

1 **IGCM4: A fast, parallel and flexible intermediate climate model**

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22

23 **Abstract**

24 The IGCM4 (Intermediate Global Circulation Model version 4) is a global spectral
25 primitive equation climate model whose predecessors have extensively been used in
26 areas such as climate research, process modelling, and atmospheric dynamics. The
27 IGCM4's niche and utility lies in its speed and flexibility allied with the complexity
28 of a primitive equation climate model. Moist processes such as clouds, evaporation,
29 atmospheric radiation and soil moisture are simulated in the model, though in a
30 simplified manner compared to state-of-the-art GCMs. IGCM4 is a parallelised model,
31 enabling both very long integrations to be conducted, and the effects of higher
32 resolutions to be explored. It has also undergone changes such as alterations to the
33 cloud and surface processes, and the addition of gravity wave drag. These changes
34 have resulted in a significant improvement to the IGCM's representation of the mean
35 climate as well as its representation of stratospheric processes such as sudden
36 stratospheric warmings. The IGCM4's physical changes and climatology are
37 described in this paper.

38 **1. Introduction**

39 In order to better understand the physical processes that underpin climate and climate
40 change, it is necessary to examine not only state-of-the-art climate models, but also
41 simpler models which can have fewer degrees of freedom. In such a manner,
42 commonly referred to as the hierarchy of models approach, a more robust picture of
43 the causative mechanisms underlying climate processes can emerge. This paper
44 describes the IGCM4 (Intermediate General Circulation Model 4), which is the latest
45 incarnation of a collection of simplified climate models, collectively and usually
46 referred to as ‘Reading IGCM’ models, after the institution where much of their
47 development has taken place.

48 The rationale for such a model in the hierarchy of potential model codes is now
49 addressed. Understanding key scientific questions related to climate and climate
50 changes relies on understanding processes within the atmosphere, whose complex and
51 nonlinear nature entails the use of global circulation models. However, understanding
52 such complex processes in models is extremely challenging since unpicking processes
53 within state-of-the-art climate circulation models can be extremely difficult given
54 their complexity- especially when their computational demands are taken into account,
55 leading to limits in both integration times and data storage.

56 Having said that, it is necessary for models to be complex enough to simulate the
57 processes that are relevant to understanding a given question of interest. This is the
58 niche which intermediate circulation models such as the IGCM occupies. This niche
59 consists of models that are complex enough in terms of dynamical processes to
60 represent a wide variety of processes from monsoonal circulations to extratropical
61 storm tracks. However, their relative simplicity compared to state-of-the-art climate
62 models that are employed by the Intergovernmental Panel on Climate Change

63 (henceforth IPCC), enable process-level understanding to become more tractable
64 because of (a) computational speed enabling long integrations or large ensemble
65 members, and (b) flexibility and ease of use enabling the examination of idealised
66 scenarios. Examples where the IGCM4 might be used are e.g.; conducting
67 integrations of idealised perturbations to boundary conditions such as sea-surface
68 temperature, topography, or continental distributions; conducting ensembles of multi-
69 century integrations to collect robust statistics of small-amplitude responses to
70 particular forcings.

71 The base model which IGCM4 will be compared with is the so-called IGCM3
72 (Forster et al. 2000). The model has had many incremental updates since IGCM3, but
73 since that was the last documented model and climatology, all improvements to
74 IGCM4 are described with respect to IGCM3.

75 The IGCM has a number of configurations which are briefly described here in order
76 to clarify where IGCM4 sits in relation to the others. IGCM1 is a spectral primitive
77 equation model which can be run in global or hemispheric modes, and is based on the
78 spectral model of Hoskins and Simmons (1975). The vertical coordinate is the σ
79 terrain-following coordinate, where $\sigma = \text{pressure}/\text{surface pressure}$. Diabatic processes
80 in IGCM1 include spectral hyperdiffusion to remove noise at small scales, linear or
81 ‘Newtonian’ relaxation to a reference temperature state, and linear or ‘Rayleigh’
82 friction at any number of model layers. Examples of research conducted with this
83 configuration are studies of baroclinic lifecycles on Earth (Hoskins and Simmons
84 1975, James and Gray 1986, Thorncroft et al. 1993) and Mars (Collins and James
85 1995), as well as studies of the stationary circulation on Earth (Valdes and Hoskins
86 1991), Mars (Joshi et al. 1994) and other planets (Joshi et al 1997).

