



1 **Variational regional inverse modeling of reactive species emissions with PYVAR-CHIMERE**

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9

10 **Abstract**

11 Up-to-date and accurate emission inventories for air pollutants are essential for understanding their  
12 role in the formation of tropospheric ozone and particulate matter at various temporal scales, for  
13 anticipating pollution peaks and for identifying the key drivers that could help mitigate their  
14 emissions. This paper describes the Bayesian variational inverse system PYVAR-CHIMERE, which  
15 is adapted to the inversion of reactive species. Complementarily with bottom-up inventories, this  
16 system aims at updating and improving the knowledge on the high spatio-temporal variability of  
17 emissions of air pollutants and their precursors. The system is designed to use any type of  
18 observations, such as satellite observations or surface stations. The potential of PYVAR-CHIMERE  
19 is illustrated with one-day inversions of CO and NO<sub>2</sub> emissions in Europe, using the MOPITT and  
20 OMI satellite observations (for CO and for NO<sub>2</sub>, respectively).

21

22 **1. Introduction**

23 The degradation of air quality is a worldwide environmental problem: 91% of the world's population  
24 have breathed polluted air in 2016 according to the World Health Organization (WHO), resulting in  
25 4.2 millions of premature deaths every year [WHO, 2016]. The recent study of Lelieveld et al. [2019]  
26 even suggests that the health impacts attributable to outdoor air pollution are substantially higher than  
27 previously assumed (with 790,000 premature deaths in the 28 countries of the European Union against  
28 the previously estimated 500,000 [EEA, 2018]). The main regulated primary (i.e. directly emitted in  
29 the atmosphere) anthropogenic air pollutants are carbon monoxide (CO), nitrogen oxides (NO<sub>x</sub>  
30 =NO+NO<sub>2</sub>), sulfur dioxide (SO<sub>2</sub>), ammonia (NH<sub>3</sub>), volatile organic compounds (VOCs), and  
31 primary particles. These primary gaseous air pollutants are precursors of secondary (i.e. produced in  
32 the atmosphere through chemical reactions) pollutants such as ozone (O<sub>3</sub>) and Particulate Matter  
33 (PM), which are also threatening to both human health and ecosystems. Monitoring concentrations  
34 and quantifying emissions are still challenging and limit our capability to forecast air quality to warn  
35 population and to assess i) the exposure of population to air pollution and ii) the efficiency of  
36 mitigation policies.



37

38 Bottom-up (BU) inventories are built in the framework of air quality policies such as The Convention  
39 on Long-Range Transboundary Air Pollution (LRTAP) for air pollutants. Based on national annual  
40 inventories, research institutes compile gridded global or regional, monthly inventories (mainly for  
41 the US, Europe and China) with a high spatial resolution (currently regional or city scale inventories  
42 are typically finer than  $0.1^{\circ} \times 0.1^{\circ}$ ). These inventories are constructed by combining available  
43 (economic) statistics data, from different detailed activity sectors, with the most appropriate emission  
44 factors (defined as the average emission rate of a given species for a given source or process, relative  
45 to the unit of activity). It is important to note that the activity data (often statistical data) has an  
46 inherent uncertainty and its reliability may vary between countries or regions. In addition, the  
47 emission factors bear large uncertainties in their quantification [Kuenen et al., 2014; EMEP/EEA,  
48 2016; Kurokawa et al., 2013]. Moreover, these inventories are often provided at the annual or monthly  
49 scale with typical temporal profiles to build the weekly, daily and hourly variability of the emissions.  
50 The combination of uncertain activity data, emission factors and emission timing can be a large source  
51 of uncertainties, if not errors, for forecasting or analyzing air quality [Menut et al., 2012]. Finally,  
52 since updating the inventories and gathering the required data for a given year is costly in time,  
53 manpower and money, only a few institutes have offered estimates of the gaseous pollutants for each  
54 year since 2011 (i.e., EMEP updated until the year 2017, MEIC updated until the year 2017 to our  
55 knowledge). Nevertheless, using knowledge from inventories and air quality modelling, emissions  
56 could have been mitigated. For example, from 2010 to nowadays, emissions in various countries have  
57 been modified and/or regional trends have been reversed (e.g., the decrease of  $\text{NO}_x$  emissions over  
58 China since 2011 [de Foy et al., 2016]), leading to significant changes in the atmospheric  
59 composition. Consequently, the knowledge of precise and updated budgets, together with seasonal,  
60 monthly, weekly and daily variations of gaseous pollutants driven, amongst other processes, by the  
61 emissions are essential for understanding their role in the formation of tropospheric ozone and PMs  
62 at various temporal scales, for anticipating pollution peaks and for identifying the key drivers that  
63 could help mitigate these emissions.

64

65 In this context, complementary methods have been developed for estimating emissions using  
66 atmospheric observations. They operate in synergy between a chemistry-transport model (CTM) -  
67 which links the emissions to the atmospheric concentrations-, atmospheric observations of the species  
68 of interest, and statistical inversion techniques. A number of studies using inverse modeling were first  
69 carried out for long-lived species such as greenhouses gases (GHGs) (e.g., carbon dioxide  $\text{CO}_2$  or  
70 methane  $\text{CH}_4$ ) at the global or continental scales [Hein et al., 1997; Bousquet et al. 1999], using  
71 surface measurements. Later, following the development of monitoring station networks, the progress



72 of computing power, and the use of inversion techniques more appropriate to non-linear problems,  
73 these methods were applied to shorter-lived molecules such as CO. For these various applications  
74 (e.g., for CO<sub>2</sub>, CH<sub>4</sub>, CO), the quantification of sources was solved at the resolution of large regions  
75 [Pétron et al., 2002]. Finally, the growing availability and reliability of observations since the early  
76 2000s (in-situ surface data, remote sensing data such as satellite data), the improvement of the global  
77 CTMs, of the computational capacities and of the inversion techniques have increased the achievable  
78 resolution of global inversions, up to the global transport model grid cells, i.e. typically with a spatial  
79 resolution of several hundreds of square kilometers [Stavrakou and Muller, 2006; Pison et al., 2009;  
80 Fortems-Cheiney et al., 2011; Hooghiemstra et al., 2012; Yin et al., 2015; Miyazaki et al., 2017,  
81 Zheng et al., 2019].

