

HighResMIP versions of EC-Earth: EC-Earth3P and EC-Earth3P-HR. Description, model computational performance and basic validation

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Abstract A new global high-resolution coupled climate model, EC-Earth3P-HR has been developed by the EC-Earth consortium, with a resolution of approximately 40 km for the atmosphere and 0.25 degree for the ocean, alongside with a standard resolution version of the model, EC-Earth3P (80 km atmosphere, 1.0 degree ocean). The model forcing and simulations follow the HighResMIP protocol. According to this protocol all simulations are made with both high and standard resolutions. The model has been optimized with respect to scalability, performance, data-storage and post-processing. In accordance with the HighResMIP protocol no specific tuning for the high resolution version has been applied.

Increasing horizontal resolution does not result in a general reduction of biases and overall improvement of the variability, and deteriorating impacts can be detected for specific regions and phenomena such as some Euro-Atlantic weather regimes, whereas others such as El Niño-Southern Oscillation show a clear improvement in their spatial structure. The omission of specific tuning might be responsible for this.

The shortness of the spin-up, as prescribed by the HighResMIP protocol, prevented the model to reach equilibrium. The trend in the control and historical simulations, however, appeared to be similar, resulting in a warming trend, obtained by subtracting the control from the historical simulation, close to the observational one.

1 Introduction

Recent studies with global high-resolution climate models have demonstrated the added value of enhanced
45 horizontal atmospheric and oceanic resolution compared to the output from models in the coupled model
intercomparison project phase 3 and 5 (CMIP3 and CMIP5) archive. An overview and discussion of those studies
has been given in Haarsma et al. (2016) and Roberts et al. (2018). Coordinated global high-resolution experiments
were, however, lacking, which induced the launch of the CMIP6 endorsed High Resolution Model
50 Intercomparison Project (HighResMIP). The protocol of HighResMIP is described in detail in Haarsma et al.
(2016). Due to the large computational cost that high horizontal resolution implies, the time period for simulations
in the HighResMIP protocol ranges from 1950 to 2050. The minimal required atmospheric and oceanic resolution
for HighResMIP is about 50 km and 0.25° respectively.

EC-Earth is a global coupled climate model (Hazeleger et al., 2010, 2012) that has been developed by a consortium
55 of European institutes consisting to this day of 27 research institutes. Simulations with EC-Earth2 contributed to
the CMIP5 archive, and numerous studies performed with the EC-Earth model appeared in peer-reviewed
literature and contributed to the fifth assessment report (AR5) of the IPCC (Intergovernmental Panel on Climate
Change) (IPCC, 2013). EC-Earth is used in a wide range of studies from paleo-research to climate projections,
including also seasonal (Bellprat et al. 2016; Prodhomme et al., 2016; Haarsma et al., 2019) and decadal forecasts
60 (Guemas et al., 2013, 2015; Doblas-Reyes et al., 2013; Caron et al., 2014, Solaraju Murali et al., 2019, Koenigk
et al., 2013, Koenigk and Brodeau, 2014, Brodeau and Koenigk, 2016).

In preparation for CMIP6, a new version of EC-Earth, namely EC-Earth3, has been developed (Doescher et al.,
2020). This has been used for the DECK (Diagnostic, Evaluation and Characterization of Klima) simulations
65 (Eyring et al., 2016) and several CMIP6-endorsed MIPs. The standard resolution of EC-Earth3 is T255 (~80 km)
for the atmosphere and 1.0° for the ocean, which is too coarse to contribute to HighResMIP. A higher resolution
version of EC-Earth3, therefore, had to be developed. In addition, the HighResMIP protocol demands simplified
aerosol and land schemes (Haarsma et al., 2016).

In section 2, we will describe the HighResMIP version of EC-Earth3 which has been developed within the
European Horizon2020 project PRIMAVERA (Roberts et al., 2018). For a detailed description of the standard
CMIP6 version of EC-Earth3 and its technical and scientific performances, we refer to Doescher et al. (2020).
High-resolution modeling requires special efforts on scaling, optimization and model performance, which will be
discussed in section 3. In section 3 we also discuss the huge amount of data that is produced by a high-resolution
75 climate model, that requires an efficient post-processing and storage workflow. A summary of the model results
will be given in section 4. In that section we also discuss the issue that for a high resolution coupled simulation it
is not possible to produce a completely spun up state that has reached equilibrium due to limited computer
resources. As a result, the HighResMIP protocol prescribes that the simulations start from an observed initial state.
The drift due to an imbalance of the initial state is then accounted for by performing a control run with constant
80 forcing alongside the transient run.

2 Model description

85 The model used for HighResMIP is part of the EC-Earth3 family. EC-Earth3 is the successor of EC-Earth2 that was developed for CMIP5 (Hazeleger et al., 2010, 2012; Sterl et al., 2012). Early versions of EC-Earth3 have been used by e.g. Batté et al. (2015), Davini et al. (2015) and Koenigk and Brodeau (2017). The versions developed for HighResMIP are EC-Earth3P (T255 (~100 km) atmosphere, 1 degree ocean) for standard resolution and EC-Earth3P-HR (T511 (~50 km) atmosphere, 0.25 degree ocean) for high resolution and will henceforth be referred to as EC-Earth3P(-HR), respectively. In addition, a very high resolution version EC-Earth3P-VHR
90 (T1279 (~15 km) atmosphere, 0.12 degree ocean) has been developed and simulations following the HighResMIP protocol are presently being performed, but not yet available. Compared to EC-Earth2, EC-Earth3P(-HR) include updated versions of its atmospheric and oceanic model components, as well as a higher horizontal and vertical resolution in the atmosphere.

95 The atmospheric component of EC-Earth is the Integrated Forecasting System (IFS) model of the European Centre for Medium-Range Weather Forecasts (ECMWF). Based on cycle 36r4 of IFS, it is used at T255 and T511 spectral resolution for EC-Earth3P and EC-Earth3P-HR, respectively. The spectral resolution refers to the highest retained wavenumber in linear triangular truncation. The spectral grid is combined with a reduced Gaussian grid where the nonlinear terms and the physics are computed, with a resolution of N128 for EC-Earth3P, N256 for EC-Earth3-
100 HR and N640 for EC-Earth3P-VHR. Because of the reduced Gaussian grid the grid box distance is not continuous, with a mean value of 107 km for EC-Earth3P and 54.2 km for EC-Earth3P-HR (Klaver et al., 2020). The number of vertical levels is 91, vertically resolving the middle atmosphere up to 0.1 hPa. The H-TESSSEL model is used for the land surface (Balsamo et al., 2009) and is an integral part of IFS: for more details see Hazeleger et al. (2012).

105 The ocean component is the Nucleus for European Modelling of the Ocean (NEMO; Madec, 2008). It uses a tri-polar grid with poles over northern North America, Siberia and Antarctica and has 75 vertical levels (compared to 42 levels in the CMIP5 model version and standard EC-Earth3). The so-called ORCA1 configuration (with a horizontal resolution of about 1 degree) is used in EC-Earth3P whereas the ORCA025 (resolution of about 0.25
110 degree) is used in EC-Earth3P-HR. The ocean model version is based on NEMO version 3.6 and includes the Louvain-la-Neuve sea-ice model version 3 (LIM3; Vancoppenolle et al., 2012), which is a dynamic-thermodynamic sea-ice model with five ice thickness categories. The atmosphere/land and ocean/sea-ice components are coupled through the OASIS (Ocean, Atmosphere, Sea Ice, Soil) coupler (Valcke and Morel, 2006; Craig et al., 2017).

115 The NEMO configuration is based on a set-up developed by the ShaCoNEMO initiative lead by Institute Pierre Simon Laplace (IPSL) and adapted to the specific atmosphere coupling used in EC-Earth. The remapping of runoff from the atmospheric grid points to runoff areas on the ocean grid has been re-implemented to be independent of the grid resolution. This was done by introducing an auxiliary model component and relying on the interpolation
120 routines provided by the OASIS coupler. In a similar manner, forcing data for atmosphere-only simulations are

passed through a separate model component, which allows to use the same SST and sea-ice forcing data set for different EC-Earth configurations.

125 IFS and NEMO have the same time steps: 45 min in EC-Earth3P and 15 min in EC-Earth3P-HR. The coupling between IFS and NEMO is 45 min in both configurations.

The CMIP6 protocol requests modeling groups to use specific forcing datasets that are common for all participating models. Table 1 lists the forcings that have been implemented in EC-Earth3P(-HR). Because of the HighResMIP protocol, EC-Earth3P(-HR) distinguish themselves in several aspects from the model configurations used for the CMIP6 experiments (Doescher et al., 2020):
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- 135 1. The stratospheric aerosol forcing in EC-Earth3P(-HR) is handled in a simplified way that neglects the details of the vertical distribution and only takes into account the total aerosol optical depth in the stratosphere which is then evenly distributed across the stratosphere. This approach follows the treatment of stratospheric aerosols as it was used by EC-Earth2 for the CMIP5 experiments yet with the stratospheric aerosol optical depth (AOD) at 500 nm updated to the CMIP6 data set.
- 140 2. A sea surface temperature (SST) and sea-ice forcing data set specially developed for HighResMIP is used for AMIP experiments (Kennedy et al., 2017). The major differences compared to the standard SST forcing data sets for CMIP6 are the higher spatial (0.25 deg vs. 1 deg) and temporal (daily vs. monthly) resolution. For the Tier 3 HighResMIP SST forced future AMIP simulations (see section 4.1) an artificially produced data set of SST and sea ice concentration (SIC) is used that combines observed statistics and modes of variability with an extrapolated trend (<https://esgfnod.llnl.gov/search/input4mips/>).
- 145 3. The HighResMIP protocol requires the simulations to start from an atmosphere and land initial state from the 1950 of the ECMWF ERA-20C (Poli et al., 2016) reanalysis data. Because the soil moisture requires at least 10 years to reach equilibrium with the model atmosphere, a spin-up of 20 years under 1950 forcing has been made before starting the Tier 1 simulations.
- 150 4. In agreement with the HighResMIP protocol, the vegetation is prescribed as a present-day climatology that is constant in time.
- 155 5. The climatological present-day vegetation, based on ECMWF ERA-Interim (Dee et al., 2011), and specified as albedos and leaf area index (LAI) from the Moderate-resolution Imaging Spectroradiometer (MODIS) is used throughout all runs. In contrast, the model version for other CMIP6 experiments uses lookup tables to account for changes in land-use. In addition, that version is consistent with the CMIP6 forcing data set and not based on ERA-Interim.
- 160 6. Another difference is the version of the pre-industrial aerosols background derived from the TM5 model (Van Noije et al., 2014; Myriokefalitakis et al., 2020, and references therein): version 2 in PRIMAVERA, version 4 in other CMIP6 model configurations using prescribed anthropogenic aerosols. This affects mainly the sea-spray source, and in turn the tuning parameters.

3 Model performance and data handling

165 New developments in global climate models require special attention in terms of high-performance computing (HPC) due to the demand for increased model resolution, large numbers of experiments and increased complexity of Earth System Models (ESMs). EC-Earth3P-HR (and VHR) is a demanding example where an efficient use of the resources is mandatory.

170 The aim of the performance activities for EC-Earth3P-HR is to adapt the configuration to be more parallel, scalable and robust, and to optimize part of the execution when this high-resolution configuration is used. The performance activities are focused on three main challenges: (1) scaling of EC-Earth3P-HR to evaluate the ideal number of processes for this configuration, (2) analyses of the main bottlenecks of EC-Earth3P-HR and (3) new optimizations for EC-Earth3P-HR.

3.1 Scalability

175 The results of the scalability analyses of the atmosphere (IFS) and ocean (NEMO) components of EC-Earth3P-HR are shown in Fig. 1, and for the fully coupled model in Fig. 2. Acosta et al. (2016) showed that, while for coupled application the load balance between components has to be taken into account in the scalability process, the process needs to start with a scalability analysis of each individual component. Moreover, the user could experience that the speeding up of one component (e.g. the reduction of the execution time of IFS) does not reduce the execution time of the coupled application. This could be because there is one synchronization point at the end each coupled time-step, where both components exchange fields. If the other non-optimized components are slower, a load rebalance will be required. The final choice depends on the specific problem, where either time or energy can be minimized. In section 3.2, we describe how the optimal load balance between the two components, where NEMO is the slowest component, was achieved (Acosta et al., 2016).

3.2 Bottlenecks

190 For the performance analysis, the individual model components (IFS, NEMO and OASIS) are benchmarked and analyzed using a methodology based on extracting traces from real executions. These traces are displayed using the Paraver software and processed to discover possible bottlenecks (Acosta et al., 2016). Eliminating these bottlenecks not only involves an adjustment of the model configuration and a balance of the number of cores devoted to each one of its components, but also modifications of the code itself and work on the parallel programming model adopted in the different components.