87 In IGCM2, the linear diabatic processes in IGCM1 are replaced by more realistic
88 nonlinear diffusive processes. Radiative processes are parameterised simply using a
89 prescribed surface temperature and a constant cooling rate of 1.25 Kday^{-1} representing
90 infra-red radiation to space. The effects of moisture are included in IGCM2,
91 necessitating the inclusion of evaporation, parameterisation of deep and shallow
92 convection, and the potential for moisture transport. Such a configuration represents
93 moist processes allowing the study of tropical regions, and has accordingly been used
94 in studies of mesoscale tropical dynamics and circulation (Cornforth et al. 2009).

95 IGCM3 is a full climate model in which the prescribed surface can be replaced by one
96 or both of a two-level interactive land surface, and a slab or ‘q-flux’ ocean model. The
97 constant radiative cooling is replaced by a radiative scheme which calculates clear sky
98 fluxes in 2 visible bands and 6 infra-red bands, and accounts for the radiative effects
99 of clouds. This model is described fully in an appendix to Forster et al. (2000). This
100 configuration has been used in many studies of tropospheric climate (Forster et al
101 2000, Joshi et al. 2003) and stratospheric climate (Rosier and Shine 2000, Winter and
102 Bourqui 2011 a,b). A coupled ocean-atmosphere model (FORTE) has been created in
103 the past by coupling the IGCM3 to the MOMA ocean model (e.g. Sinha et al. 2012).
104 A similar process is underway for IGCM4, and the resulting coupled model is the
105 subject of an accompanying paper.

106 We now set out the climatology of the new IGCM4 model in addition to changes
107 since the last published detailed version IGCM3. Section 2 details changes since
108 IGCM3, section 3 details the new model climatology, and section 4 details the
109 climatic performance of the IGCM4.

110

111 **2. Model changes from IGCM3**

112 **2.1 IGCM4 Configurations**

113 IGCM4 exists in two standard configurations: a spectral truncation of T42 (having a
114 128 x 64 horizontal grid) and 20 layers in the vertical, denoted T42L20, which is the
115 standard configuration for studies of the troposphere and climate, and T42L35, which
116 enables study of the stratosphere on climate. In addition, a configuration of T170L20,
117 which enables study of mesoscale phenomena such as weather fronts and tropical
118 waves, is also under development, but its description is beyond the scope of this paper.
119 The L20 and L35 configurations reach from the surface to 50 hPa and 0.1 hPa
120 respectively, and are shown in Figure 1. The lowest 19 model layers in each
121 configuration have exactly the same values, so that only the stratosphere is different,
122 enabling more traceability when comparing different model configurations.

123 The spectral code is parallelised using a so-called 2D decomposition (Foster and
124 Worley 1997, Kanamitsu et al. 2005). In a 2D decomposition, two of the three
125 dimensions are divided across the processors, and so there is a column and row of
126 processors, with the columns divided across one dimension and the rows across
127 another. Compared with a 1D decomposition, a 2D decomposition increases the
128 number of transpositions that need to be made to go from spectral-space to grid-space
129 and back again. However the advantage is that each transposition is only amongst
130 processor elements (henceforth PEs) either on the same column or the same row. Any
131 transposition for 1D decomposition requires all the PEs to communicate with one
132 another, which increases the size of buffers passed between PEs, communication
133 latency, and slows down the model. Han and Juang (2004) found that a 2D
134 decomposition is about twice as fast as a 1D decomposition. More details on the
135 decomposition are given on the IGCM website (Stringer 2012)

136 The model's performance on a parallel cluster using an intel compiler and MPI
137 parallelisation libraries is as follows: T42L35: ~ 75 model years/day on 32 processors;
138 96 timesteps per day); T42L20: ~ 200 model years/day on 32 processors; 72 timesteps
139 per day).

140

141 **2.2 Surface and boundary layer processes**

142 Over land, each grid point has a land-surface type based on present-day observations:
143 there are 8 types (ice, inland water, forest, grassland, agriculture, tundra, swamp,
144 desert). Each land-surface type has its own value for snow-free albedo A , snow-
145 covered albedo S , the height (in metres) at which total albedo reaches $(A+S)/2$, and
146 roughness length. The values of these quantities for each surface type are shown in
147 Table 1.