82

83 Today, the scientific and societal issues require an up-to-date quantification of pollutant emissions at  
84 a higher spatial resolution than the global one and imply to widely use regional inverse systems.  
85 However, although they are suited to reactive species such as CO, NO<sub>x</sub>, and their very large spatial  
86 and temporal variability, they have hardly been used to quantify pollutant emissions. Some studies  
87 inferred NO<sub>x</sub> [Pison et al., 2007; Tang et al., 2013] and VOC emissions [Koohkan et al., 2013] from  
88 surface measurements. Konovalov et al. [2006, 2008, 2010], Mijling et al. [2012, 2013], Van der A.  
89 [2008], Lin et al. [2012] and Ding et al. [2017] have also shown that satellite observations are a  
90 suitable source of information to constrain the emissions of NO<sub>x</sub>. These regional inversions using  
91 satellite observations were often based on Kalman Filter (KF) schemes [Mijling et al., 2012, 2013;  
92 Van der A., 2008; Lin et al., 2012; Ding et al., 2017]. However, these inversion problems may be  
93 impacted by the non-linearity of atmospheric chemistry, for which variational methods are more  
94 suitable than KFs by design.

95

96 Here, we present the Bayesian variational inverse system PYVAR-CHIMERE, that has reached a  
97 good level of maturity (robustness of calculations, clarity, portability, modularity of the code) and  
98 that is adapted to reactive species such as CO and NO<sub>2</sub>. This adaptation for reactive species takes  
99 advantage of the previous developments for the quantification of fluxes of long-lived GHG species  
100 such as CO<sub>2</sub> [Broquet et al., 2011] and CH<sub>4</sub> [Pison et al., 2018] at the regional to the local scales,  
101 paving the way for monitoring anthropogenic emissions at high spatial resolutions. The PYVAR-  
102 CHIMERE inverse modeling system is based on the Bayesian variational assimilation code PYVAR  
103 [Chevallier et al. 2005] and on the regional state-of-the-art CTM CHIMERE, dedicated to the study  
104 of regional atmospheric pollution events [Menut et al., 2013, Mailler et al., 2017]. Variational  
105 techniques require the adjoint of the model to compute the sensitivity of simulated atmospheric  
106 concentrations to corrections of the fluxes. CHIMERE is one of the few CTMs all over the world



107 possessing its adjoint code (e.g., for global models: IMAGES [Stavrakou and Muller, 2006], TM5  
108 [Krol et al., 2008], GELKA [Belikov et al., 2016] and LMDz [Chevallier et al., 2005; Pison et al.,  
109 2009] ; for limited-area models: CMAQ [Hakami et al., 2007], EURAD-IM [Elbern et al., 2007],  
110 RAMS/CTM-4DVAR [Yumimoto et Uno, 2006]). Originally, the sequential adjoint code was  
111 developed at LMD [Menut et al., 2000; Menut et al., 2003]. Then, it has been further developed and  
112 parallelized at LISA [Pison et al., 2007] and LSCE (see Section 3).

113

114 The principle of variational atmospheric inversion and the configuration of PYVAR-CHIMERE are  
115 described in Section 2 and in Section 3, respectively. Details about the forward and adjoint codes are  
116 also given. Then, the potential of PYVAR-CHIMERE is illustrated in Section 4 with the optimization  
117 of European CO and NO<sub>2</sub> emissions, constrained by observations from the Measurement of Pollution  
118 in the Troposphere (MOPITT) instrument and from the Ozone Monitoring Instrument (OMI).

119

## 120 **2. Principle of Bayesian variational atmospheric inversion**

121 The Bayesian variational atmospheric inversion method adjusts a set of control parameters in input  
122 of the CTM, including parameters related to the emissions whose estimate is the primary target of the  
123 inversion. The control parameters may also include initial and boundary conditions for the  
124 atmospheric species of interest in the CTM. The adjustments are applied to prior values, usually taken,  
125 for the emissions, from pre-existing BU inventories. The principle is to minimize, on the one hand,  
126 the departures from the prior estimates of the control parameters (term  $J_b$  in Equation 1), which are  
127 weighted by the uncertainties in these estimates (called hereafter “prior uncertainties”), and, on the  
128 other hand, the differences between simulated and observed concentrations (term  $J_o$  in Equation 1),  
129 which are weighted by all other sources of uncertainties explaining these differences (called hereafter  
130 all together “observation errors”). In statistical terms, the inversion searches for the most probable  
131 estimate of the control parameters given their prior estimates, the observations, the CTM and the  
132 associated uncertainties. The solution, which will be called posterior estimate in the following, is  
133 found by the iterative minimization of a cost function  $J$  [Talagrand et al., 1997], defined as:

$$134 \quad J(\mathbf{x}) = \underbrace{(\mathbf{x} - \mathbf{x}_b)^T \mathbf{B}^{-1} (\mathbf{x} - \mathbf{x}_b)}_{J_b} + \underbrace{(\mathbf{H}(\mathbf{x}) - \mathbf{y})^T \mathbf{R}^{-1} (\mathbf{H}(\mathbf{x}) - \mathbf{y})}_{J_o} \quad (\text{Eq. 1})$$

135

136 The control vector  $\mathbf{x}$  contains the variables to be optimized during the inversion process (surface  
137 fluxes but also initial or boundary conditions for example, see Section 3.3).  $\mathbf{H}$  is the non-linear  
138 observation operator that projects the state vector  $\mathbf{x}$  onto the observation space. In most of the  
139 variational atmospheric inversion cases (such as those described in Section 4), the observation  
140 operator includes the CTM and an interpolation or an extraction and averaging of the simulated fields



141 (see Section 3.4). The observations in  $\mathbf{y}$  could be surface measurements and/or remote sensing data  
142 such as satellite data. The prior uncertainties and the observation errors are assumed to be centered  
143 and to have a Gaussian distribution. Consequently, the prior uncertainties are characterized by their  
144 covariance matrix  $\mathbf{B}$  and the observation errors are characterized by their covariance matrix  $\mathbf{R}$ . By  
145 definition, the observation errors combine errors in both the data and the observation operator, in  
146 particular the combination of measurement errors and errors in the conversion of satellite  
147 measurement into concentration data, errors from the CTM, representativity errors due to the  
148 comparison between point measurements and gridded models or due to the representation of the  
149 fluxes as gridded maps at a given spatial resolution, and aggregation errors associated with the control  
150 of emissions at a given spatial and/or temporal resolution that is different from (usually coarser than)  
151 that of the CTM.

152

153 For non-linear problems, as it is the case with reactive species, the minimum of  $J$  may be reached  
154 iteratively with a descent algorithm. In this case, the iterative minimization of  $J$  is based on a gradient  
155 method.  $J$  is calculated with the forward observation operator (including the CTM) and its gradient  
156 relative to the control parameters  $\mathbf{x}$  is provided by the adjoint operator (including the adjoint of the  
157 CTM):

158 
$$\text{Grad}J(\mathbf{x}) = \mathbf{B}^{-1}(\mathbf{x} - \mathbf{x}_b) + \mathbf{H}^T \mathbf{R}^{-1}(\mathbf{H}\mathbf{x} - \mathbf{y}) \text{ (Eq. 2)}$$

159 As shown in Figure 1, the minimization algorithm repeats the forward-adjoint cycle to seek an optimal  
160 solution for the control parameters.

