG) iteration. The little I2 projection corresponding to the velocity modes are subtracted every inner iteration. The little I2 projection corresponding to constant pressure is subtracted every outer iteration. We hope this clears up the final one of your comments. We have tried to clarify this paragraph in revision. Without going into to much details on the solver strategy (which is also described in Davies et al. '11), the point we try to bring across is that the null-modes you have to provide to ensure convergence of the iterative solvers, which because of their abstraction are typically formulated in an I2-inner product context, are insufficient to ensure convergence to the analytical solution because of the difference between I2 and L2 projections which in particular for non-uniform meshes can be significant.

## Suggestions:

p. 4, line 19 and p. 9, line 9: A reference for the biharmonic equation resp. its solutions might be helpful for the general audience; it is not quite as common knowledge as the harmonic equation and its solutions

- p. 4, line 14: You might consider changing 'top' and 'bottom' to 'inner and outer' for the cylindrical domain
- p. 6, line 18: Shouldn't the strip be defined as $\left(r^{\prime}-\epsilon, r^{\prime}+\epsilon\right) \times(0,2 \pi)$ ?
- p. 9, line 3: The standard definition of the term domain in calculus is an open and connected set, so connected domain sounds like a pleonasm ;-)
- p. 4, equation (19) and line 19: I must admit that as I reader I very much behave like a one-pass-compiler, as soon as I encounter something I cannot follow I stumble. For the sake of people like me you might consider moving that sentence which explains why there's only a cos term in (19) up front. Also you might state that you consciously neglect other harmonic functions (such as $\ln (r)$ ), for similar reasons as given in line 13.
- p. 1, line 2: IMHO 'within' does not sound quite proper here?
- In your paper you are using the term 'natural' boundary condition. If I understand correctly, you mean 'inspired' by nature? I am asking as in classical FE analysis there is that distinction between 'natural' and 'essential' boundary conditions, so I was at first glance a little confused. Maybe change it to 'physical', if that still fully expresses what you want to convey.
- p. 1, line 24: Aren't 3D global mantle convection models being simulated routinely today, and not only becoming more common? I mean, you site references from the last 35 years ;-)
- equation (A7): Is there any specific motivation for defining the 2D polar curl in this way? It seems to be just the negative of what one would obtain by constantly extending a 2D field in z-direction and taking the z-component of the curl in 3D cylindrical coordinates?

Your list of suggestions are all very much to the point and have been incorporated in the revised manuscript. Regarding, the sign in the 2D curl definition in the appendix this is indeed not standard and not as we intended. What I think happened was that I worked back from the curl of the streamfunction with a different sign convention for the streamfunction in mind. In any case we have corrected this to follow the usual definition. This does not affect the derivations in the rest of the paper.

## Questions out of curiosity

- I found your discussion of the reason for the reduced convergence rates in the $\delta$ function case with the $P_{2}-P_{1}$ Taylor-Hood element very interesting. As (68) only contains the $H_{1}$ semi-norm of $u$, is there an easy way to see from (70) why we get a similar $\frac{2}{3}$ order reduction in the $H_{0}$ norm of velocity as we do for pressure? Maybe I am missing some standard FE-analysis argument?
I believe the next step, to derive an L2 error bound for $u$ out of the H 1 seminorm bound, is the so called Aubin-Nitsche trick. The most compact write up that I could find is (4.1)-(4.3) in Verfuehrt'84. Here, in turn, you need to find a bound in the $H^{1}$-norm for the solution $z$ and numerical solution $z_{h}$ of an arbitrary righthand side $v$ in $L_{2}$ (rather than $H^{-} 1$ ) for which, with a domain of the right regularity, you can find a solution $v$ in $H^{2}$ (instead of $H^{1}$ ). The $H^{1}$-norm of the difference between $z$ and $z_{h}$ is then first order in $h$. This together with the $h^{1 / 2}$ bound we already had, gives you the expected $h^{3 / 2}$ (see also Theorem 1.9 in Girault). Note that this relies on having this $h^{1 / 2}$ bound in the first place which I believe isn't guaranteed for arbitrary $H^{-1}$ right-hand sides (hence Theorem 1.9 in Girault is only valid for H 2 velocity solutions, i.e. the original rhs should be L2). The bound in (70) is based on the assumption of a single jump in pressure where the number of simplices affected by this jump scales with $1 / h^{\mathrm{dim}-1}$ (i.e. occurs along a dim-1 "front" that can be well resolved).
- Do you have any (speculative) idea why that convergence issue is not observed when one only examines surface quantities?
I am speculating that a similar bound to (68) is possible where the numerical error in the solution restricted to the boundary is related to the best possible approximation on the boundary for the given discrete function space. Note that the bound in (70) for the best possible approximation of the analytical solution $p$ is actually just bounded by the (worst) regularity of $p$ in the triangles near the jump (you can derive it from the generic finite element approximation error bounds in Bernardi '89). However this does not mean that the numerical solution of the discretised PDE has its $h^{1 / 2}$ error concentrated in these cells; this l've tried: even if you restrict the $L_{2}$ error integral to integrate only over parts of the domain well away from the jump, the convergence is still $h^{1 / 2}$. So I suspect this really only works for an integral just over the boundary, or a volume integral at some limited distance from the boundary where that distance scales to zero as $h \rightarrow 0$.
- In your 3D test cases you always select two similar combinations of degree and order $(l, m)$, which is $m=l$ (sectoral) and $m=l / 2$ (tesseral); I was a little surprised that the errors seem to be fully identical for the two choices, because the number and direction of nodal lines differs. Are the differences just too small to be visible in the figures? Can you comment on that?
We believe this is just an artefact of the choice of $k$, the degree of the radial dependence, where the combinations you mention that seems to have similar error, also have the same degree $k$. So it seems that for the ratio of tangential and radial resolution we have chosen the tangential variation is well resolved and the error is probably more restricted by the vertical resolution.