148 Whenever snowmelt occurs in the model, the snowmelt moistens soil so that the soil
149 water is $2/3$ of the saturated value. This is a very simple parameterisation of snowmelt
150 percolating through soil and helps to alleviate warm biases in late spring and summer
151 in Eastern Eurasia, consistent with more complex GCMs such as HadGEM2 (Martin
152 et al 2010).

153 A maximum effective depth for snow of 15m exists to prevent slow drifts in heat
154 capacity and hence temperature and energy balance, since there is no physics in the
155 IGCM4 to represent the melting of ice fields at their bases. In addition, the 'land ice'
156 surface type has a fixed snow depth, so that points diagnosed as 'ice' are not subject to
157 slowly emerging model biases in temperature appearing because of snow depth
158 slowly being eroding away over decades. At present, these fixed land-ice points are
159 set to be Antarctica and Greenland.

160 The effect of sea-ice in IGCM is implemented by assuming a linear change from 0°C
161 to -2°C in these surface properties: roughness, albedo and heat capacity. This replaces
162 the sudden change of surface properties at -2°C, which is unrealistic given partial ice
163 cover in most oceans, and also removes a bias in that while sea-ice forms from saline
164 water at -2°C, it melts at 0°C, since ice is mostly composed of fresh water. A
165 combination of ice and open water is therefore desirable between -2°C and 0°C.

166 The amount that surface heat fluxes can be amplified by convectively unstable
167 conditions above their values at neutral or zero stability has been limited to 4.0. This
168 value has been chosen to limit latent heat fluxes over the ocean and sensible heat
169 fluxes over the land to better match observations, although it is still a simplification of
170 more complex schemes that involve the Richardson number (e.g. Louis 1979), since it
171 is entirely stability-based.

172

173 **2.3 Radiation, convection, clouds and aerosol**

174 The NIKOSRAD radiation scheme in IGCM3 (Forster et al. 2000) has been replaced
175 with a modified version of the Morcrette radiation scheme (Zhong and Haigh 1995)
176 which was originally written for the ECMWF model. This is because the NIKOSRAD
177 scheme was found to produce $2\Delta z$ oscillations under certain conditions in the
178 stratosphere. A transitional version of IGCM3, called IGCM3.1, has existed with the
179 Morcrette radiation scheme for some time, and many climatic (e.g. Bell et al. 2009,
180 Cnossen et al. 2011) and climate-chemistry (e.g. Highwood and Stevenson 2003,
181 Taylor and Bourqui 2005) studies have been conducted with it. The Morcrette
182 radiation scheme has a representation of O₃ absorption of UV between 0.12µm and

183 0.25 μ m, 2 visible bands (0.25-0.68 μ m, 0.68-4 μ m), and 5 infra-red (henceforth IR)
184 bands.

185 The radiatively active species in the IGCM are H₂O, CO₂, CH₄, O₃, N₂O, CFC-11 and
186 CFC-12. H₂O is advected self-consistently in the model, but prescribed above a
187 seasonally varying climatological tropopause. O₃ is specified from a zonally averaged
188 climatology (Li and Shine 1995), which is then interpolated to model levels. All other
189 gases are assumed to be well-mixed throughout the GCM domain, and are easily
190 changed via a namelist.

191 The solar constant in IGCM4 is 1365 Wm⁻², which is more consistent with
192 observations than the older value of 1376 Wm⁻² in IGCM3 and IGCM3.1. The ocean
193 albedo A_o varies with latitude φ in this manner:

$$194 \quad A_o = 0.45 - 0.30 \cos \varphi. \quad (1)$$

195 This is a simple parameterisation of the effects of aerosols and solar zenith angle on
196 albedo based on observations so that at the equator $A_o = 0.15$, increasing to 0.3 at
197 60°S/N.