161

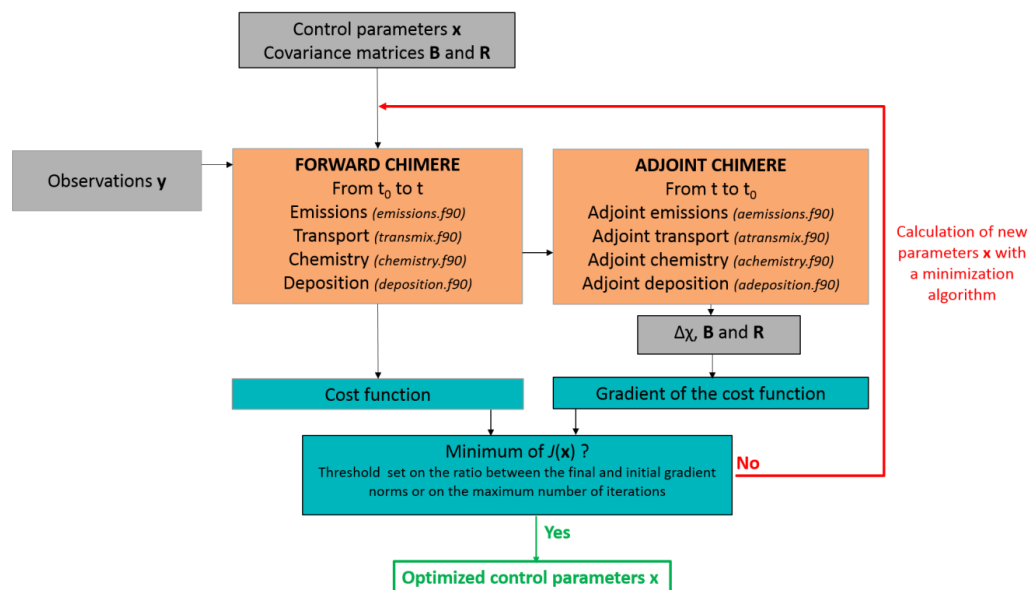
162 In PYVAR-CHIMERE, we use the M1QN3 limited memory quasi-Newton minimization algorithm  
163 [Gilbert and Lemaréchal, 1989]. As most quasi-Newtonian methods, it requires an initial  
164 regularization of  $\mathbf{x}$ , the vector to be optimized, for better efficiency. We adopt the most generally used  
165 regularization, made by optimizing  $\chi = \mathbf{B}^{1/2}(\mathbf{x} - \mathbf{x}_b)$  instead of  $\mathbf{x}$ . Although more optimal regularizations  
166 can be chosen, the minimization with  $\chi$  is preferred for its simplifying the equation to solve. This  
167 transformation translates in Equation 2 as follows:  $\text{Grad}J\chi = \chi + \mathbf{B}^{1/2}\mathbf{H}^*(\mathbf{R}^{-1}(\mathbf{H}(\mathbf{x}) - \mathbf{y}))$ . The criterion  
168 for stopping the algorithm is based on a threshold set on the ratio between the final and initial gradient  
169 norms or on the maximum number of iterations to perform. Due to the non-linearity of the problem,  
170 the minimization may reach only a local minimum.

171

172 Finally, the calculation of the posterior uncertainty is challenging in a variational inverse system.  
173 Even though the analysis error covariance matrix can be explicitly written in various analytical forms,  
174 it requires the inversion of matrices that are too large to invert given the current computational  
175 resources in our variational approach. As a trade-off between computing resources and



176 comprehensiveness, the analysis error may be evaluated by an approach based on the spread of  
 177 sensitivity tests (e.g., as in [Fortems-Cheiney et al., 2012]). It can also be estimated through a Monte  
 178 Carlo Ensemble [Chevallier et al., 2007], implemented in PYVAR.  
 179



180  
 181 **Figure 1.** Simplified scheme of the iterative minimization in PYVAR-CHIMERE. PYVAR, CHIMERE  
 182 and text sources are respectively displayed in blue, in orange and in grey.  
 183

### 184 3. The PYVAR-CHIMERE configuration

#### 185 3.1. PYVAR adapted to CHIMERE

186 The PYVAR-CHIMERE inverse modeling system is based on the Bayesian variational assimilation  
 187 code PYVAR [Chevallier et al. 2005] and on the previous inversion system based on CHIMERE  
 188 [Pison et al., 2007]. PYVAR is an ensemble of Python scripts, which deals with preparing the vectors  
 189 and the matrices for the inversion, driving the required Fortran codes of the transport model and  
 190 computes the minimization of the cost function to solve the inversion. Previously used for global  
 191 inversions with the LMDz model (e.g., Pison et al., 2009; Chevallier et al., 2010; Fortems-Cheiney  
 192 et al., 2011; Yin et al., 2015; Locatelli et al., 2015; Zheng et al., 2019), PYVAR has been adapted to  
 193 the state-of-the-art CHIMERE regional CTM. Including the elements of the previous inversion  
 194 system built around CHIMERE (coded in Fortran90) [Pison et al., 2007] in PYVAR (coded in Python,  
 195 see Section 3.5) lead to an up-to-date, more flexible and more robust system.

196

#### 197 3.2. Development of the adjoint of CHIMERE

198 To compute the sensitivity of simulated atmospheric concentrations to corrections to the fluxes, the  
 199 adjoint of CHIMERE has been developed. Originally, the sequential adjoint was coded at LISA



200 [Menut et al., 2000; Menut et al., 2003; Pison et al., 2007]. Then, it has been parallelized and further  
201 developed at LSCE and LISA, together with its tangent-linear (TL) code. The adjoint has been coded  
202 by hand line by line, following the principles formulated by Talagrand [1997]. It contains exactly the  
203 same processes as the CHIMERE forward model. PYVAR-CHIMERE is currently operational for  
204 the full module of gaseous chemistry. As a compromise between the robustness of the method for  
205 reactive species, the time required coding the adjoint and the computational cost with a full chemical  
206 scheme, the aerosols modules of CHIMERE have not been included in the adjoint of CHIMERE yet  
207 and are therefore not available in PYVAR-CHIMERE.