195 The first step of a performance analysis consists in analyzing parallel programming model codes using targeted performance tools. Figure 3a illustrates an example of the performance tool's output from one single EC-Earth3P-HR model execution as provided by the Paraver tool, focusing only on its two main components: NEMO and IFS. This figure is very useful to determine the communications within the model and identify sources of bottlenecks, especially those resulting from communication between components. It displays the communication pattern as a function of time. The vertical axis corresponds to the different processes executing the model, the top part for IFS

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and the lower part for NEMO. The different colors correspond to different MPI communication functions, except the light blue, which corresponds to no communication. Red, yellow and purple colors are related to MPI communications. The green color represents the waiting time needed to synchronize the coupled model for the next time step, which means an unloaded balance in the execution. In summary, light blue areas are pure computation and should be maximized. On the other hand, yellow, red and purple are representing overhead from parallel computation and should be minimized if possible. Additionally, green areas are preferably to be also reduced, for example increasing the number of parallel resources of the slowest component, but no optimizations are needed. From this analysis, several things can be concluded related to the overhead from parallel computation:

210 1) Figure 3 shows the coupling cost from a computational point of view, including one regular time step of IFS and NEMO and one time step including the coupling process. In the top part of Fig. 3a, we notice that during the first half of the first time step, the IFS component model reserves most of its processors for execution (512 processes). To simplify, it can be said that the first half of the time step has less MPI communication, with more computation-only regions, while the second half of the time step is primarily about broadcasting messages (yellow and white colour block), which corresponds to the coupling computation and to send/receive files from the atmospheric to the ocean model. These calculations impact the scalability of the code dramatically. This configuration increases the overhead when more and more processes are used and represents more than 50% of time execution when 1024 processes are used. The coupling process can be analyzed in detail in Fig. 3a (Coupling zoom, top image), where the same pattern of communications is repeated four times. This occurs because the different fields from IFS to NEMO are sent in three different groups, followed by an additional group of fields sent from IFS to the runoff mapper component. The communication of three different groups of fields to the same component is not taking advantage of the bandwidth of the network, thus increasing the overhead produced by MPI communications. However, these three groups are using the same interpolation method and they could be gathered into the same group.

220 2) From other parts of the application, (not shown in the figure) we also notice the expensive cost of the IFS output process for each time step. A master process gathers the data from all MPI sub-domains and prints the complete outputs at a regular time interval of three and six hours. During this process, the rest of processes are waiting for this step to be completed. Due to the large data volumes, this sequential process is very costly, increasing the execution time of IFS by about 30% when outputs are required, compared to the regular time-step of IFS (without output).

235 3) The bottom part of Fig. 3a shows that the communication in NEMO is not very effective and that a large part of it is devoted to global communications, which appear in purple. Those communications belong to the horizontal diffusion routine, inside the ice model (LIM3) used in NEMO. The high frequency of communications in this routine prevented the model to scale. More information about MPI overhead of NEMO can be found in Tintó et al. (2019).

240 4) Due to the domain decomposition used by NEMO some of the MPI processes, which are used to run part of the ocean domain in parallel, were computing without use. This is because domain decomposition is done on a

regular grid and a mask is used to discriminate between land and sea points. The mask creates subdomains of land points whose calculations are not used. This is illustrated in Fig. 4 showing a particular case in which 12 % of the depicted subdomains do not contain any sea-point.

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250 3.3 New optimizations for the specific configuration

According to the profiling analysis done, different optimizations were implemented to improve the computational efficiency of the model:

255 1) The optimization (“opt”) option of OASIS3-MCT was used. This activates an optimized global conservation transformation. Using this option, the coupling time from IFS to NEMO is reduced by 90% for EC-Earth3P-HR. This is because all-to-one/one-to-all MPI communications are replaced by global communications (gather/scatter and reduction) and the coupling calculations are done by all the IFS processes instead of only the IFS master process.

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Another functionality of OASIS consists in gathering all fields sent from IFS to NEMO in a single group (Acosta et al., 2016). Coupling field gathering, an option offered by OASIS3-MCT, can be used to optimize coupling exchanges between components. The results show that gathering all the fields that use similar coupling transformations reduces the coupling overhead. This happens because OASIS3-MCT is able to communicate and interpolate all of the fields gathered at the same time. Fig. 3a (Coupling zoom, bottom image) proves that the collection of the first three groups reduces the communication patterns from four to two, where the execution time of this part is reduced significantly (40%).

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Figure 3b shows the execution when “opt” and “gathering” options are used, with the 90% reduction in coupling time clearly visible (large green section). In the case of the first time step in the trace, the coupling time is replaced by waiting time, since NEMO is finishing its time step and both components have to exchange fields at the end of the time step.

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2) For the output problem, the integration of XIOS as the I/O server for all components of EC-Earth can increase performance dramatically. XIOS is already used for the ocean component NEMO and the I/O server receiving also all the data from IFS processes and doing the output work in parallel and in an asynchronous way is the best solution to remove the sequential process when an IFS master process is required to do this work. This is being developed and will be included in the next version of EC-Earth.

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3) Based on the performance analysis, the amount of MPI communications can be reduced (Tintó et al., 2019) achieving a significant improvement in the maximum model throughput. In the case of EC-Earth3P-HR, this translated into a reduction of 46% in the final execution time.

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285 4). Using the tool ELPiN (Exclude Land Processes in NEMO) the optimal domain decomposition for NEMO has
been implemented (Tintó et al., 2017), with computation of only ocean subdomains and finding the most efficient
number of MPI processes. This substantially improves both the throughput and the efficiency (in case of 2048
processor cores 41% faster using 25% less resources). The increase in throughput was due to less computations
and related to that less communications. In addition, ELPiN allows for the optimal use of the available resources
in the domain decomposition depending on the shape and overlap of the subdomains.

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3.4 Post-processing and data output

295 At the T511L91 resolution, the HighResMIP data request translates into an unprecedented data volume for EC-
Earth. Because the atmosphere component IFS is originally a numerical weather prediction (NWP) model, it
contains no built-in functionality for time-averaging the data stream during the simulation. The model was
therefore configured to produce the requested three-dimensional fields (except radiative fluxes on model levels,
which cannot be output by the IFS) on six-hourly basis and surface fields with three-hourly frequency. As a
consequence, the final daily and monthly averages for instantaneous fields have been produced from sampling at
300 these frequencies, whereas fluxes are accumulated in the IFS at every time step. Vertical interpolation to requested
pressure or height levels is performed by the model itself.

For the ocean model, the post-processing is done within NEMO by the XIOS library which can launch multiple
processes writing netCDF files in parallel, alleviating the I/O footprint during the model run. The XIOS
305 configuration XML files were extended to produce as many of the ocean and sea ice variables as possible.

The combination of the large raw model output volume, the increased complexity of the requested data and the
new format of the CMOR tables (Climate Model Output Rewriter, an output format in conformance with all the
CMIP standards) required a major revision of the existing post-processing software. This has resulted in the
310 development of the ece2cmor3 package. It is a python package that uses Climate Data Operators (CDO) [CDO,
2015] bindings for (i) selecting variables and vertical levels, (ii) time-averaging (or taking daily extrema), (iii)
mapping the spectral and gridpoint atmospheric fields to a regular Gaussian grid and (iv) computing derived
variables by some arithmetic combination of the original model fields. Finally, ece2cmor3 uses the PCMDI
CMOR-library for the production of netCDF files with the appropriate format and metadata. The latter is the only
315 supported step for the ocean output.

To speed up the atmosphere post-processing, the tool can run multiple CDO commands in parallel for various
requested variables. Furthermore, we optimized the ordering of operations, performing the expensive spectral
transforms on time-averaged fields wherever possible. We also point out that the entire procedure is driven by the
320 data request, i.e. all post-processing operations are set up by parsing the CMOR tables and a single dictionary
relating EC-Earth variables and CMOR variables. This should make the software easy to maintain with respect to
changes in the data request and hence useful for future CMIP6 experiments.

4.1 Outline of HighResMIP protocol

330 The protocol of the HighResMIP simulations consists of Tiers 1, 2 and 3 experiments, that represent simulations of different priority (1 highest, 3 lowest), and a spin-up procedure. The protocol also excludes specific tuning for the high resolution version compared to the standard resolution version. Below we give a short summary of the protocol. The experiment names in the CMIP6 data base are given in italics.

- Tier 1: Forced-atmosphere simulations 1950-2014; *highresSST-present*

335 The Tier 1 experiments are atmosphere only simulations forced using observed sea surface temperature for the period 1950-2014.

- Tier 2: Coupled simulations 1950-2050

340 The period of the coupled simulations is restricted to 100 years because of the computational burden brought about by the model resolution and the limited computer resources. The period 1950-2050 covers historical multi-decadal variability and near-term climate change. The coupled simulations consist of a spin-up, control, historical and future simulation.

-Spin-up simulation; *spinup-1950*

345 Due to the large computer resources needed, a long spin-up to (near) complete equilibrium is not possible at high resolution. Therefore, as an alternative approach an analyzed ocean state representative of the 1950s is used as the initial condition for temperature and salinity (Good et al., 2013, EN4 data set). To reduce the large initial drift a spin-up of about 50 years is made using constant 1950s forcing. The forcing consists of greenhouse gases (GHG), including O₃ and aerosol loading for a 1950s (~10 year mean) climatology. Output from the initial 50 year spin-up is saved to enable analysis of multi-model drift and bias, something that was not possible in previous CMIP exercises, with the potential to better understand the processes causing drift in different models.

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- Control simulation; *control-1950*

355 This is the HighResMIP equivalent of the pre-industrial control, but using fixed 1950s forcing. The length of the control simulation should be at least as long as the historical plus future transient simulations. The initial state is obtained from the spin-up simulation.

- Historical simulation; *hist-1950*

360 This is the coupled historical simulation for the period 1950-2014, using the same initial state from the spin-up as the control run.

- Future simulation; *highres-future*

This is the coupled scenario simulation 2015-2050, effectively a continuation of the *hist-1950* experiment into the future. For the future period the forcing fields are based on the CMIP6 SSP5-8.5 scenario.

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- Tier 3: Forced-atmosphere 2015-2050 (2100); *highresSST-future*

The Tier 3 simulation is an extension of the Tier 1 atmosphere-only simulation to 2050, with an option to continue to 2100. To allow comparison with the coupled integrations, the same scenario forcing as for Tier 2 (SSP5-8.5) is used.

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A schematic representation of the HighResMIP simulations is given in figure 5.

4.2 Main results of EC-Earth3P(-HR) HighResMIP simulations

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For each of the HighResMIP tiers more than one simulation was produced. An overview of the simulations is given in Table 2.

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The data is stored on the JASMIN server at CEDA (<http://www.ceda.ac.uk/projects/jasmin/>) and available from ESGF. During the PRIMAVERA project the data was analyzed at the JASMIN server. For the *highresSST-present* and *highresSST-future* simulations the ensemble members were started from perturbed initial states. These were created by adding small random perturbations from a uniform distribution over $[-5e-5, +5e-5]$ degree to the three-dimensional temperature field. For the *control-1950* and *hist-1950*, the end of the spin-up was taken as the initial condition of the first member. For the two extra members the initial conditions were generated by continuing the spin-up for 5 years after perturbing the fields that are exchanged between atmosphere and ocean. The *highres-future* members are the continuation of the *hist-1950* members.

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The Atlantic Meridional Overturning Circulation (AMOC) in the *control-1950* of EC-Earth3P had unrealistically low values of less than 10 Sv. It was therefore decided to change the ocean mixing parameters, which improved the AMOC. The main difference compared to the first ensemble member of EC-Earth3P is that the parameterization of the penetration of turbulent kinetic energy (TKE) below the mixed layer due to internal and inertial waves is switched off ($nn_tau=0$; Madec et al. 2016). The mixing below the mixed layer is an ad-hoc parameterization into the TKE scheme (Rodgers et al. 2014,) and is meant to account for observed processes that affect the density structure of the ocean's boundary layer. In EC-Earth3P, this penetration of TKE below the mixed layer caused a too deep surface layer of warm summer water masses in the North Atlantic convection areas which lead to a breakdown of the Labrador Sea convection within a few years and a strongly underestimated AMOC in EC-Earth3P. An additional minor modification compared to ensemble member 1 is an increased tuning parameter $m_lc (=0.2)$ in the TKE turbulent closure scheme that directly relates to the vertical velocity profile of the Langmuir Cell circulation. Consequently, the Langmuir Cell circulation is strengthened.

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The new mixing scheme was also applied to EC-Earth3P-HR, to ensure the same set of parameters for both versions of EC-Earth3P(-HR). The simulations with the new ocean mixing are denoted with 'p2' for the coupled simulations in Table 2. The atmosphere is unchanged and therefore the atmosphere simulations are denoted as 'p1'. Because of the unrealistic low AMOC in EC-Earth3P in the 'p1' simulations we focus on 'p2' for the coupled simulations.

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Below we will briefly discuss the mean climate and variability of the *highresSST-present*, *control-1950* and *hist-1950* simulations. The main differences between EC-Earth3P and EC-Earth3P-HR will be highlighted. In addition, the spin-up procedure for the coupled simulations, *spinup-1950*, will be outlined. A more extensive analysis of the HighResMIP simulations will be presented in forthcoming papers.