198 The convection scheme in the IGCM4 is identical to that described in Forster et al
199 (2000) and is based on the scheme of Betts (1986), with separate adjustment processes
200 for shallow and deep convection; the adjustment process for deep convection takes
201 place over 3 hours as in Forster et al (2000). Rainout of shallow convective
202 precipitation is now allowed in IGCM4 over a timescale of 6 hours. This rainout helps
203 to slow down the Hadley circulation, whilst removing some of the shallow convective
204 cloud that occurs over subtropical regions. Stratiform precipitation is as in Forster et
205 al (2000): grid-scale supersaturation is removed. Above a gridpoint relative humidity

206 (henceforth RH) of 0.8, clouds are formed whose fraction F is given by $F = ((RH-$
207 $0.8)/0.2)^2$. No cloud can form in the very lowest model layer.

208 The clouds have been tuned to better match observations of outgoing infra-red
209 radiation and downward surface solar radiation: the cloud base fraction for deep
210 convective cloud is 4 times the fraction at all other levels, which is consistent with
211 observed convective cloud profiles (Slingo 1987). A version of the Kawai and Inoue
212 (2006) parameterisation for marine stratocumulus cloud has also been implemented in
213 IGCM4. This diagnoses low cloud at ocean points depending on the stability of the
214 lowest two model sigma half layers, i.e. between the surface and layer 1, and layer 1
215 and layer 2, and deposits cloud in the second-to-lowest model layer if diagnosed.

216 Aerosols are not in the standard IGCM4: their effect on surface temperatures have
217 been parameterised by slightly raising the albedo of land and ocean by 0.05. This is
218 because even CMIP5 GCMs have trouble accurately representing the forcing due to
219 different types of aerosol. In addition, even the aerosol scheme in the IGCM only
220 deals with the direct effect, and not the different indirect effects such as cloud lifetime
221 and particle size that are also present in reality. However, both specific case studies of
222 tropospheric and stratospheric aerosols have been studied using IGCM3.1 (Highwood
223 and Stevenson 2003, Ferraro et al. 2014), so future study using IGCM4 remains
224 technically very feasible.

225

226 **2.4 Stratosphere**

227 A simple gravity wave drag scheme based on Lindzen (1981) had previously been
228 implemented in both IGCM1 (Joshi et al 1995) and IGCM3 (Cnossen et al. 2011).

229 The IGCM4 scheme is as above, but calculates drag based on orographic drag, as well

230 as 2 non-orographic modes having horizontal phase speeds of $\pm 10 \text{ ms}^{-1}$. The
231 orographic drag source amplitude is the magnitude of the zonal wind in the lowest
232 model layer multiplied by the subgrid-scale standard deviation of topography; the
233 non-orographic source amplitude is the magnitude of the zonal wind in the lowest
234 model layer multiplied by a constant value of 90m.

235 Stratospheric water vapour (henceforth SWV) is calculated by adding a fixed value (3
236 ppmv) onto an amount calculated by a parameterisation that considers the
237 stratospheric radiative effects of changing tropospheric methane concentrations.
238 Methane oxidation in the stratosphere depends on the stratospheric chemical
239 environment and stratospheric residence time. While both the chemical environment
240 and the Brewer-Dobson circulation may change in a changing climate, coupled
241 chemistry-climate model integrations show that their effects on stratospheric methane
242 (and hence on SWV) is small compared to the effect of the changes in methane
243 entering the stratosphere (Eyring et al 2010), which in turn is given by the change in
244 average tropospheric methane to a good approximation. Hence, the impact of
245 changing tropospheric methane can be approximated by calculating the stratospheric
246 distribution of the fraction of oxidised methane, which then is multiplied by the
247 amount of tropospheric methane to give the change in stratospheric methane and its
248 contribution to changes in SWV. We define the oxidised fraction β :

$$249 \quad \beta(\varphi, z) = 1 - CH_4(\varphi, z) / CH_4_{troposphere} \quad (2)$$

250 where z is altitude, φ is latitude, and any longitudinal variation is assumed to be
251 averaged. $CH_4(\varphi, z)$ is obtained from satellite measurements by the Halogen
252 Occultation Experiment (HALOE, Russell et al. 1993) over the period 1995-2005.
253 Assuming that two water molecules form for each methane molecule, the water
254 vapour change occurring over a given time interval is given by combining the change

255 in CH₄ over the same time interval with the scaling factor β in a similar manner to
256 Fueglistaler and Haynes (2005) giving:

$$257 \quad dH_2O(\varphi, z) = 2 * \beta(\varphi, z) * dCH_{4troposphere} \quad (3)$$

258 These calculated SWV anomalies are then supplied to the IGCM to allow calculation
259 of the influence of this additional effect on climate. This approach provides excellent
260 predictions of stratospheric methane changes in CCMVal2 models for the period
261 1960-2008 (REF-1B runs) (Eyring et al 2010).