208

209 It should also be noted that other minor changes against the CHIMERE 2013 version [Menut et al.,  
210 2013] have been implemented, including:

211 • For the geometry, the possibility of polar domains and the use of the coordinates of the corners  
212 of the cells instead of only the centers

213 • For the transport, the non-uniform Van Leer transport scheme on the horizontal,

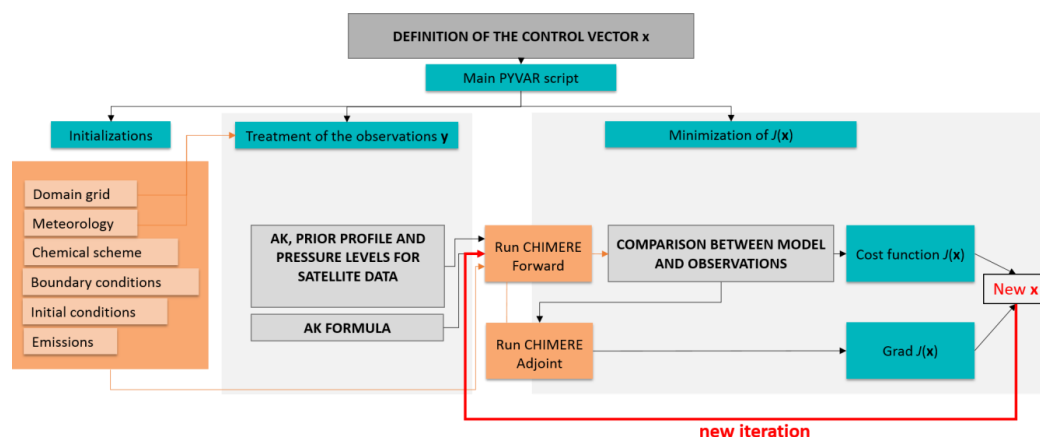
214 • For chemistry, various switches have been added to avoid going into the chemistry, deposition  
215 and wet deposition routines when no species requires them (e.g. no chemistry for methane at a  
216 regional scale).

217 It should be noted that the development and maintenance of the adjoint means that the version used  
218 is necessarily one or two versions behind the distributed CHIMERE version  
219 (<http://www.lmd.polytechnique.fr/chimere/>).

220

221 As an example, Figure 2 presents a simplified scheme of how PYVAR scripts are used to drive this  
222 version of CHIMERE for forward simulations and inversions using satellite observations. A mode is  
223 also available to test the adjoint: it runs the TL code.

224



225



226 **Figure 2.** Simplified scheme of how PYVAR scripts are used to drive CHIMERE for an inversion  
 227 using satellite observations. PYVAR, CHIMERE and text sources are respectively displayed in blue,  
 228 in orange and in grey. Note that the term AK refers to Averaging Kernel as seen in Section 3.4.

229

### 230 3.3. Definition of the control vector

231 The control vector is specified by the user in a text file. This file is formatted following Figure 3. The  
 232 inputs to constrain could be fluxes, initial concentrations, boundary conditions at the top or boundary  
 233 conditions on the four lateral sides. The spatial resolution of the inversion could be the grid-cell  
 234 resolution or one whole encompassing region. A choice of simple but efficient ways of building the  
 235 covariance matrix **B** are implemented in PYVAR-CHIMERE. The user has only to choose:

- 236 - the **B** variance coefficient, to indicate how to get the variances for these components,
- 237 - Correlation lengths *L* through time and space to define the correlations. For example, the following
- 238 formula is used to construct the distance correlation *r* between two fluxes  $x_i$  and  $x_j$ :  $r(x_i, x_j) =$

239 
$$\exp\left(\frac{-d(x_i, x_j)}{L}\right)$$

240 where *r* is the correlation between the fluxes  $x_i$  and  $x_j$ , *d* is the distance between these two fluxes in  
 241 kilometers.

Constrained species	Input to constrain	Correction type	Spatial resolution	Temporal resolution (in hours)	B variance coefficient	Time correlations in B (in hours)	Decorrelation length on land (in km)	Decorrelation length on sea (in km)
CO	Fluxes Initial conditions Boundary conditions	added	At the CHIMERE grid-cell resolution of 0.5°x0.5°	24	100 pc 10 pc 10 pc	-	-	-
NO <sub>2</sub>	Fluxes Initial conditions Boundary conditions	added	At the CHIMERE grid-cell resolution of 0.5°x0.5°	24	100 pc 10 pc 10 pc	-	-	-

242

243 **Figure 3.** Definition of the control vector and example for the illustration of Section 4.

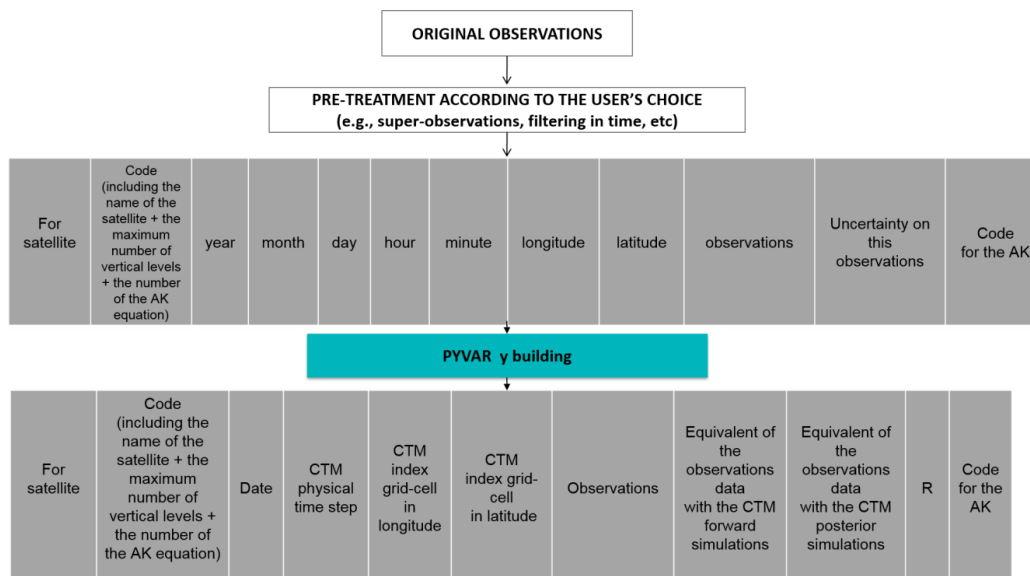
### 244 3.4 Equivalents of the observations

245 The individual data given as constraints in the system are first formatted into a text file described in  
 246 Figure 4. During forward simulations, the equivalents of the components of **y** (i.e, the equivalents of  
 247 the individual data) are calculated by PYVAR-CHIMERE. It includes the CTM and an interpolation  
 248 (see below the vertical interpolation from the model's grid to the satellite levels) or an extraction and  
 249 averaging (e.g. extracting the grid cell matching the geographical coordinates of a surface station and  
 250 averaging over one hour). As a compromise between technical issues such as the time required for  
 251 reading/writing files, the observation operator *H* that generates the equivalent of the observations by





252 the model (i.e.  $H(\mathbf{x})$ ) has been so far partly embedded in the code of CHIMERE. It makes it easier to  
 253 use finer time intervals than available in the usual hourly outputs of CHIMERE to compute the  
 254 required information (e.g., within the finer CTM physical time steps).