4.2.1 *highresSST-present*

The *highresSST-present* simulations will be compared with ERA5 (Hersbach et al., 2020) (1979-2014) except for precipitation where GPCP V2.3 (1979-2014) (Adler et al., 2003) data will be used. EC-Earth, GPCP and ERA5 data are regridded to a common grid (N128) before comparison. Seasonal means (Dec.-Feb. (DJF) and Jun.-Aug. (JJA)) will be analyzed. Ensemble mean fields will be displayed.

Due to the prescribed SST the largest surface air temperature (SAT) biases are over the continents (Fig. 6). The most negative biases are over the Sahara for DJF and Greenland in JJA while the largest positive biases are located over Antarctica in JJA and northeastern Siberia in DJF. Over most areas EC-Earth3P-HR is slightly too cold. Over most of the tropics the mean sea level pressure (MSLP) is underestimated, whereas over Antarctica and surrounding regions of the Southern Ocean it has a strong positive bias (Fig. 7). Further noteworthy is the positive bias south of Greenland during DJF. The largest precipitation errors are seen in the tropics over the warm pool regions in the Pacific and the Atlantic with too much precipitation (Fig. 8). The planetary wave structure of the geopotential height at 500 hPa (Z500) during DJF is well represented with the exception of the region south of Greenland (Fig. 9), which is consistent with the MSLP bias (Fig. 7a). The physical causes of the aforementioned biases can include a wide range of deficiencies in the parameterizations of cloud physics, land-surface and snow to mention a few. In forth coming papers this will be investigated in further detail.

Doubling of the atmospheric horizontal resolution has only modest impact on the large-scale structures of the main meteorological variables, as illustrated by the global MSLP, SAT, and precipitation (Fig. 10). For SAT the differences are generally less than 1 K, for MSLP 1 hPa except for the polar regions. A remarkable result is the worsening of the bias over Antarctica during JJA. Because the dynamics of the polar vortex, which is sensitive to horizontal resolution, is strongest during austral winter we speculate that this enhanced bias is associated with it. The exact mechanism falls outside the scope of this basic validation and will be explored in forthcoming studies. For precipitation the difference can be larger than 1.5 mm/day in the tropics. It is possible to conclude that the increase of resolution does not have a clear positive impact on the climatology of any of those variables. For instance for precipitation it results in an increase of the wet bias over the warm pool (compare with Fig. 8). Also measured by the root mean square error (RSME) (see figure captions for the numbers) the impact of resolution is small, in the order of 10% or less depending on the variable and the season. Enhancing resolution reduces the RMSE for SAT and MSLP, whereas it slightly enhances for precipitation.

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4.2.2 spinup-1950

450 As discussed in the outline of the HighResMIP protocol, the spin-up was started from an initial state that is based on observations for 1950. For the ocean this is the EN4 ocean reanalysis (Good et al., 2013) averaged over the 1950-1954 period, with 3m sea-ice thickness in the Arctic and 1m in the Antarctic. The atmosphere-land system was initialized from ERA-20C for 1950 January 1st, and spun-up for 20 years to let the soil moisture reach equilibrium. For the ocean no data assimilation has been performed, which can result in imbalances between the
455 density and velocity fields giving rise to initial shocks and waves.

During the first years of the spin-up there is a strong drift in the model climate (not shown). For the fast components of the climate system like the atmosphere and the mixed layer of the ocean the adjustment is in the order of one year, whereas the slow components such as the deep ocean require thousand years or more to reach
460 equilibrium. For the land component this is on the order of a decade. As a consequence after a spin-up of 50 years the atmosphere, land and upper ocean are approximately in equilibrium while the deeper ocean is still drifting. The largest drift occurs in the layer 100-1000 m with a drift of 0.5 °C/100 year. This drift also has an impact on the fast components of the climate system, which therefore still might reveal trends.

465 4.2.3 control-1950

After the spin-up, the SAT each of the three members of EC-Earth3P-HR is in quasi-equilibrium and the global mean temperature oscillates around 13.9 °C (Fig. 11-left, black). The ocean is still warming as expressed by a negative net surface heat flux in the order of -1.5 Wm⁻² (positive is upward) (Fig. 11-right, black). This imbalance
470 is reduced during the simulation, but without an indication that the model is getting close to its equilibrium state.

Contrary to EC-Earth3P-HR, the global annual mean SAT of EC-Earth3P displays a significant upward trend, with an indication of stabilizing after about 35 years (Fig. 11-left, red). This warming trend is caused by a large warming of the North Atlantic as revealed by Fig. 12 showing the difference between the first and last 10 years
475 of the *control-1950* run. This warming is caused by the activation of the deep convection in the Labrador Sea (not shown), that started about 10 years after the beginning of the control simulation, which was absent in the spin-up run. Associated with that also the AMOC shows an upward trend (see Fig. 17 below). This switch to a warmer state does not strongly affect the slow warming of the deeper ocean, which is reflected in a similar behavior of the net surface heat flux as for EC-Earth3P-HR (Fig. 11-right). The reasons for the initial absence of deep convection
480 in the Labrador Sea in EC-Earth3P and the difference with EC-Earth3P-HR are not clear and presently under investigation. Possible candidates are that the differences in ocean resolution affect the sea-ice dynamics and deep convection, but also changes in ocean temperature and salinity distribution may play a role.

The *control-1950* experiment is also analyzed to evaluate model performance of internally-generated variability
485 in the coupled system; the targets are: El Niño-Southern Oscillation (ENSO), the North Atlantic Oscillation (NAO), sudden stratospheric warmings (SSWs) and the Atlantic Meridional Overturning Circulation (AMOC).

ENSO

Figure 13 depicts the seasonal cycle of the Niño3.4 index (SST anomalies averaged over 5°S-5°N/170°W-120°W). As it was also shown for EC-Earth3.1 (Yang et al., 2019), both EC-Earth3P and EC-Earth3P-HR still have a systematic underestimation of the ENSO amplitude from late-autumn to mid-winter and yield the minimum in July, 1-2 months later than in observations. Increasing model resolution reduces the bias in early-summer (May-June) but worsens it in late-summer (July-August). Overall, EC-Earth3P-HR shows lower ENSO variability than EC-Earth3P, which following Yang et al.'s (2019) arguments suggests that the ocean-atmosphere coupling strength over the tropical Pacific is weaker in the high-resolution version of the model. On the other hand, Fig. 14 displays the spatial distribution of winter SST variability and the canonical ENSO pattern, computed as linear regression onto the Niño3.4 index. Increasing model resolution leads to a reduction in the unrealistic zonal extension of the cold tongue towards the western tropical Pacific, which was also present in EC-Earth3.1 (Yang et al., 2019) and is a common bias in climate models (e.g. Guilyardi et al., 2009): EC-Earth3P reaches longitudes of Papua New Guinea (Fig. 14a), while EC-Earth3P-HR improves its location (Figs. 14b), yet overestimated as compared to observations (Fig. 14c). Note that the reduction of this model bias is statistically significant (Fig. 14-bis). Consistently, the improvement in the cold tongue translates into a better representation of the ENSO pattern (Fig. 14-bottom). Nonetheless, the width of the cold tongue in EC-Earth3P-HR is still too narrow in the central tropical Pacific (see also Yang et al., 2019), which again is a common bias in climate models (e.g. Zhang and Jin, 2012). Both EC-Earth3P and EC-Earth3P-HR realistically simulate the wave-like structure of the ENSO teleconnection in the extratropics (Fig. 14-bottom).

On another matter, note that EC-Earth3P-HR (Fig. 14b) captures much better the small-scale features and meanderings along the western boundary currents, Kuroshio-Oyashio and Gulf Stream, and the sea-ice edge over the Labrador Sea than EC-Earth3P (Fig. 14a). In these three areas there is a substantial increase in SST variability (Fig. 14bis), which following Haarsma et al. (2019) is likely due to increasing ocean resolution rather than atmosphere resolution.

NAO

Figure 15 illustrates how EC-Earth3P(-HR) simulates the surface NAO and its hemispheric signature in the middle troposphere. The NAO (here measured as leading EOF of the DJF SLP anomalies over 20°N-90°N/90°W-40°E) accounts for virtually the same fraction of SLP variance in both model versions, i.e. 42.70% in EC-Earth3P (Fig. 15d) and 42.74% in EC-Earth3P-HR (Fig. 15e), and still slightly underestimates the observed one (~50% in ERA-Interim, Fig. 15f); the same applied to EC-Earth2.2 when compared to ERA-40 (Hazeleger et al. 2012). EC-Earth3P correctly captures the circumglobal pattern associated with the NAO at upper levels (e.g. Branstator, 2002; García-Serrano and Haarsma, 2017), particularly the elongated lobe over the North Atlantic and the two centers of action over the North Pacific (Fig. 15-top). A close inspection to the barotropic structure of the NAO reveals that the meridional dipole is shifted westward in EC-Earth3P-HR (Fig. 15b,e) as compared to EC-Earth3P (Fig. 15a,d), which according to Haarsma et al. (2019) could be related to increasing ocean resolution and a stronger forcing of the North Atlantic storm-track.

SSWs

530 Also, the simulation of SSW occurrence is assessed (Fig. 16); the identification follows the criterion in Palmeiro
et al. (2015). The decadal frequency of SSWs in EC-Earth is about 8 events per decade regardless model
resolution, which is underestimated when compared to ERA-Interim (~11 events per decade) but in the range of
observational uncertainty (e.g. Palmeiro et al., 2015; Ayarzagüena et al., 2019). The same underestimation was
535 diagnosed in EC-Earth3.1 (Palmeiro et al., 2020a). The reduced amount of SSWs is probably associated with a
too-strong bias at the core of the polar vortex, still present in EC-Earth3.3 (Palmeiro et al., 2020b). It is thus
concluded that increasing horizontal resolution does not affect the model bias in the strength of the polar vortex.
The seasonal cycle of SSWs in reanalysis is quite robust over the satellite period, showing one maximum in
December-January and another one in February-March (Ayarzagüena et al., 2019), which was properly captured
by EC-Earth3.1 in control, coupled simulations with fixed radiative forcing at year 2000 (Palmeiro et al., 2020a).
540 Here in *control-1950*, EC-Earth does not reproduce such bimodal cycle, with EC-Earth3P-HR (blue) yielding a
peak in January-February and EC-Earth3P (red) two relative maxima in January and March. Interestingly, the
seasonal cycle of SSWs over the historical, pre-satellite period shows a different distribution with a prominent
maximum in mid-winter and a secondary peak in late-winter, although it is less robust among reanalysis products
(Ayarzagüena et al., 2019). The impact of the radiative forcing on SSW occurrence deserves further research.

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AMOC

The AMOC index was computed as the maximum stream function at 26.5N and between 900 and 1200 m depth.
The annual AMOC index of EC-Earth3P-HR for the *control-1950* runs (Fig. 17-left, black) is about 15 SV, which
is lower than the values from the RAPID array (Smeed et al., 2019) that are measured since 2004 (Fig. 17, stars
550 in middle panel). It reveals interannual and decadal variability, without an evident trend. As already discussed at
the beginning of section 4.2.3, the AMOC of EC-Earth3P shows an upward trend (Fig. 17-left, red) associated
with the activation of convection in the Labrador Sea.

555 **4.2.4 *hist-1950***

The *hist-1950* ensemble simulations differ from the *control-1950* simulations by the historical GHG and aerosol
concentrations. The global mean annual temperature in EC-Earth3P-HR displays an increase similar to the ERA5
data set (Fig. 18-left). The warming seems to be slightly larger in the model. We remind, however, the enhanced
560 observed warming after 2014, which might result in a similar trend in the model simulations compared to
observations up to present day. The cooling due to the Pinatubo eruption in 1991 is clearly visible in all members
and the ensemble mean. The amplitude and period compare well with ERA5. On its part, the AMOC in EC-
Earth3P-HR reveals a clear downward trend in particular from the 1990s onward (Fig. 17-middle, black). This is
consistent with a slowdown of the Atlantic overturning due to global warming in CMIP5 models (Cheng et al.,
565 2013).

Similarly to *control-1950*, the *hist-1950* simulations with EC-Earth3P show an upward drift in SAT (Fig. 18-left,
red) and AMOC (Fig. 17-middle, red) that are smaller (SAT) or absent (AMOC) in EC-Earth3P-HR. The
HighResMIP protocol (Haarsma et al., 2016) of having a control and a historical simulation starting from the same

570 initial conditions was designed to minimize the consequences of such trends. Under the assumption that the model trend is similar for both simulations, it can be eliminated by subtracting the control from the historical simulation. Indeed the global annual mean SAT and the AMOC of *hist-1950* minus *control-1950* display a very similar behavior in EC-Earth3P and EC-Earth3P-HR (Fig. 18-right and Fig. 17-right) with an upward trend for SAT and a downward trend for the AMOC. For SAT the upward trend compares well with ERA5.

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Weather regimes

Another way to test the representation of the mid-latitude atmospheric flow, with a focus on the low frequency variability (5 – 30 days), is to assess how well the models reproduce the winter (DJF) Euro-Atlantic weather regimes (Corti et al., 1999; Dawson et al., 2012).