262 Figure 2 (top right) shows an analytical approximation to this distribution, which is
263 then used to calculate β . The effect is demonstrated by showing the SWV perturbation
264 in ppmv for pre-industrial CH₄ concentrations of 0.75 ppmv (bottom left), and
265 potential future concentrations of CH₄ of 2.5 ppmv (bottom right), as might be
266 expected in the mid 21st century under the Representative Concentration Pathway
267 (RCP) 8.5 scenario (Holmes et al 2013). For reference the background SWV
268 concentration to which this perturbation is added is 3 ppmv.

269

270 **3. Model Evaluation**

271 **3.1 Surface and top-of-atmosphere model climatology**

272 The following results are all from the most commonly used configuration of the
273 IGCM4: sea surface temperature (henceforth SST) is prescribed as a monthly-varying
274 climatology based on ERA-40 reanalysis (Forster et al 2000), but land temperature is
275 calculated self-consistently from surface fluxes at each timestep. For this section, the
276 20-layer T42L20 model has been used, which has been integrated for 100 model years
277 in total.

278 Figure 3 shows the comparison between NCEP-DOE Reanalysis 2 (Kanamitsu et al.
279 2002) and IGCM4 surface temperature. During Boreal winter (December- February,
280 or DJF), Figure 3 (bottom left panel) shows that the model displays a slight cold bias
281 in Northern Eurasia, and a warm bias in the tropical regions and Antarctica. The bias
282 is mostly below 10K in amplitude, which is good for intermediate models of this type.
283 The boreal summer response (June-August, or JJA) is shown in Figure 3 (bottom right
284 panel). Here, a warm bias is present over most of the land surface. The warm bias in
285 both summer hemispheres is likely due to an absence of aerosols in the IGCM,
286 especially over North Africa and Australia where high amounts of dust occur in
287 reality. However, even during JJA the magnitude of the bias is less than 10K almost
288 everywhere, which is reasonable when compared to biases even in CMIP5 models
289 (e.g. Flato et al. 2013, Figure 9.2). Both ice caps display too large a seasonal cycle,
290 which we attribute to the simplicity of the snow scheme in the model, which has no
291 facility for changing density or conductivity when snow is compacted into ice. This
292 could be a source for future model improvement.

293 Figure 4c (top right panel) shows the precipitation bias in DJF in the IGCM compared
294 to the CMAP dataset (Xie and Arkin 1997) shown in Figure 4a (top left panel). In
295 general the comparison is quite good, with the major convergence zones (as
296 diagnosed by the 4 mm day⁻¹ contour in black) being represented quite well. As a
297 guide to the IGCM's performance in the context of other models, the mean \pm one
298 standard deviation precipitation bias amongst a subset of models present in the
299 CMIP5 archive being used for the UN Intergovernmental Panel on Climate Change's
300 5th assessment report (IPCC AR5) is also shown (Figures 4d and 4f respectively): the
301 comparison is for the CMIP5 model configuration using prescribed "AMIP" SSTs,
302 since coupled ocean-atmosphere biases tend to worsen model performance.

303 The IGCM's precipitation bias (top right panel) lies within one standard deviation of
304 the AMIP ensemble biases; for instance the dry bias in the Southern Pacific
305 Convergence Zone (SPCZ) in the IGCM (top left panel) is 2-5 mm day⁻¹, which is
306 similar in magnitude to the mean minus one standard deviation, suggesting that the
307 IGCM's performance in this region is within the envelope of state-of-the-art GCMs
308 forced by observed SSTs.

309 Figure 5 is the same as Figure 4, but for the JJA period. There are some notable wet
310 biases in IGCM4 as shown by Figure 5c (top right panel), particularly in the northern
311 Indian Ocean and Central American regions: however such wet biases are not outside
312 the envelope of the CMIP5 ensemble when comparing the IGCM to the "mean plus
313 one standard deviation" (Figure 5f- bottom right panel). Thus, for the JJA season as
314 well as the DJF season, the precipitation bias in IGCM4 is within the range of state-
315 of-the-art GCMs forced by observed SSTs, which provides a good justification for the
316 use of IGCM4 as a simplified climate model.