255  
 256 **Figure 4.** Simplified scheme of how PYVAR scripts prepares the  $y$  observations, using satellite data.  
 257 PYVAR, and text sources are respectively displayed in blue and in grey.

258  
 259 To make comparisons between simulations and satellite observations, the simulated vertical profiles  
 260 are first interpolated on the satellite’s levels (with a vertical interpolation on pressure levels). Then,  
 261 the averaging kernels (AKs), when available, are applied to represent the vertical sensitivity of the  
 262 satellite retrieval. Two types of formula, depending on the satellite observations used, have been  
 263 detailed in PYVAR-CHIMERE for the use of AKs:  $C_m = AK \cdot C_{m(o)}$  (e.g., used by the satellite  
 264 instrument OMI, see Section 4.3) or  $C_m = x_a + AK(C_{m(o)} - x_a)$  where  $C_m$  is the modeled column,  
 265 AK contains the averaging kernels,  $x_a$  is the prior profile (provided together with the AKs when  
 266 relevant) and  $C_{m(o)}$  is the vertical distribution of the original model partial columns interpolated to the  
 267 pressure grid of the averaging kernels.

268  
 269 **3.5. Numerical language**

270 The PYVAR code is in Python 2.7, the CHIMERE CTM is coded in Fortran90. The CTM requires  
 271 several numerical tools, compilers and libraries. The CTM was developed and tested using the  
 272 software versions as described in Table 1.

273



		URL	Version
Software	Python	<a href="https://www.python.org/downloads/">https://www.python.org/downloads/</a>	2.7
	Fortran compiler ifort	<a href="https://software.intel.com/en-us/fortran-compilers">https://software.intel.com/en-us/fortran-compilers</a>	Composer- xe- 2013.2.146
Libraries or packages	UnidataNetCDF	<a href="https://www.unidata.ucar.edu/">https://www.unidata.ucar.edu/</a>	3
	Open MPI	<a href="https://www.open-mpi.org/">https://www.open-mpi.org/</a>	1.10.5
	GRIB_API	<a href="https://confluence.ecmwf.int/display/GRIB/Releases">https://confluence.ecmwf.int/display/GRIB/Releases</a>	1.14
	nco	<a href="http://nco.sourceforge.net/#Source">http://nco.sourceforge.net/#Source</a>	4.6.3

274 **Table 1.** URL addresses for the development and the use of the PYVAR-CHIMERE system and its  
275 modules.

276  
277 PYVAR-CHIMERE's computation time for one node of 10 CPUs is about 3h for 1 day of inversion  
278 (with ~10 iterations) for the European domain size of 101 (longitude) x 85 (latitude) x 17 (vertical  
279 levels) used in Section 4. It is also important to note that an integration of CHIMERE's adjoint is  
280 about three times longer than a CHIMERE forward simulation. The model parallelism results from a  
281 Cartesian division of the main geographical domain into several sub-domains, each one being  
282 processed by a worker process. To configure the parallel sub-domains, the user has to specify two  
283 parameters in the model parameter file: the number of sub-domains for the zonal and meridional  
284 directions. The total number of CPUs used is therefore the product of these two numbers plus one for  
285 the master process.

286

#### 287 **4. Potential of PYVAR-CHIMERE for the inversion of CO and NO<sub>2</sub> emissions**

288 The potential of the PYVAR-CHIMERE system to invert emissions of reactive species is illustrated  
289 with the inversion of the emissions of CO and NO<sub>2</sub> over Europe. We present results for 1-day  
290 inversions. We select two different days, respectively for CO and NO<sub>2</sub>: the 7<sup>th</sup> March 2015 and the  
291 19<sup>th</sup> February 2015. These particular days have been chosen as they present a typical number of super-  
292 observations (with respectively, 1587 and 3330) during winter, when the prior emissions are high.

293

##### 294 **4.1. Set-up**

295 For both examples, CHIMERE is run over a 0.5°×0.5° regular grid (about 50x50km<sup>2</sup>) and 17 vertical  
296 layers, from the surface to 200hPa (about 12km), with 8 layers within the first two kilometers. The  
297 domain includes 101 (longitude) x 85 (latitude) grid-cells (15.5°W-35°E; 31.5°N-74°N, see Figure  
298 5). Meteorological fields are provided by ECMWF meteorological forecast [Owens and Hewson,  
299 2018]. The chemical scheme used in PYVAR-CHIMERE is MELCHIOR-2, with more than 100  
300 reactions [Derognat et al., 2003], including 22 for inorganic chemistry.



301 **4.1.1. Control vector  $\mathbf{x}$**

302 All the information required by the inverse system to run the inversion for CO or NO<sub>2</sub> emissions are  
303 listed in Table 1. The control vector  $\mathbf{x}$  has 17542 components:

- 304     ▪ the CO or NO<sub>2</sub> emissions at a 1- day and at a 0.5° ×0.5° (longitude, latitude) resolutions i.e.  
305         1 day times 101x85 grid cells = 8585 components in  $\mathbf{x}$  for each one-day-long inversion
- 306     ▪ the CO or NO<sub>2</sub> 3D initial conditions at a 0.5° ×0.5° (longitude, latitude) resolution i.e. 8585  
307         components in  $\mathbf{x}$
- 308     ▪ the CO or NO<sub>2</sub> boundary conditions at a 1- day and at a 0.5° ×0.5° (longitude, latitude)  
309         resolutions i.e. 2x101 + 2x85=372 components in  $\mathbf{x}$

310

311 The prior anthropogenic emissions for CO and NO<sub>x</sub> emissions come from the TNO inventory  
312 [Dellaert et al., 2018]. The prior anthropogenic emissions for VOCs come from the EMEP inventory.  
313 Climatological values from the LMDZ-INCA global model [Szopa et al., 2008] are used to prescribe  
314 concentrations at the lateral and top boundaries and the initial atmospheric composition in the domain.