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The analysis has been applied here to the EC-Earth3P and EC-Earth3P-HR *hist-1950* simulations. Following recent works (Dawson, 2015; Strommen, 2019), we computed the regimes via k-means clustering of daily geopotential height anomalies at 500 hPa over 80W-40E/30N-85N. As a reference, we considered the ECMWF reanalysis data from ERA40 (1957-1978) and ERA-Interim (1979-2014). The clustering is performed in the space spanned by the first 4 Principal Components obtained from the reference dataset. More details on the technique used and, on the metrics discussed here can be found in Fabiano et al. (2020) and references therein. Each row in Fig. 19 shows the resulting mean patterns of the four standard regimes - NAO+, Scandinavian Blocking, Atlantic Ridge and NAO- - for ERA (top), EC-Earth3P (middle) and EC-Earth3P-HR (bottom). The regimes are quite well represented in both configurations. However, the matching is better in the standard resolution version both in terms of RMS and pattern correlation averaged over all regimes (see Table 3). Only the Scandinavian (Sc) blocking pattern is improved in the HR, whereas the other patterns are degraded. The most significant degradation is seen for the NAO- pattern, which is shifted westward in the HR simulation. The result for EC-Earth3P(-HR) goes in the opposite direction of what has been observed in Fabiano et al. (2020), where most models showed a tendency for improving the regime patterns with increased resolution. Concerning the regime frequencies, both model versions show a tendency to produce less NAO+ cases than the observations and more Atlantic Ridge cases (Fig. 19).

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Another quantity of interest is the persistence of the regimes, since models usually are not able to reach the observed persistence of the NAO+/- states (Fabiano et al., 2020). As stated in Table 3, this is also observed for the EC-Earth3P *hist-1950* simulations and the effect of the HR is to increase the persistence of NAO+, but decrease that of NAO-.

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Even if the HR is degrading the regime patterns, it produces a small but positive effect on the geometrical structure of the regimes. This is shown by the last two quantities in Table 3: the optimal ratio and the sharpness. The optimal ratio is the ratio between the mean inter-cluster squared distance and the mean intra-cluster variance: the larger the optimal ratio, the more clustered are the data. The sharpness is an indicator of the statistical significance of the regime structure in the dataset in comparison with a randomly sampled multinormal distribution (Straus et al., 2007). The closer the value is to 100, the more significant is the multimodality of the distribution. The sharpness tends to saturate at 100 for very long simulations, so the values reported in Table 3 are obtained from a bootstrap

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on 30 years randomly chosen. Both the optimal ratio and the sharpness are too low in the EC-Earth3P simulations, as is usually seen for all models. A significant increase with EC-Earth3P-HR is seen for the optimal ratio, and a smaller (non-significant) one is seen for the sharpness.

615 The increased resolution simulations have a stronger regime structure and are closer to the observations in this sense. However, the regime patterns are degraded in the HR version and this affects the resulting atmospheric flow. A similar result was obtained by Strommen et al. (2019), for a different version of EC-Earth and two other climate models.

620

5 Discussion and conclusions

As contribution of the EC-Earth consortium to HighResMIP, a new version of EC-Earth has been developed with two horizontal resolutions: the standard resolution EC-Earth3P (T255, ORCA1) and the high-resolution EC-
625 Earth3P-HR (T511, ORCA0.25). Simulations following the HighResMIP protocol (Haarsma et al., 2016) for all three tiers have been made using both resolutions, with an ensemble size of three members. Only the spin-up consists of one member.

630 Performing 100-yr simulations for the high-resolution version (EC-Earth3P-HR) required specific developments for the hard and soft ware to ensure efficient production, post-processing and storage of the data. In addition, the model must be able to run on different platforms with similar performance. Large efforts have been dedicated to scalability, reducing bottlenecks during performance, computational optimization and efficient post-processing and data output.

635 Enhancing resolution does not noticeably affect most model biases and there are even locations and variables where increasing the resolution has a deteriorating effect such as an increase of the wet bias over the warm pool seen in the *highresSST-present* simulations or the representation of Euro-Atlantic weather regimes found in the *hist-1950* experiments. Also, the variability reveals examples of improvement such as the zonal extension of the ENSO pattern or the representation of meandering along the western boundary currents, as revealed in the *control-*
640 *1950* simulations. The lack of re-tuning the high-resolution version of the model compared to the standard-resolution version, in accordance with the HighResMIP protocol, might be responsible for this.

The short spin-up as prescribed by the HighResMIP protocol prevented the simulations to reach an equilibrium state. This happened in particular for the *control-1950* and *hist-1950* simulations of EC-Earth3P where a transition
645 to a warmer state occurred due to enhanced convection in the Labrador Sea, with an accompanying increase of the AMOC. Because this transition occurred almost concurrently in the *control-1950* and *hist-1950* simulations the greenhouse-forced warming from 1950 onward in EC-Earth3P can be inferred by subtracting both simulations. The resulting warming pattern compares well with the observed one and is similar to the warming pattern simulated by EC-Earth3P-HR. Due to the transition, the *control-1950* does not provide a near-equilibrium state.

650 It was therefore decided to extend the *control-1950* run for another 100 year to allow process studies, that will be documented elsewhere.

Analysis of the kinetic energy spectrum indicates that the sub-synoptic scales are better resolved at higher resolution (Klaver et al., 2020) in EC-Earth. Despite the lack of a clear improvement with respect to biases and
655 synoptic scale variability for the high resolution version of EC-Earth, the better representation of sub-synoptic scales results in better representation of phenomena and processes on these scales such as tropical cyclones (Roberts et al., 2020) and ocean-atmosphere interaction along western boundary currents (Tsartsali et al. in preparation).

660 **Code/data availability,**

Model codes developed at ECMWF, including the IFS and FVM, are intellectual property of ECMWF and its member states. Permission to access the EC-Earth source code can be requested via the EC-Earth website <http://www.ec-earth.org> (The EC-Earth Consortium, 2019) and may be granted if a corresponding software
665 licence agreement is signed with ECMWF. The repo tags for the versions of IFS and EC-EARTH that are used in this work are 3.0p (see section 4.2, “p1” version) and 3.1p (“p2” version), and are available through r7481 and r7482 on ECSF respectively. The model code evaluated in this manuscript has been provided for anonymous review by the topical editor and anonymous reviewers.

The doi of the data used in the analyses and available from ESGF are:

670 EC-Earth3P doi:10.22033/ESGF/CMIP6.2322

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Forcing	Dataset	Version
Solar	http://solarisheppa.geomar.de/solarisheppa/c mip6	3.1
Well-mixed GHG concentrations	CMIP6_histo_mole_fraction_of_XX_in_air_input4MIPs_gr1-GMNHSH.nc from input4mips with XXX being carbon_dioxide, cfc11eq, cfc12, methane or nitrous_oxide	1.2.0
Tropospheric aerosols	Anthropogenic part: MACv2.0-SP_v1.nc Pre-industrial part: based on TM5	2.0
Stratospheric aerosols	Simplified approach. CMIP6 stratospheric AOD at 550 nm, vertically integrated	2.1.0
Ozone	vmro3_input4MIPs_ozone_CMIP6_UReading-CCMI from input4mips	1.0
Vegetation	Present day climatology. Vegetation type and cover from ERA-Interim. Albedo and LAI derived from MODIS. Same procedure as used for ERA-20C	
AMIP SST + SIC	HadISST2 from input4mips	2.2.0.0.

Table 1 CMIP6 forcing details.

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	<i>highresSST-present</i>	<i>highresSST-future</i>	<i>control-1950</i>	<i>hist-1950</i>	<i>highres-future</i>
EC-Earth3P-HR	3 members: r1i1p1f1 r2i1p1f1 r3i1p1f1	3 members: r1i1p1f1 r2i1p1f1 r3i1p1f1	4 members: r1i1p1f1 r1i1p2f1 r2i1p2f1 r3i1p2f1	4 members: r1i1p1f1 r1i1p2f1 r2i1p2f1 r3i1p2f1	4 members: r1i1p1f1 r1i1p2f1 r2i1p2f1 r3i1p2f1
EC-Earth3P	3 members: r1i1p1f1 r2i1p1f1 r3i1p1f1	3 members: r1i1p1f1 r2i1p1f1 r3i1p1f1	4 members: r1i1p1f1 r1i1p2f1 r2i1p2f1 r3i1p2f1	4 members: r1i1p1f1 r1i1p2f1 r2i1p2f1 r3i1p2f1	4 members: r1i1p1f1 r1i1p2f1 r2i1p2f1 r3i1p2f1

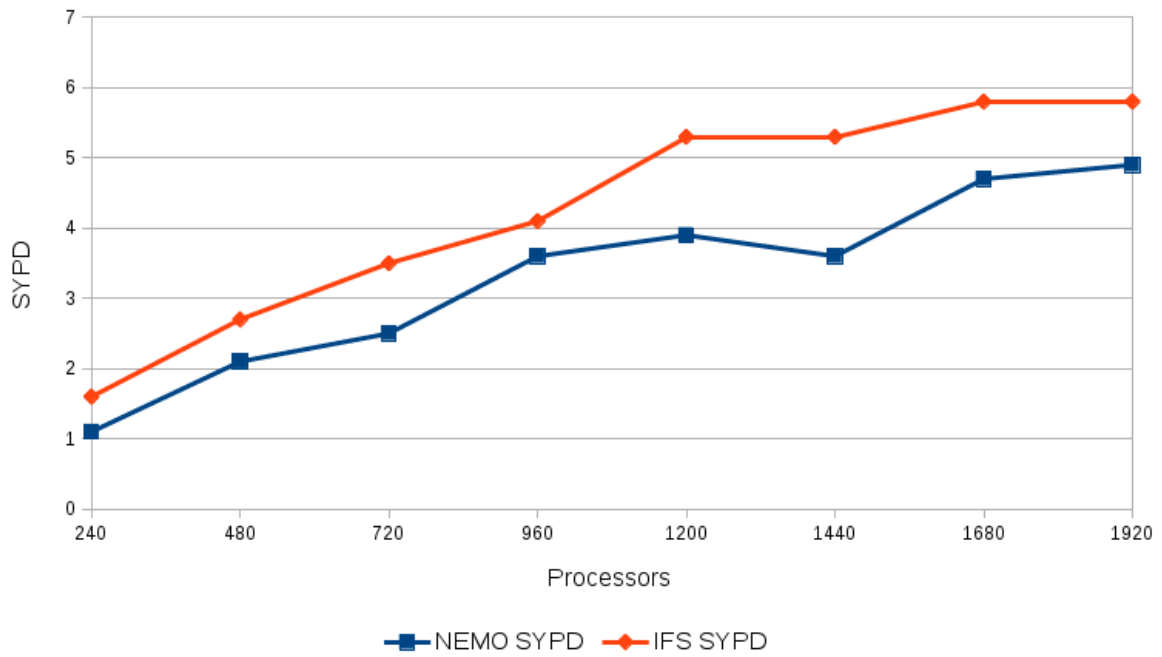
Table 2 Overview of the HighResMIP simulations of EC-Earth3P-HR and EC-Earth3P.

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	ERA	EC-Earth-3P	EC-Earth3P-HR
RMS (mean)	/	18 +/- 8 m	22 +/- 8 m
Patt. corr. (mean)	/	0.90 +/- 0.08	0.86 +/- 0.11
Av. persistence (NAO +)	6.1 days	5.4 +/- 0.2 days	5.7 +/- 0.5 days
Av. persistence (NAO -)	7.0 days	6.0 +/- 0.2 days	5.5 +/- 0.3 days
Optimal ratio	0.841	0.759 +/- 0.010	0.771 +/- 0.007
Significance (30 yr)	95.6	80.2 +/- 6.0	82.3 +/- 0.4

Table 3 Some metrics to assess the overall performance in *hist-1950* of the EC-Earth3P and EC-Earth3P-HR simulations in terms of weather regimes. The table shows: the average RMS deviation from the observed patterns and the relative average pattern correlation among all regimes; the average persistence of the two NAO states in days; the optimal ratio and the sharpness. The errors refer to the spread between members (standard deviation).