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## Technical Corrections

- equation (14): please check sign, I might have miscalculated, but I think it should read $H_{n}=+\ldots$
- p. 6, line 15: 'expect a continuity' $\rightarrow$ 'expect continuity'
- p. 21, line 12: solution $\rightarrow$ solution(s)
- equation (A3): $\hat{\varphi} \cdot \nabla \hat{\varphi}=\frac{1}{r} \frac{\partial \hat{r}}{\partial \varphi} \quad \rightarrow \quad \frac{1}{r} \frac{\partial \hat{\varphi}}{\partial \varphi}$
- equation (A12): spurious + near end of equation
- reference Hernlund, Tackley, 2007: IMHO that should be 2008

I have double checked (14), and I must admit at first I came to the same conclusion of $\mathrm{a}+\mathrm{in}$ the $H_{n}$ equation. This confused me a lot as we had been using a minus sign in our convergence analysis. Luckily, after going through the derivation once more, and more carefully, I am now convinced the original minus sign is correct.

For completeness let me include 'my working'. Assume $p=H_{n} r^{-n} \cos (n \varphi)$ and $\psi=$ $D_{n} r^{-n+2} \sin (n \varphi)$, then

$$
\begin{aligned}
\frac{\partial p}{\partial r} & =-n r^{-n-1} H_{n} \cos (n \varphi), \\
\frac{1}{r} \frac{\partial p}{\partial \varphi} & =-n r^{-n-1} H_{n} \sin (n \varphi), \\
\nabla^{2} \psi & =\left((-n+2)^{2}-n^{2}\right) D_{n} r^{-n} \sin (n \varphi) \\
& =-4(n-1) D_{n} r^{-n} \sin (n \varphi), \\
\frac{\nu}{r} \frac{\partial \nabla^{2} \psi}{\partial \varphi} & =-4 \nu n(n-1) D_{n} r^{-n-1} \cos (n \varphi),
\end{aligned}
$$

$$
-\nu \frac{\partial \nabla^{2} \psi}{\partial r}=-4 \nu n(n-1) D_{n} r^{-n-1} \sin (n \varphi) .
$$

So that equations (7) and (8) become respectively:

$$
\begin{array}{ll}
-n\left[4 \nu(n-1) D_{n}+H_{n}\right] r^{-n-1} \cos (n \varphi) & =0, \\
-n\left[4 \nu(n-1) D_{n}+H_{n}\right] r^{-n-1} \sin (n \varphi) & =0
\end{array}
$$

which indeed implies that $H_{n}=-4 \nu(n-1) D_{n}$ as in equation (14).
We would like to thank the reviewer, once again, for their very thorough and constructive review of our derivations and the manuscript. It was clearly a huge effort and we are very grateful. The other technical corrections have been included in the revised manuscript.

## References

References HMB20, Thi17, BMP16, PLP14 as in your referee comment.
Bernardi '89: Christine Bernardi, "Optimal Finite-Element Interpolation on Curved Domains", SIAM Journal on Numerical Analysis 26.5 (1989): 1212-1240, https://doi.org/10.1137/0726068

Verfuhrt '84: R. Verfuhrt, "A combined conjugate gradient - multi-grid algorithm for the numerical solution of the Stokes problem", IMA J Numer Anal, Volume 4, Issue 4, 1984, Pages 441-455, https://doi.org/10.1093/imanum/4.4.441

All other references appear in the submitted paper.
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