315

316 **4.1.2. Observations  $\mathbf{y}$**

317 The observational constraints for CO emissions come from the MOPITT instrument. It has been flown  
318 onboard the NASA EOS-Terra satellite, on a low sun-synchronous orbit that crosses the equator at  
319 10:30 and 22:30 LST. The spatial resolution of its observations is about 22x22 km<sup>2</sup> at nadir. It has  
320 been operated nearly continuously since March 2000. MOPITT CO products are available in three  
321 variants: thermal-infrared TIR only, near-infrared NIR only and the multispectral TIR-NIR product,  
322 all containing total columns and retrieved profiles (expressed on a ten-level grid from the surface to  
323 100 hPa). We choose to constrain CO emissions with the MOPITT surface product for our illustration.  
324 Among the different MOPITTv8 products, we choose to work with the multispectral MOPITTv8-  
325 NIR-TIR one, as it provides the highest number of observations, with a good evaluation against in  
326 situ data from NOAA stations [Deeter et al., 2019]. The MOPITTv8-NIR-TIR surface concentrations  
327 are sub-sampled into “super-observations” in order to reduce the effect of correlated errors between  
328 neighboring observations in the inversion system: we selected the median of each subset of OMI data  
329 within each 0.5°×0.5° grid-cell and each physical time step (about 5-10 minutes). After this screening,  
330 1587 “super-observations” remain in the 1-day inversion (from 10667 raw observations). These  
331 super-observations are provided to PYVAR-CHIMERE as constraints  $\mathbf{y}$ , and treated as described in  
332 Section 3.4. Any other pre-treatment of the data (e.g., no screening, or different subsampling, etc)  
333 could have been chosen. It is important to note that the potential of MOPITT to provide information  
334 at a high temporal resolution, up to the daily scale, is hampered by the cloud coverage (see the blanks  
335 in Figure 5b).



The observational constraint on NO<sub>2</sub> emissions comes from the OMI QA4ECV tropospheric columns [Muller et al., 2016; Boersma et al., 2016, Boersma et al., 2017]. The Ozone Monitoring Instrument (OMI), a near-UV/Visible nadir solar backscatter spectrometer, was launched onboard EOS Aura in July 2004. It has been flying on a 705 km sun-synchronous orbit that crosses the Equator at 13:30 LT. Our data selection follows the criteria of the OMI QA4ECV data quality statement. As the spatial resolution of the OMI data is finer than that of the chosen CHIMERE model grid (13x24 km<sup>2</sup> against 0.5°×0.5°, respectively), the OMI tropospheric columns are sub-sampled into “super-observations” (median of the OMI data within the 0.5°×0.5° grid-cell and each physical time step and its corresponding AK).

336

337

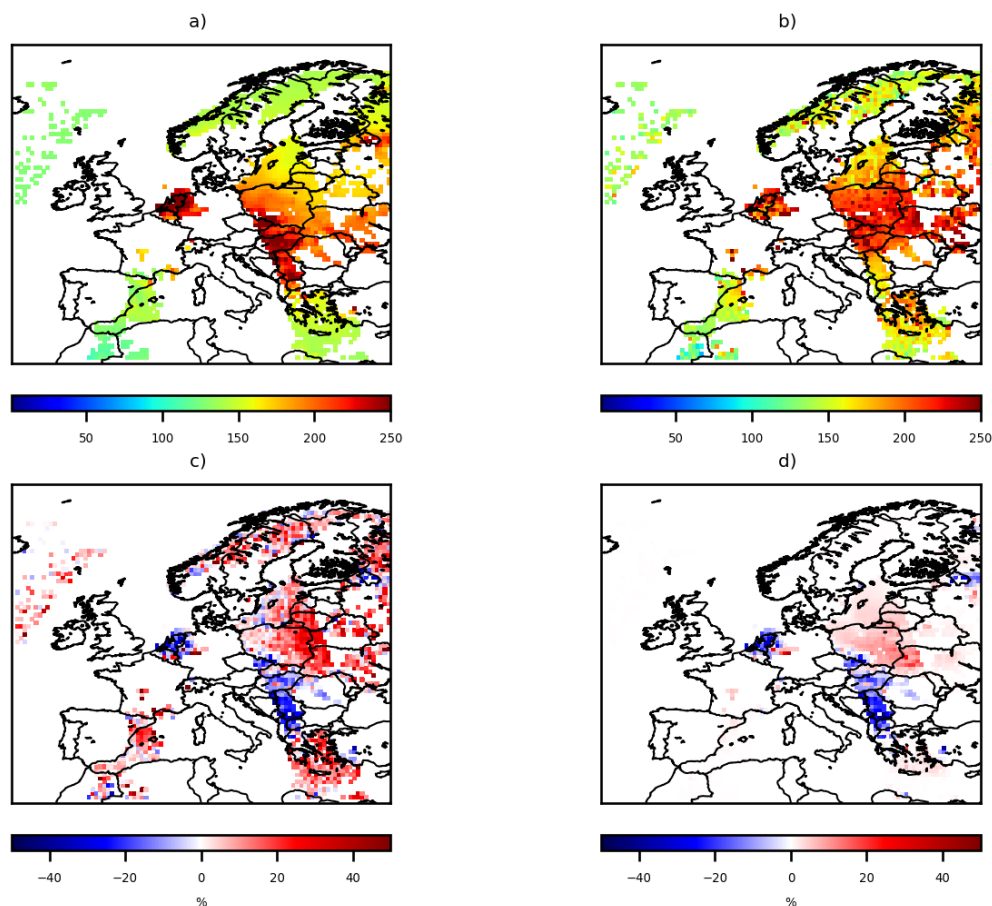
#### 4.1.3. Covariance matrices **B** and **R**

338 The covariance matrix **B** of the prior errors is defined as diagonal (i.e. only variances are taken into  
339 account). Even though total emissions are well known, large uncertainties still affect such emission  
340 inventories [Kuenen et al. 2014] at the pixel scale. Consequently, the error standard deviations  
341 assigned to the CO and NO<sub>2</sub> prior emissions in the covariance matrix **B** are set at 100%. The variance  
342 of the individual observation errors in **R** is defined as the quadratic sum of the measurement error  
343 reported in the MOPITT and the OMI data sets, and of the CTM errors (including transport errors  
344 and representativity errors) set to 20% of the retrieval values. The representativity errors could have  
345 been reduced with the choice of a finer CTM resolution (e.g., with a resolution closer to the size of  
346 the satellite pixel). Error correlations between the super-observations are neglected, so that the  
347 covariance matrix **R** of the observation errors is diagonal.

348

#### 4.2. Inversion of CO emissions

349 Large discrepancies are found between the MOPITT CO observations and the prior simulation of  
350 their equivalents by CHIMERE over Europe (Figure 5). These discrepancies might be explained by  
351 an underestimation in the BU inventory due to a general trend in emissions (if the underestimation  
352 persists throughout the year) or to an underestimation regarding particular activity sectors or the time  
353 profiles at given scales (daily, monthly). This can also be explained by uncertainties from the satellite  
354 data or from the CTM (e.g., atmospheric production, chemistry with OH).