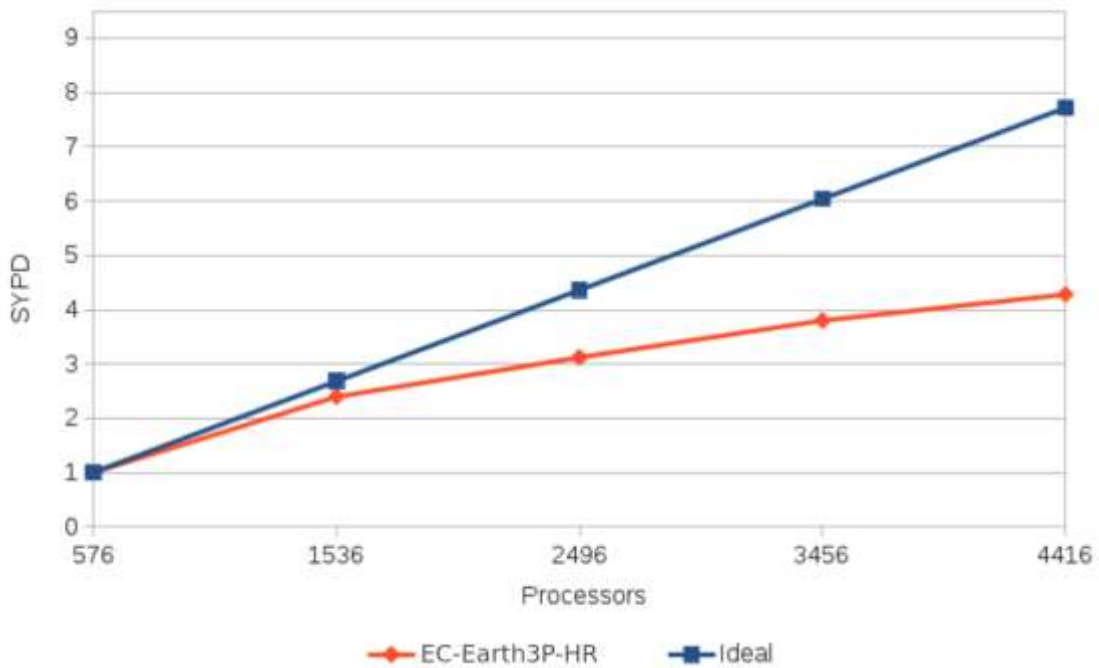
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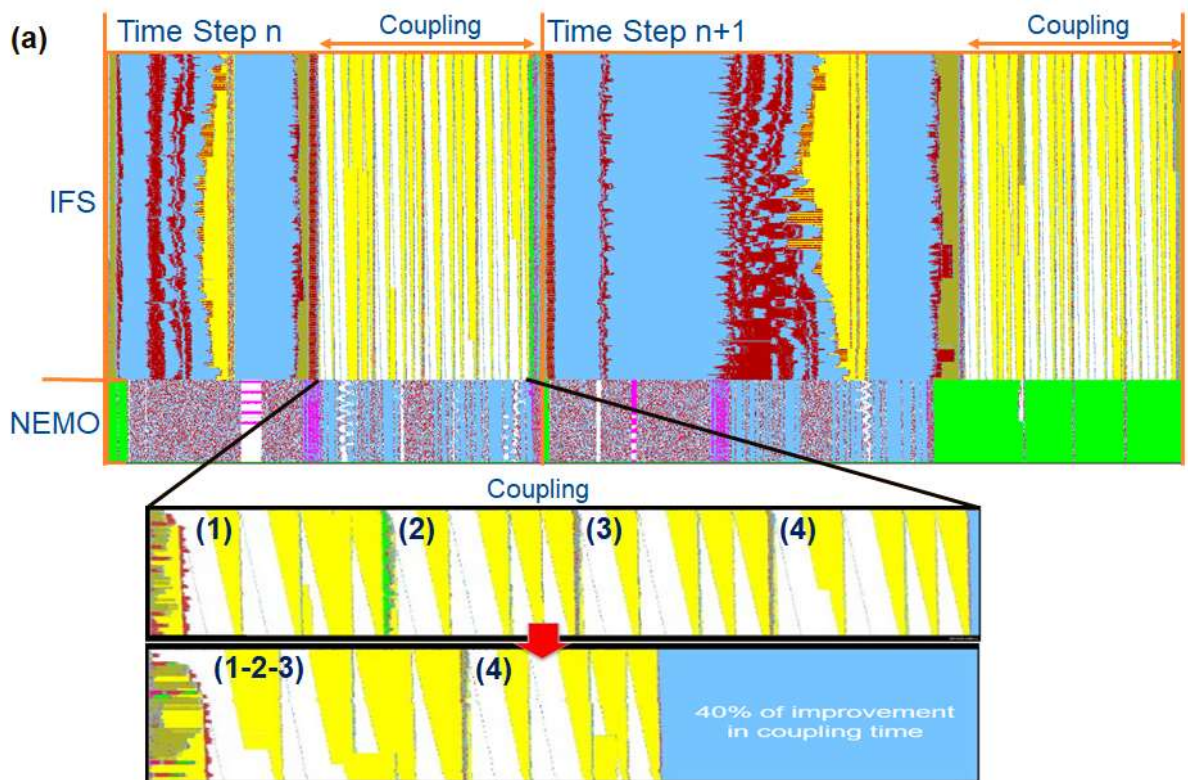
Figure 1 NEMO (red) and IFS (blue) scalability in EC-Earth3P-HR. The throughput is expressed in simulated years per day (SYPD) of wall clock time. The tests have been performed on the MareNostrum4 computer at the Barcelona Computing Centre with full output and samples of five one-month runs for each processor combination, the average of which is shown in the figure. The horizontal axis corresponds to the number of cores used.

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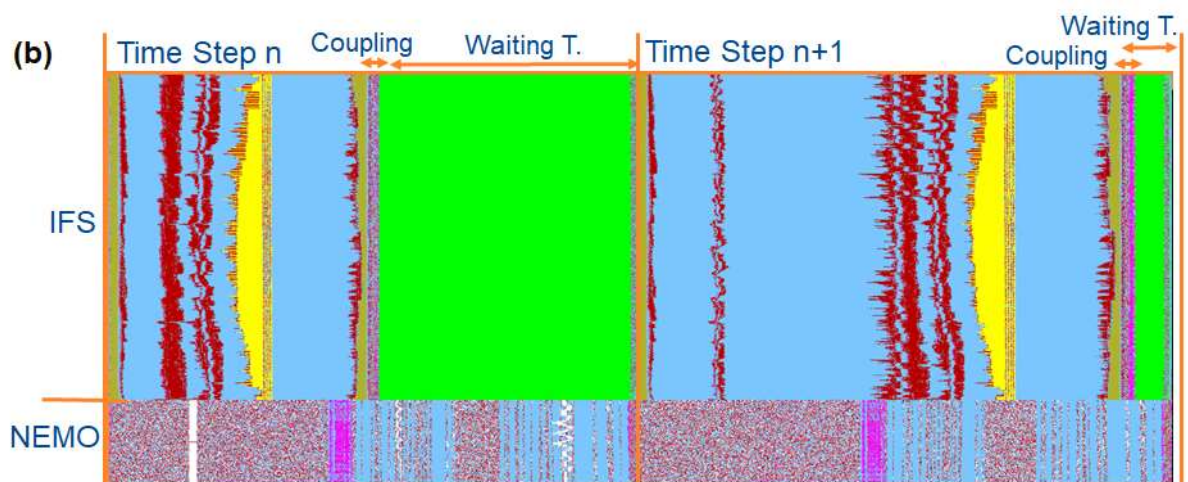


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Figure 2 As Fig. 1 but now for the scalability of the fully coupled EC-Earth3P-HR. The blue diagonal indicates perfect scalability.



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Figure 3 (a) Paraver view of the NEMO and IFS components in an EC-Earth3P-HR model execution for two time steps including the coupling process. The horizontal lines give the behaviour of the different processes (1 to 512 for IFS and 513 to 536 for NEMO) as a function of time. Each colour corresponds to a different MPI communication function. See text for explanation. (b) as (a), but now when optimization options “opt” and “gathering” for coupling are activated.

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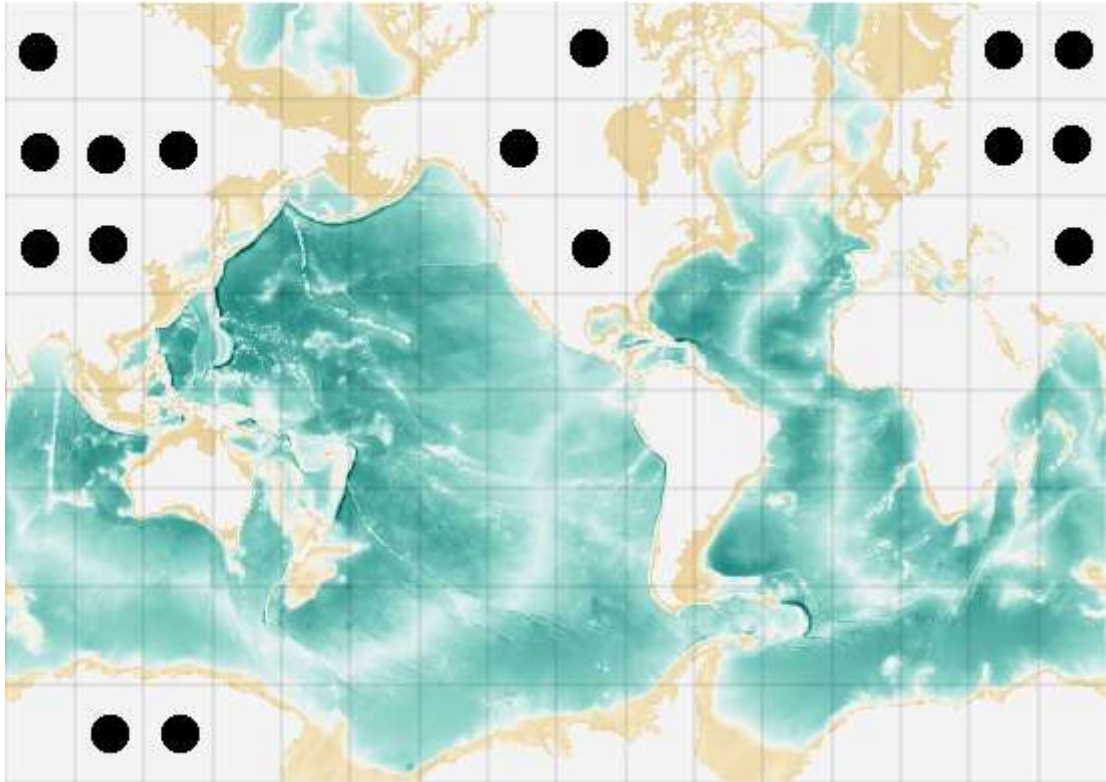


Figure 4 Domain decomposition of a tripolar grid of the ORCA family with a resolution of a of degree into 128 subdomains (16 x 8). Subdomains marked with a black dot do not contain any ocean grid point.

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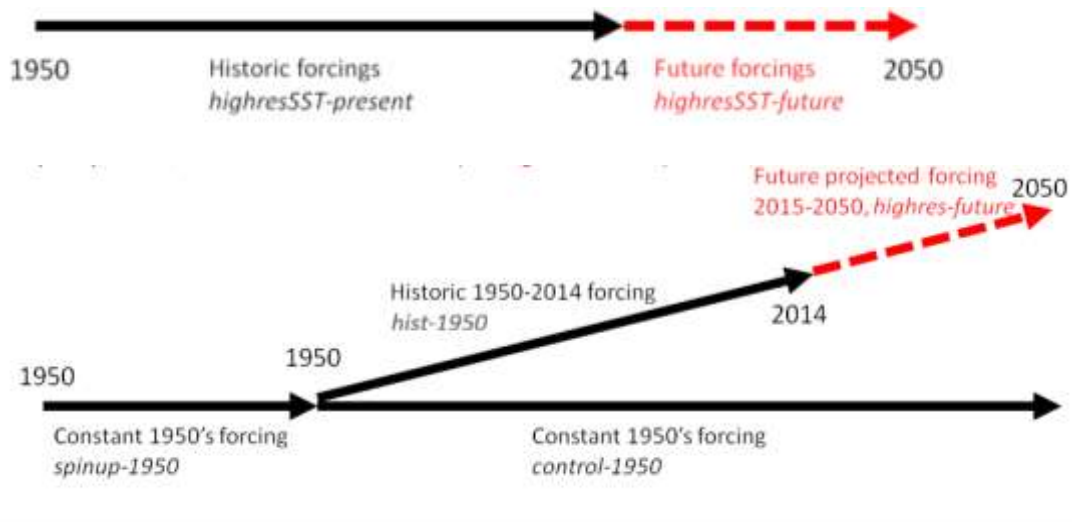
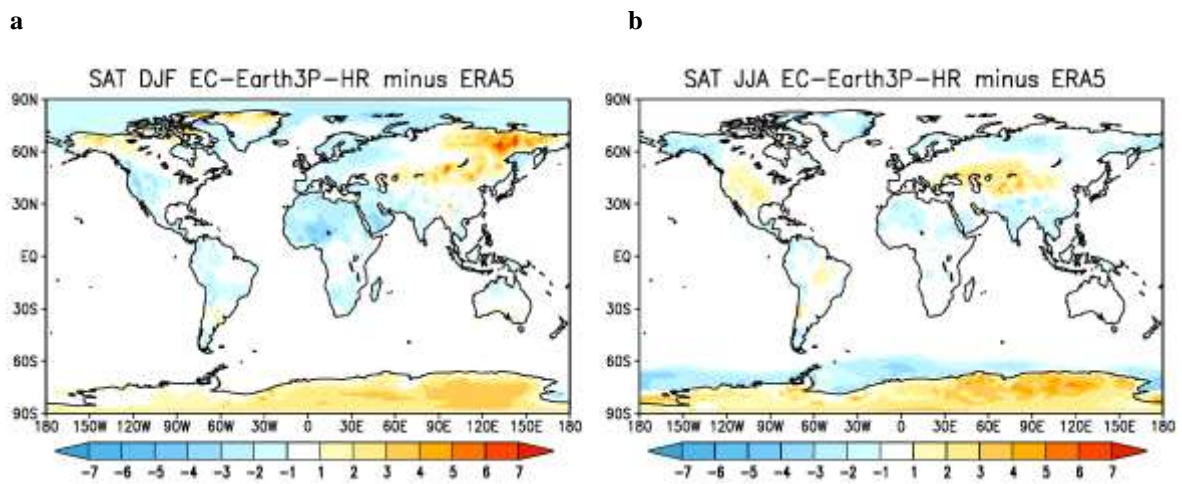


Figure 5 Schematic representation of the HighResMIP simulations.