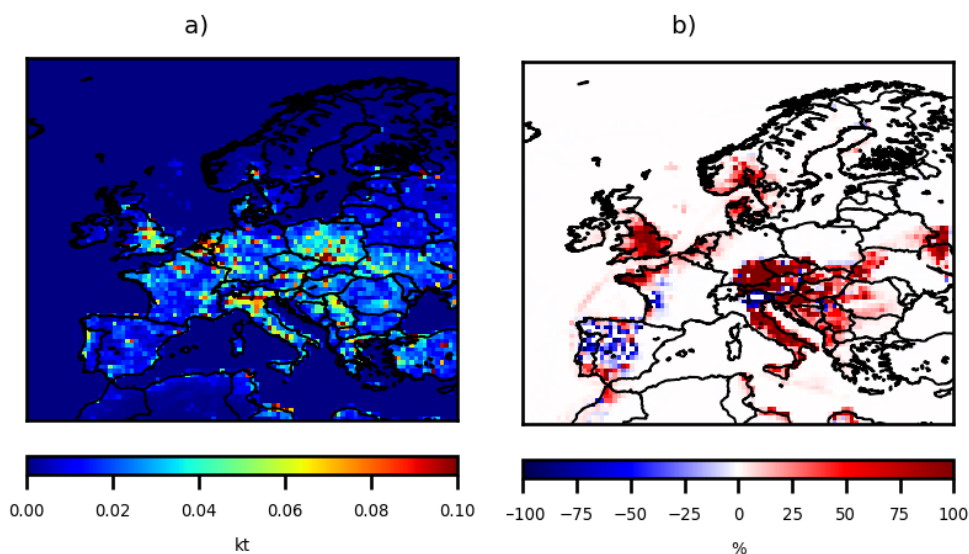
355

**Figure 5.** CO collocated surface concentrations a) simulated by CHIMERE using the prior TNO emissions and b) observed by MOPITTv8-NIR-TIR in ppbv, c) relative differences between the MOPITT observations and the CHIMERE simulated concentrations using the prior TNO emissions, and d) relative differences between the CHIMERE simulated concentrations using the posterior emissions and the CHIMERE simulated concentrations using the prior TNO emissions, in %, at the  $0.5^\circ \times 0.5^\circ$  grid-cell resolution, over Europe for the 7<sup>th</sup>, March 2015.

356 About 10 iterations are needed to reduce the norm of the gradient of  $J$  by 90% with the minimization  
357 algorithm M1QN3 and obtain the increments i.e. the corrections provided by the inversion. The prior  
358 CO emissions over Europe on the 7<sup>th</sup>, March 2015 and their increments are shown in Figure 6. As  
359 expected from the underestimation of the prior surface concentrations in Figure 5, local increments  
360 may be significant, reaching more than +50%. The posterior emissions and their uncertainties will  
361 have to be evaluated and may bring hints to the cause of the discrepancies. The analyzed



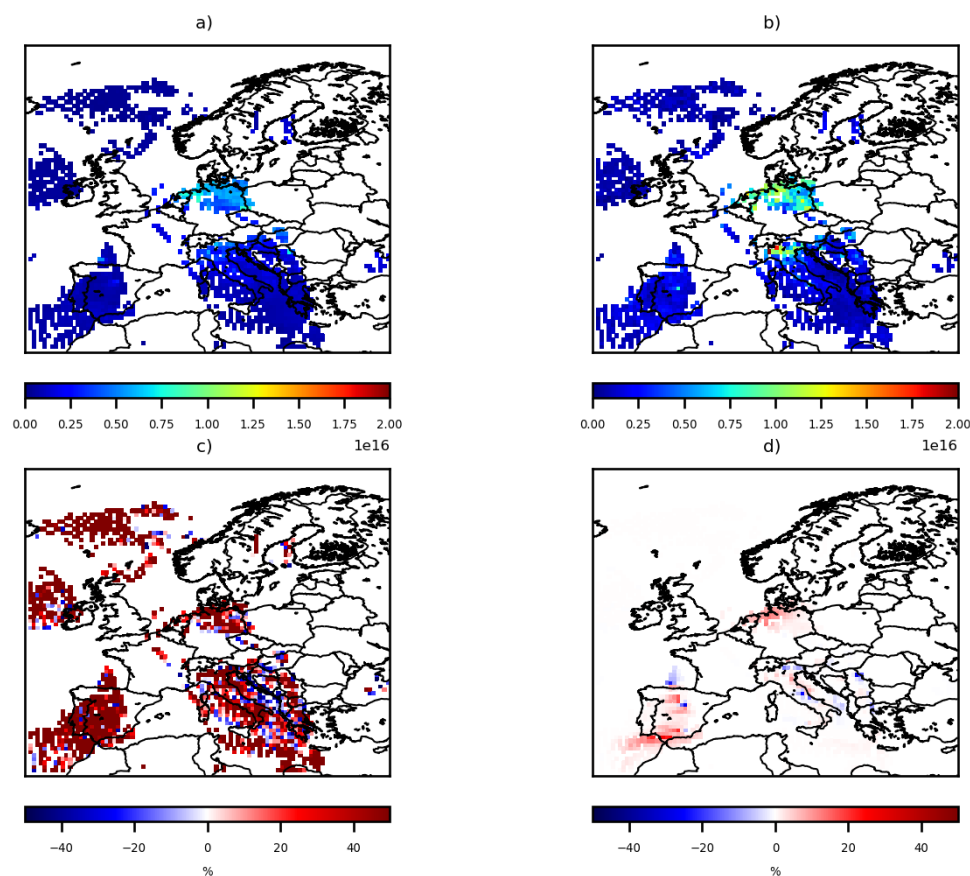
362 concentrations in Figure 5d are the concentrations simulated by CHIMERE with the posterior fluxes:  
363 as expected, the optimization of the fluxes improves the fit of the simulated concentrations to the  
364 observations.



365  
366 **Figure 6.** a) TNO CO anthropogenic prior emissions, in ktCO/grid-cell and b) increments provided  
367 by the inversion with constraints from MOPITTv8-NIR-TIR for the 7<sup>th</sup>, March 2015, in %. Note that  
368 part of Figure 6b presents blanks, as there is no observation to constrain the fluxes (see Figure 5).  
369

#### 4.3. Inversion of NO<sub>2</sub> emissions

Large discrepancies are found between the OMI NO<sub>2</sub> super-observations and the prior simulation of their equivalents by PYVAR-CHIMERE (Figure 7). Over Europe, the prior simulation strongly underestimates the tropospheric columns, and particularly over Po Valley.

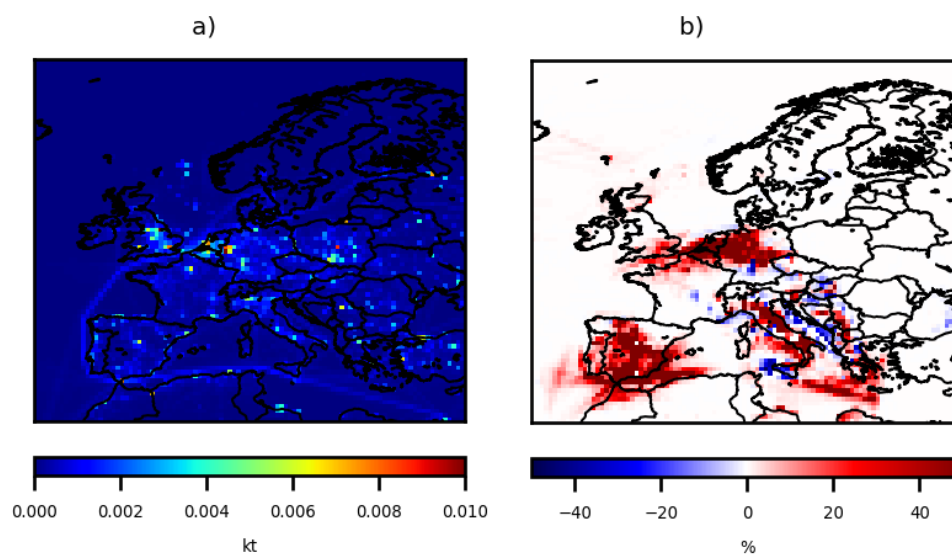


**Figure 7.**  $\text{NO}_2$  collocated tropospheric columns a) simulated by CHIMERE using the prior TNO emissions, b) observed by OMI, in  $1\text{e}16$  molec. $\text{cm}^{-2}$ , c) relative differences between the OMI observations and the CHIMERE simulated tropospheric columns using the prior TNO emissions, and d) relative differences between the CHIMERE simulated tropospheric columns using the posterior emissions and the CHIMERE simulated tropospheric columns using the prior TNO emissions, in %, at the  $0.5^\circ \times 0.5^\circ$  grid-cell resolution, for the 19<sup>th</sup> February 2015, over Europe.

370 The prior  $\text{NO}_2$  emissions and the increments obtained after inversion over Europe for the 19<sup>th</sup>,  
371 February 2015 are shown in Figure 8. As expected from the underestimation of the prior tropospheric  
372 columns in Figure 7, local increments may be large, for example over industrial areas (e.g., over the  
373 Po Valley) and over large cities (e.g, Madrid), with increments of more than +50% (Figure 8). The  
374 analyzed  $\text{NO}_2$  tropospheric columns in Figure 5c are the columns simulated by CHIMERE with the  
375  $\text{NO}_2$  posterior fluxes: as expected, the optimization of the  $\text{NO}_2$  fluxes has only a small impact on the  
376 differences between the simulated and observed  $\text{NO}_2$  columns.  $\text{NO}_x$  emissions are speciated as



377 9.2 % of NO<sub>2</sub>, 0.8 % of HONO, and 90 % of NO [Menut et al., 2013; Liu et al., 2018], following the  
378 Generation of European Emission Data for Episodes (GENEMIS) recommendations [Friedrich, 2000;  
379 Kurtenbach et al., 2001; Aumont et al., 2003]. Consequently, the NO<sub>2</sub> fluxes contribute only to about  
380 10% to the NO<sub>2</sub> tropospheric column. Further work will be done to simultaneously optimize NO and  
381 NO<sub>2</sub> fluxes, which together contribute to more than 99% to the NO<sub>2</sub> tropospheric column.



**Figure 8.** a) TNO NO<sub>2</sub> anthropogenic prior emissions, in ktNO<sub>2</sub>/grid-cell and b) increments provided by the inversion with constraints from OMI for the 19<sup>th</sup>, February 2015, in %. The configuration of the inversion is detailed in Table 1.

## 382 5. Conclusion/Discussion

383 This paper presents the Bayesian variational inverse system PYVAR-CHIMERE, which has been  
384 adapted to the inversion of reactive species such as CO and NO<sub>2</sub>, taking advantage of the previous  
385 developments for long-lived species such as CO<sub>2</sub> [Broquet et al., 2011] and CH<sub>4</sub> [Pison et al., 2018].  
386 We show the potential of PYVAR-CHIMERE, with inversions for CO and NO<sub>2</sub> illustrated over  
387 Europe for an example day. PYVAR-CHIMERE will be used to infer CO and NO<sub>x</sub> emissions over  
388 long periods, e.g. first for a whole season or year and then for the recent decade 2005-2015 in the  
389 framework of the H2020 VERIFY project over Europe, and in the framework of the ANR PoLEASIA  
390 over China, to quantify their trend and their spatio-temporal variability.

391

392 The PYVAR-CHIMERE system can handle any large number of both control parameters and  
393 observations. It will be able to cope with the dramatic increase in the number of data in the near future





394 with, for example, the high-resolution imaging (pixel of  $7 \times 3.5 \text{ km}^2$ ) of the new Sentinel-  
395 5P/TROPOMI program, launched in October 2017. These new space missions with high-resolution  
396 imaging have indeed the ambition to monitor atmospheric chemical composition for the  
397 quantification of anthropogenic emissions. Moreover, a step forward in the joint assimilation of co-  
398 emitted pollutants will soon be possible with the PYVAR-CHIMERE system and the availability of  
399 TROPOMI co-localized images of  $\text{NO}_2$  and CO for example. This should improve the consistency of  
400 the inversion results and can be used to inform inventory compilers, and subsequently improve  
401 emission inventories. Moreover, this development will help in further understanding of air quality  
402 problems and help address air quality related emissions at the national to subnational scales.

403

#### 404 **Author Contribution**

405 All authors have contributed to the manuscript writing (main authors: AFC, GB, IP and GD) and to  
406 the development of the present version of the PYVAR-CHIMERE system (main developer: IP). IP  
407 and GD have parallelized the adjoint version from Menut et al., [2000], Menut et al., [2003] and Pison  
408 et al., [2007]. IP has complemented the adjoint of new parameterizations since the CHIMERE release  
409 in 2011 and the tangent-linear model.

410

#### 411 **Code and Data Availability**

412 OMI QA4ECV  $\text{NO}_2$  product can be found here: <http://temis.nl/qa4ecv/no2.html>.  
413 MOPITTv8-NIR-TIR CO product can be found here: <ftp://15ftl01.larc.nasa.gov/MOPITT/>  
414 The CHIMERE code is available here: [www.lmd.polytechnique.fr/chimere/](http://www.lmd.polytechnique.fr/chimere/).

415

416 The associated documentation of PYVAR-CHIMERE is available on the website  
417 <https://pyvar.lscce.ipsl.fr/doku.php/3chimere:headpage>. The documentation includes a whole  
418 description of PYVAR-CHIMERE and several tutorials on how to run a first PYVAR-CHIMERE  
419 simulation or how to run an inversion.

420

#### 421 **Competing interests**

422 The authors declare that they have no conflict of interest.

423

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437

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