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Figure 6 SAT: Bias [°C] of EC-Earth3P-HR with respect to ERA5 for the period 1979-2014. (a) DJF, (b) JJA. Global mean of SAT for EC-Earth3P-HR are 11.01 (DJF); 15.85 (JJA), and for ERA5 12.43 (DJF); 15.95 (JJA). RMSE of EC-Earth3P-HR with respect to ERA5 are 1.25 (DJF); 1.06 (JJA).

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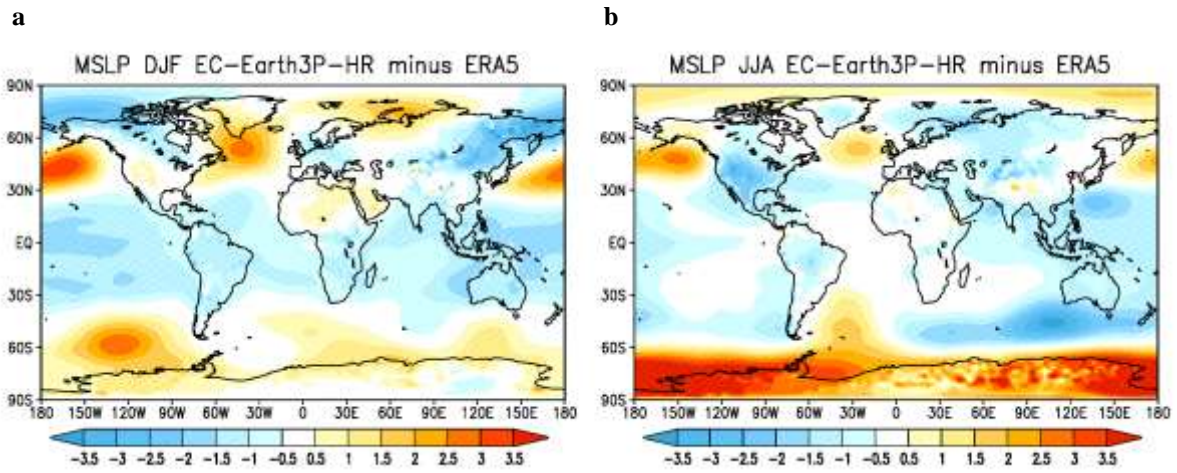


Figure 7 MSLP: Bias [hPa] of EC-Earth3P-HR with respect to ERA5 for the period 1979-2014. (a) DJF, (b) JJA.

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Global mean of MSLP for EC-Earth3P-HR are 1011.3 (DJF); 1009.4 (JJA), and for ERA5 1011.53 (DJF); 1011.24 (JJA). RMSE of EC-Earth3P-HR with respect to ERA5 are 1.11 (DJF); 1.27 (JJA).

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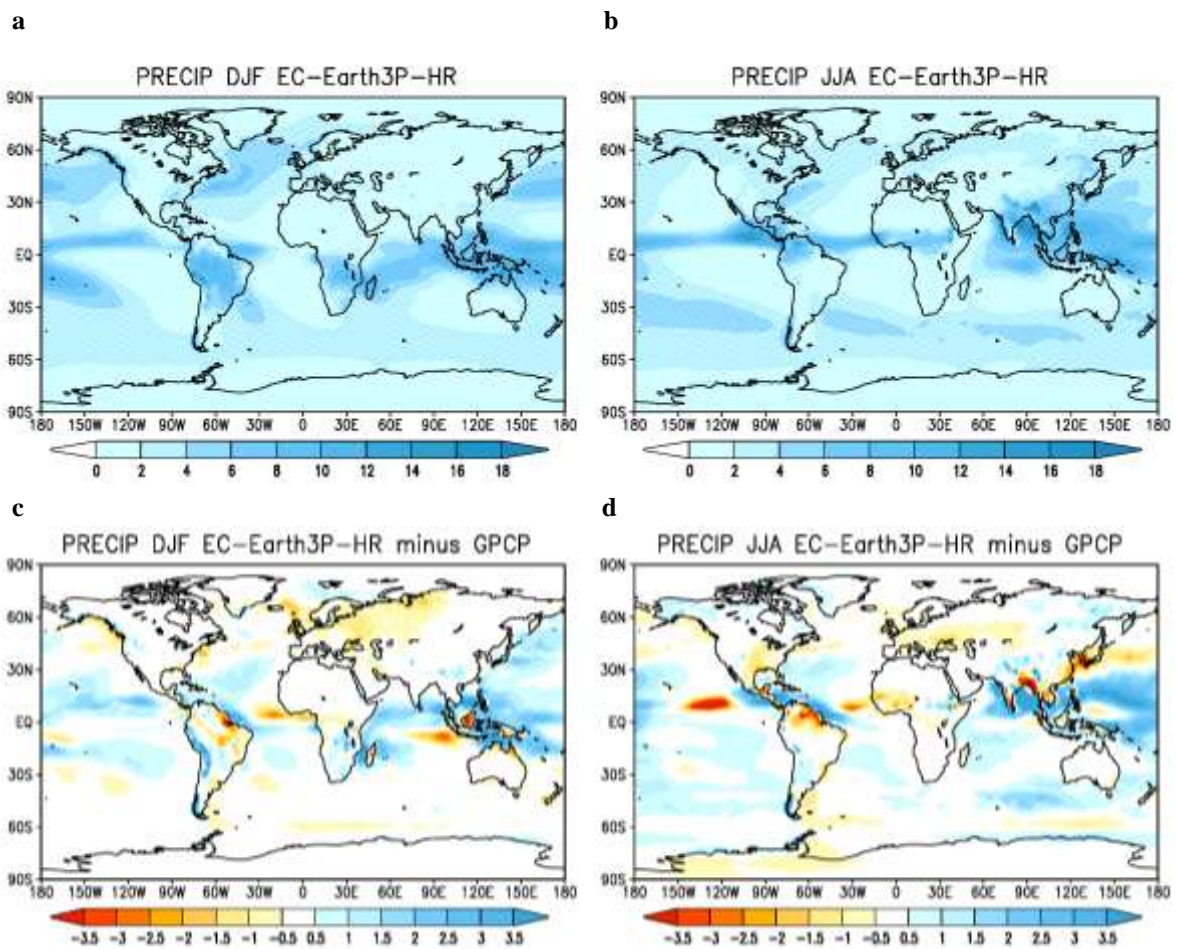


Figure 8 Precipitation and bias of EC-Earth3P-HR with respect to GPCP [mm/day] for the period 1979-2014. (a), (c) DJF, (b), (d) JJA.

1050 Global mean of precipitation for EC-Earth3P-HR are 2.91 (DJF); 3.25 (JJA), and for ERA5 2.70 (DJF); 2.71 (JJA).
RMSE of EC-Earth3P-HR with respect to ERA5 are 1.06 (DJF); 1.44 (JJA).

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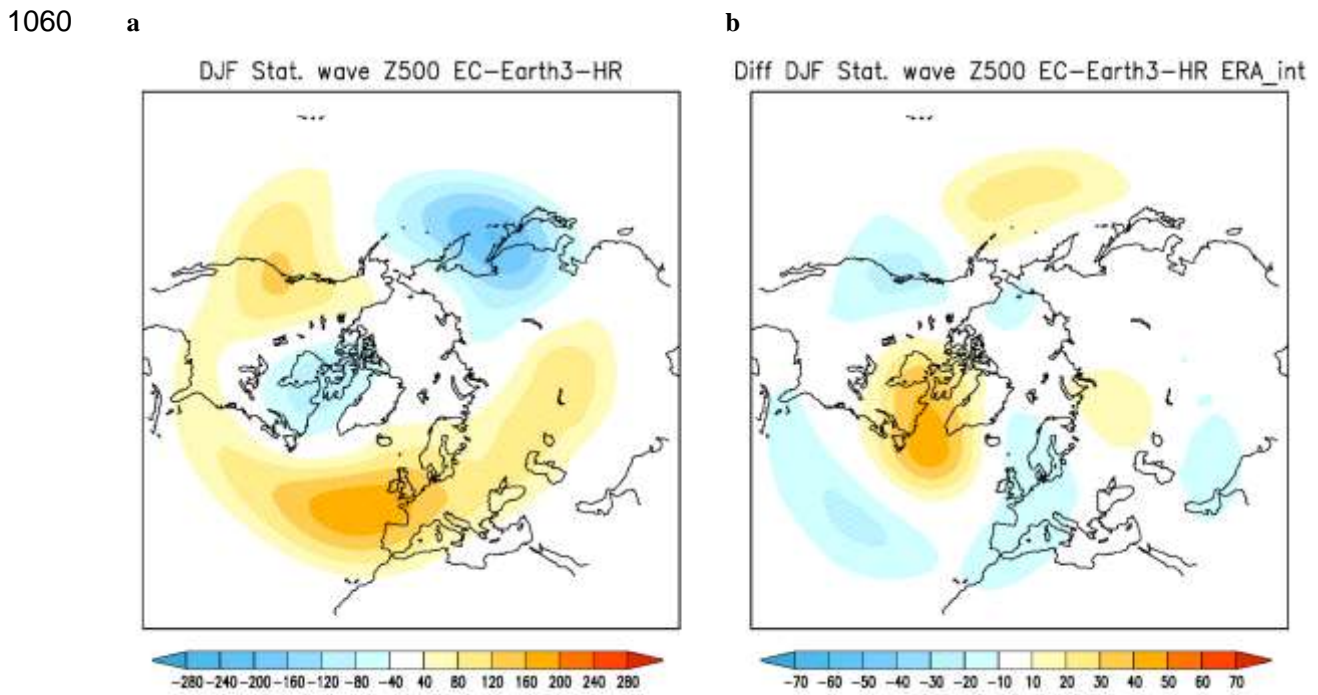
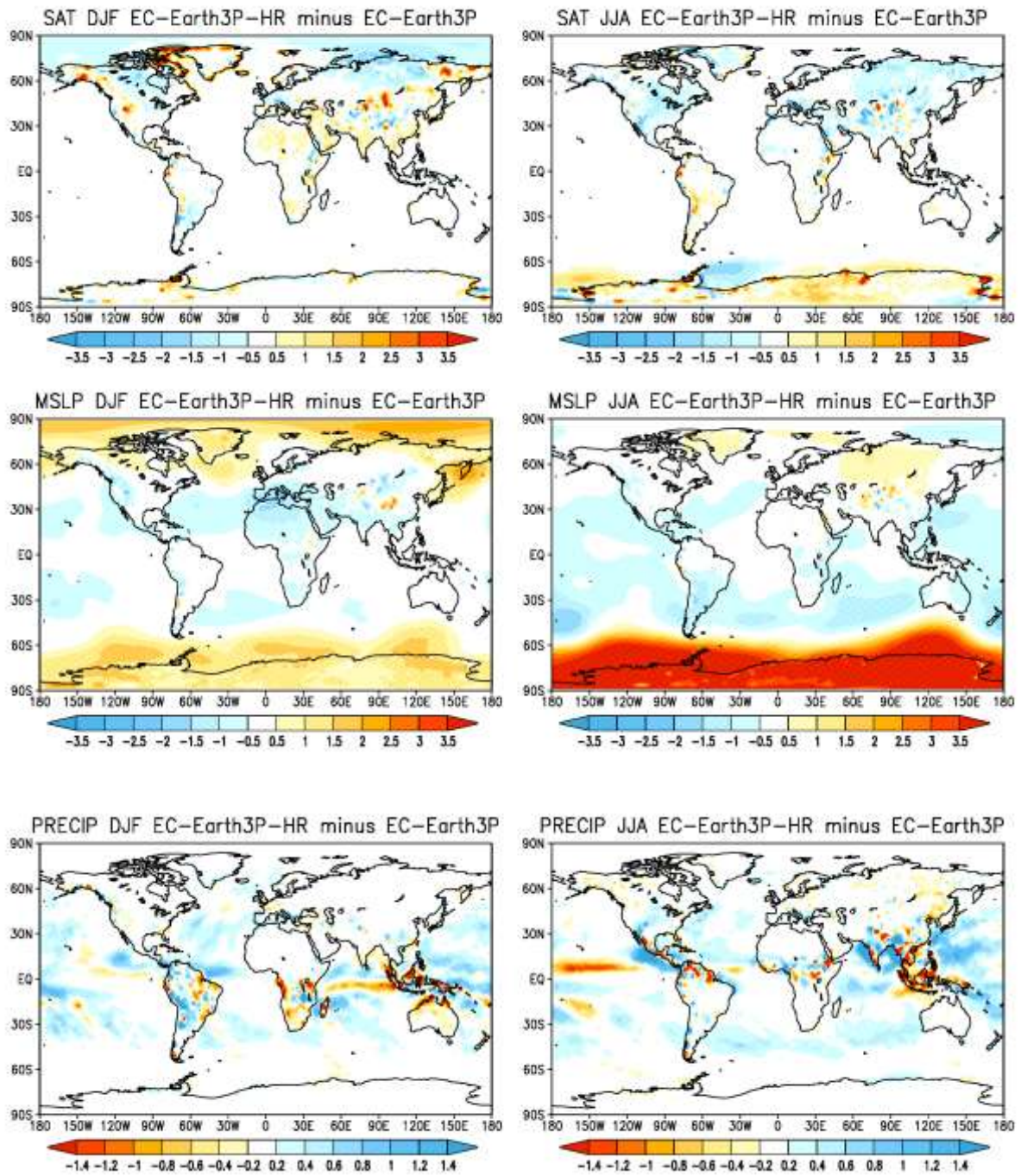


Figure 9 (a) Stationary eddy component (departure from zonal mean) of EC-Earth3P-HR of the 500-hPa geopotential height (m) in boreal winter; (b) the difference with ERA5. Note the difference in color scale between the two panels.

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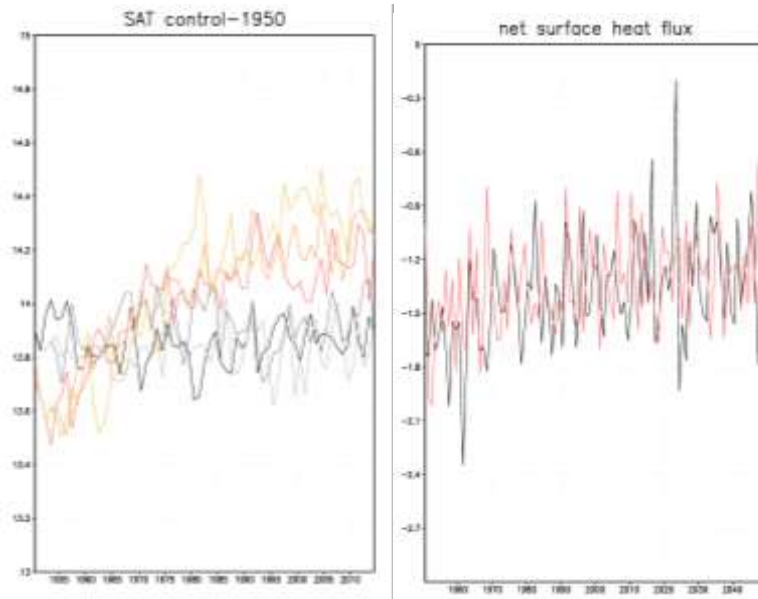


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Figure 10 Differences between EC-Earth3P-HR and EC-Earth3P for SAT [°C] (upper row), MSLP [hPa] (middle row) and precipitation [mm day⁻¹] (bottom row), for DJF (left panels) and JJA (right panels).

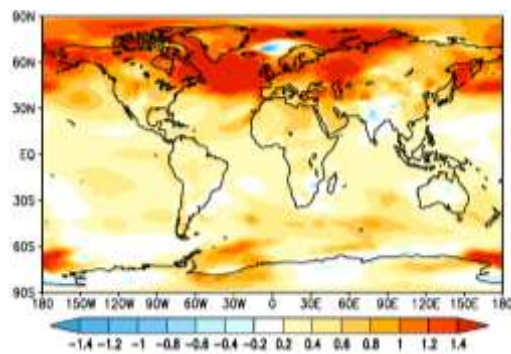
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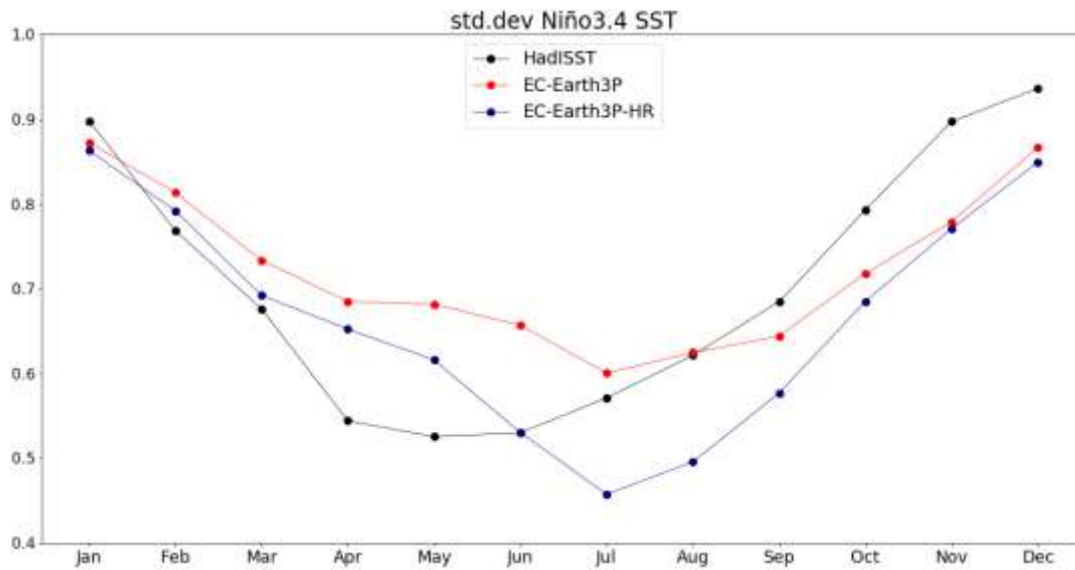


1090 **Figure 11** Left: Global mean averaged annual SAT [$^{\circ}\text{C}$] in *control-1950* for the three members of EC-Earth3P (red colors) and EC-Earth3P-HR (grey colors). Right: Global mean averaged net surface heat flux [Wm^{-2}] in *control-1950* of EC-Earth3P (red) and EC-Earth3P-HR (black), displayed only for one member (r1i1p2f1) of each model for clarity; other members display similar behavior.

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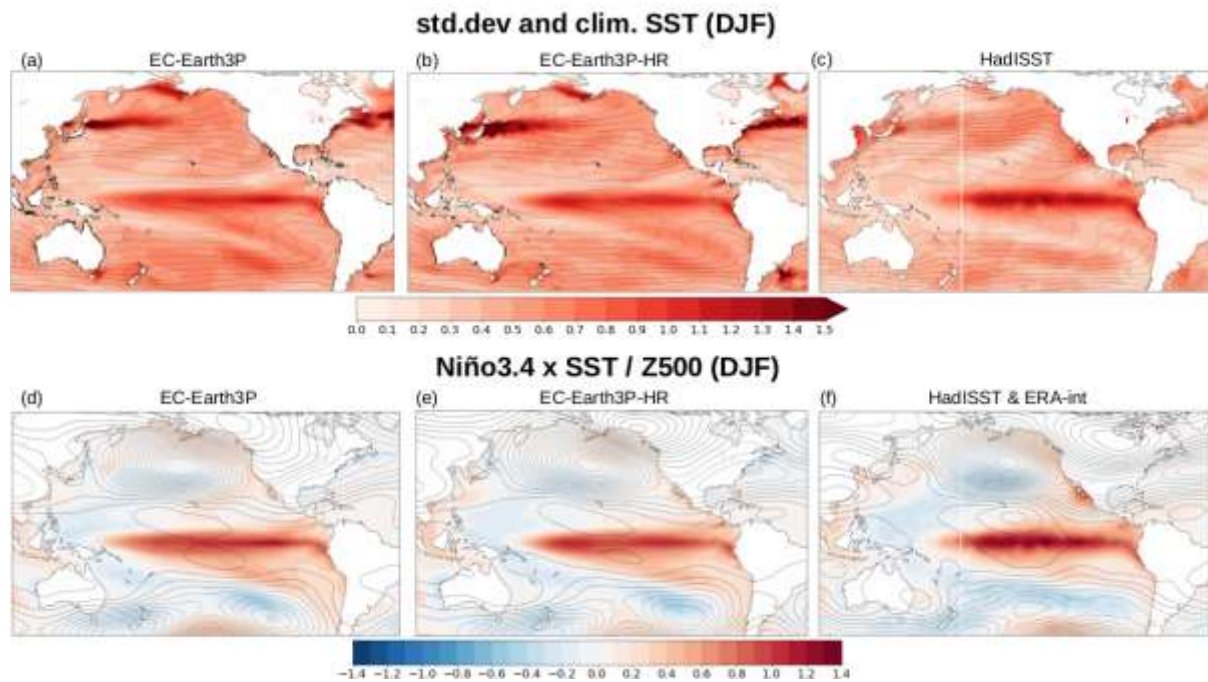
1100 **Figure 12** Ensemble mean SAT [$^{\circ}\text{C}$] of the averaged last 10 years (2040-2049) minus the averaged first 10 year (1950-1959) of the *control-1950* simulations of EC-Earth3P.



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Figure 13 Monthly standard deviation of the Niño3.4 SST index: EC-Earth3P (red) EC-Earth3P-HR (blue) from *control-1950*, and detrended HadISST over 1900-2010 (black).

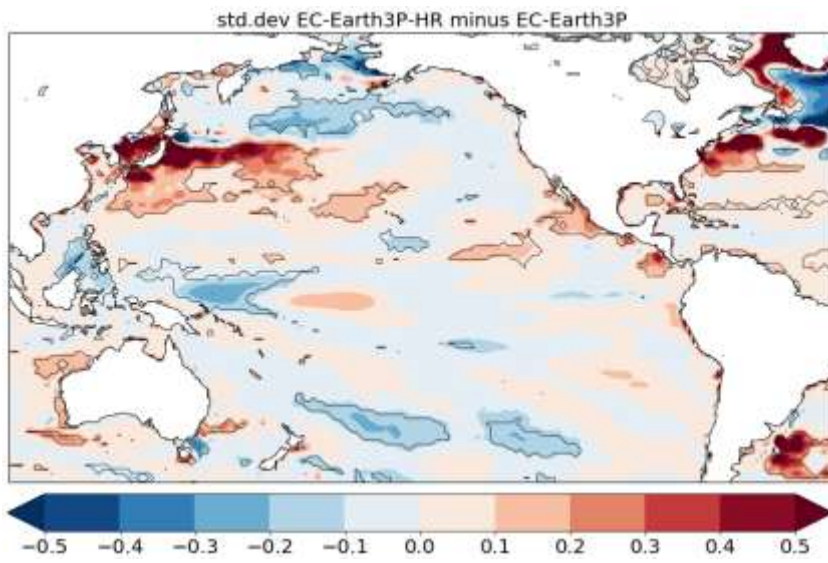
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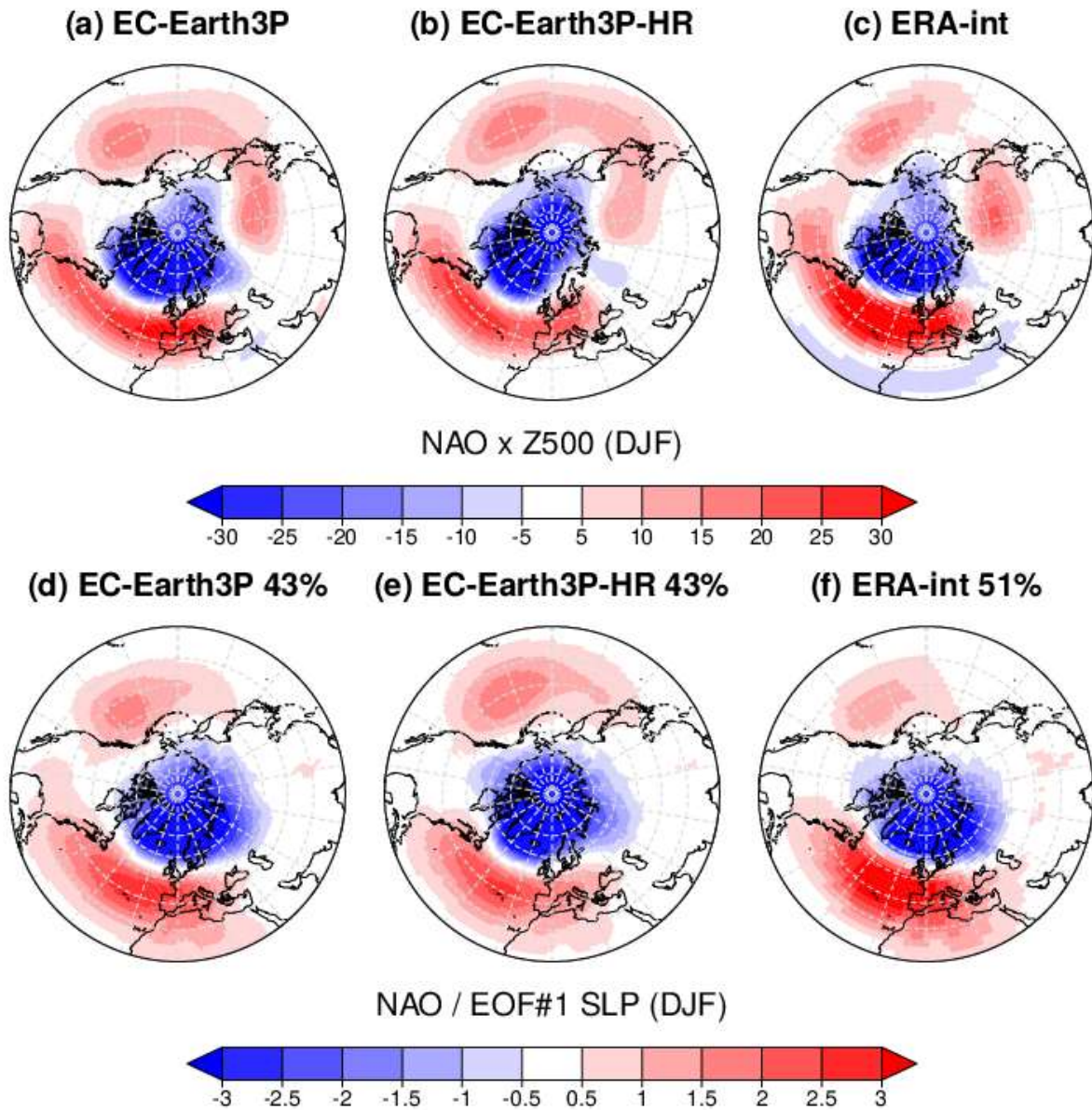
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Figure 14 Top: Boreal winter SST standard deviation from *control-1950* in EC-Earth3P (a), EC-Earth3P-HR (b), and detrended HadISST (c); overplotted with contours are the corresponding climatology (c.i. 2°C). Bottom: Regression of SST anomalies onto the Niño3.4 index from *control-1950* in EC-Earth3P (d), EC-Earth3P-HR (e), and detrended HadISST (f); overplotted with contours are the corresponding regression of 500hPa geopotential height anomalies (c.i. 2.5m), ERA-Interim in panel (f). Observational period 1979-2014.

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1125 **Figure 14 bis:** Difference in SST standard deviation between EC-Earth3P-HR (Fig. 14b) and EC-Earth3P (Fig. 14a).

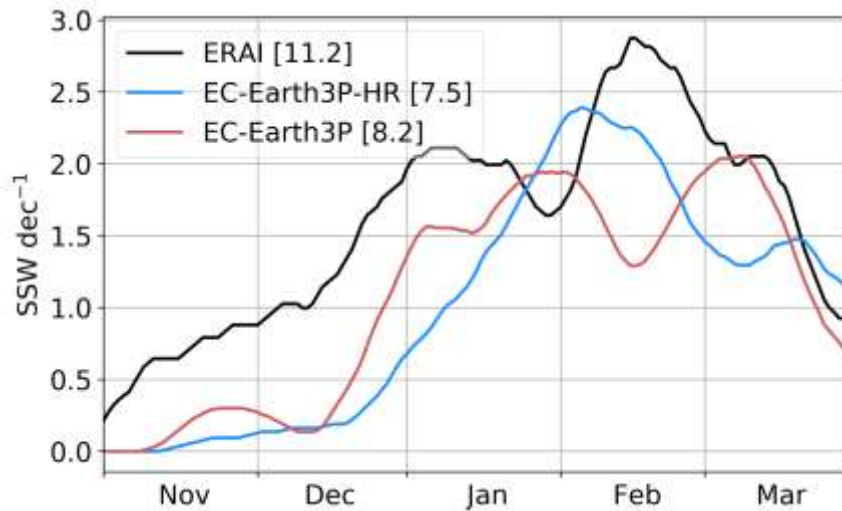


1130 **Figure 15**

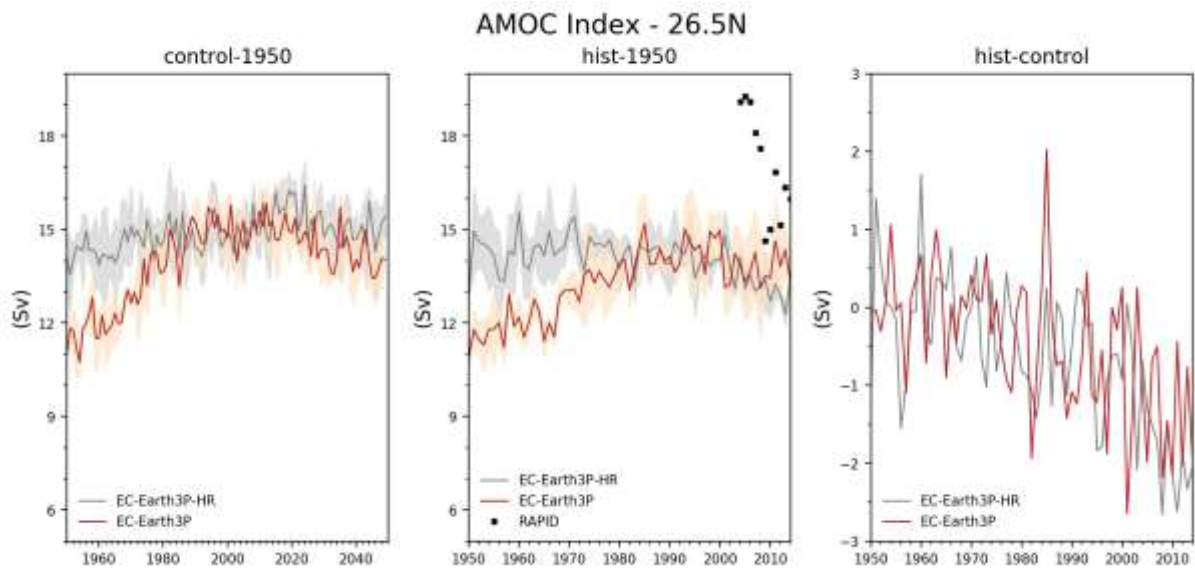
Bottom: Leading EOF of winter SLP anomalies over the North Atlantic-European region 20°N-90°N/90°W-40E from *control-1950* in EC-Earth3P (d), EC-Earth3P-HR (e), and detrended ERA-Interim (f); the corresponding fraction of explained variance is indicated in the title. Top: Regression of 500hPa geopotential height anomalies from *control-1950* in EC-Earth3P (a), EC-Earth3P-HR (b), and detrended ERA-Interim (c) onto the corresponding leading principal component, i.e. NAO index.

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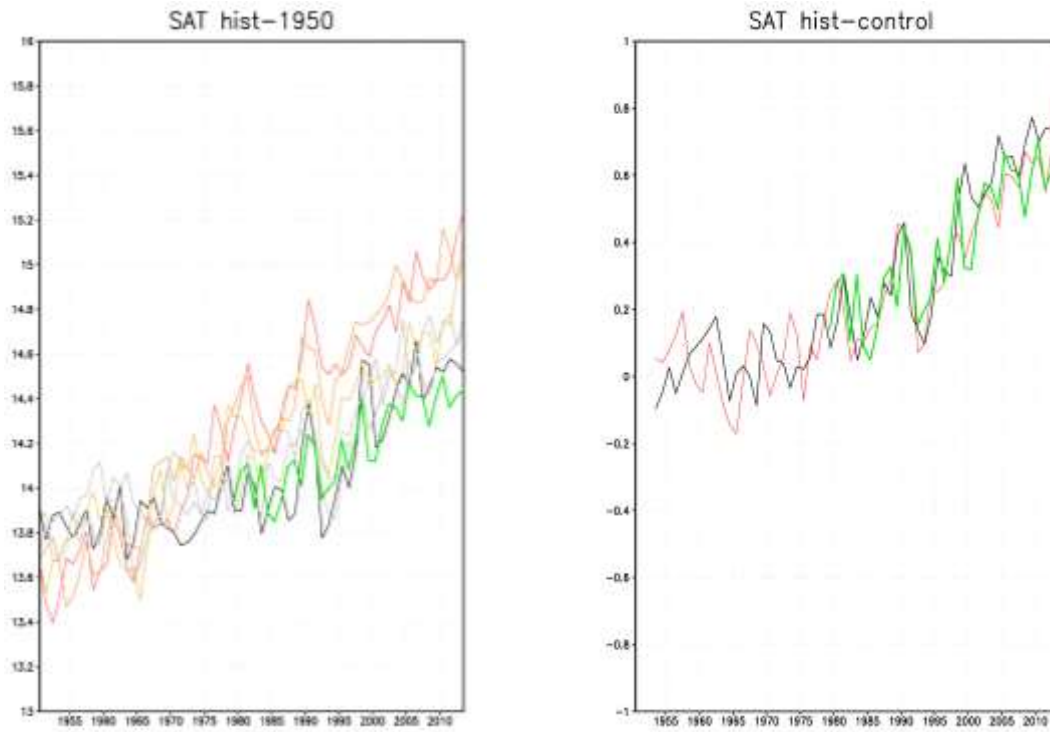


1145 **Figure 16** Seasonal distribution of SSWs per decade in a [-10, 10]-day window around the SSW date for ERA-Interim (black), EC-Earth3P (red) and EC-Earth3P-HR (blue) from *control-1950*. Time-series are smoothed with a 11-day running-mean. The total decadal frequency of SSWs is indicated in brackets.



1150 **Figure 17** Time series of the annual AMOC index for the *control-1950* (left) and *hist-1950* (middle) runs. Solid lines display the ensemble mean for the EC-Earth3P (red) and EC-EarthP-HR (black). Shaded areas represent the dispersion due to the ensemble members. Black stars in the middle panel displays values of RAPID data. Right: Mean ensemble difference between *hist-1950* and *control-1950* for Earth3P (red) and EC-Earth3P-HR (black).

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Figure 18 Global mean averaged annual SAT [$^{\circ}\text{C}$] in *hist-1950* (left) for the three members of EC-Earth3P (red colors) and EC-Earth3P-HR (grey colors). Right: Mean ensemble difference between *hist-1950* and *control-1950* for EC-Earth3P (red) and EC-Earth3P-HR (black). ERA5 is indicated by the green curves. For the right plot it is scaled so that the starting point fits with the EC-Earth curves.

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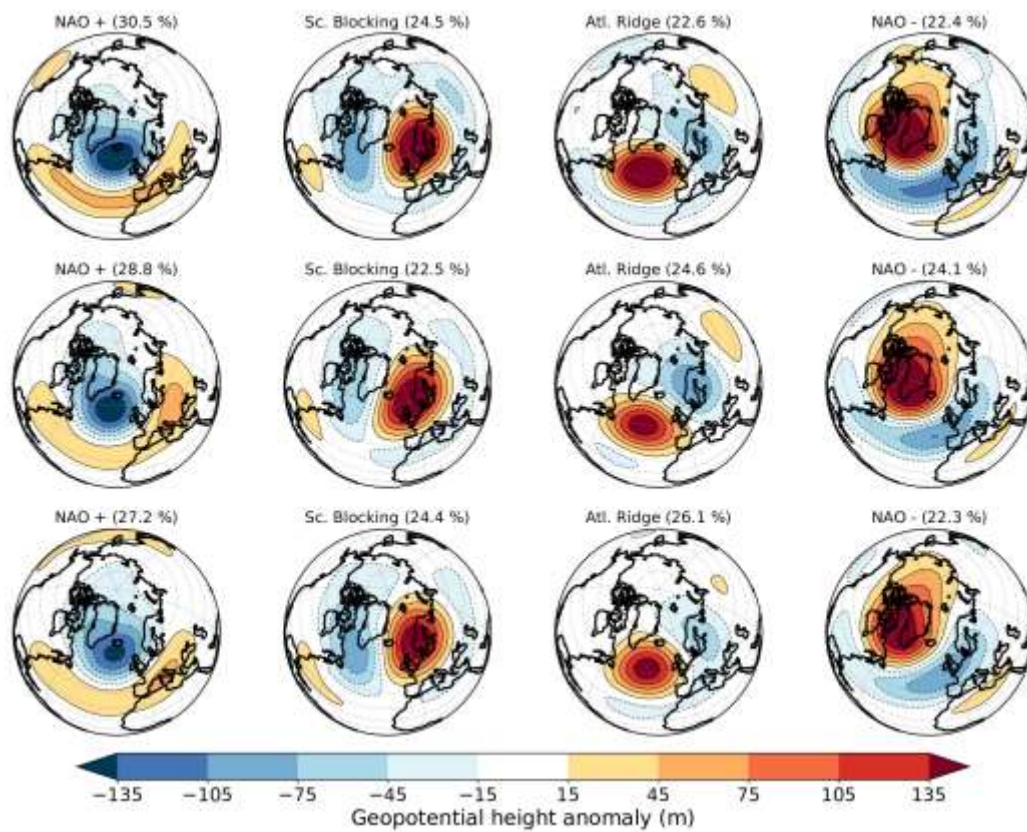


Figure 19 Observed cluster patterns for ERA (top), simulated cluster patterns in hist-1950 for EC-Earth3P (middle) and EC-Earth3P-HR (bottom). The frequency of occurrence of each regime is shown above each subplot.