317 The interaction of precipitation, cloud and radiation, can be studied by comparing the
318 outgoing long-wave radiation (OLR) field with observations (Liebmann and Smith
319 1996), which is shown in Figure 6. The bottom left panel shows that the IGCM
320 broadly simulates OLR quite well, with some differences between model and
321 observations in the Maritime continent region. During JJA (bottom right panel), there
322 is a positive bias in OLR over the Indian Ocean (Figure 6f- bottom right panel),
323 consistent with a slight dry bias there (Figure 4c). The top-of-atmosphere energy
324 imbalance in the IGCM is approximately 1-2 Wm⁻², which is similar to other climate
325 models (e.g. Roeckner et al. 2006).

326

327 **3.2 Zonal mean climatology and stratospheric performance**

328 For this section, both 20-layer T42L20 and 35-layer T42L35 configurations are
329 described; the latter has been integrated for 200 model years in total, in order to
330 average out the effect of stratospheric variability. Figure 7 shows the zonally averaged
331 temperature structure in IGCM4 for the two solstitial seasons compared to data from
332 the ERA40 reanalysis (Uppala et al 2005). In both seasons the lower stratosphere in
333 both L20 and L35 configurations is too cold in the tropics and the winter extratropics
334 by 5-10K. Elsewhere, biases are smaller than 10 K apart from near the summer
335 stratopause, perhaps due to deficiencies in the ozone heating in IGCM4. These errors
336 are comparable models that represent the stratosphere (e.g. Eyring et al. 2006).

337 A comparison between the zonally-averaged zonal wind in IGCM4 and ERA40 is
338 shown in Figure 8, and like Figure 7, also shows good agreement, perhaps not
339 surprisingly for a field that is expected to be in large-scale thermal balance with
340 temperature. In both L20 and L35 configurations, the southern hemisphere
341 tropospheric jetstream is slightly equatorward of the jet in ERA40 as shown by the
342 dipole pattern in colours in Figures 8 c-f in this region. During DJF the northern
343 hemisphere's tropospheric jetstream is slightly too strong in both L20 (Figure 8c) and
344 L35 (Figure 8e) by 5 ms^{-1} . In general, both L20 and L35 configurations display
345 similar tropospheric biases in zonal wind.

346 During DJF, the strength of the stratospheric jetstreams in the L35 configuration
347 IGCM4 compares well to ERA40 (Figure 7e). In northern winter especially this is a
348 sign that the joint effects of gravity wave drag and tropospheric wave forcing in
349 IGCM4 are approximately of the right magnitude, since these two factors play a
350 crucial role in controlling the strength of the DJF winter stratospheric jetstream. In
351 JJA however the stratospheric jetstream is weaker and less tilted in the vertical than

352 ERA40 (Figure 7f). This bias is likely due to the simplicity of the gravity wave drag
353 scheme (see above), and might be removed by more tuning of the drag scheme- but
354 this would require more multi-century L35 integrations to ensure that tuning did not
355 result in greater biases elsewhere: as such it is a source for future development.

356 The zonally asymmetric component of the circulation is apparent from Figure 9,
357 which shows the geopotential height eddy fields at 500 and 200 hPa. The IGCM4
358 reproduces the main features of the reanalysis with the standing wave patterns
359 apparent in both model configurations, although low pressure anomaly in NE Asia is
360 weaker in both model configurations compared to reanalysis. Both L35 and L20
361 configurations display a similar standing wave pattern at both pressure levels.

362 A key issue for stratospheric dynamics and its interplay with tropospheric climate,
363 which is a primary use of this model, is that the stratospheric circulation, and
364 phenomena such as sudden stratospheric warmings (henceforth SSWs) are simulated
365 as well as other models. A 200-year long integration of IGCM4 yielded 0.57 SSWs
366 per year as diagnosed by the method of Charlton and Polvani (2007); this should be
367 compared with 0.6 as diagnosed in reanalyses by Charlton and Polvani (2007). 57%
368 of the SSWs were categorised as “displacement” events using a vortex moment
369 method based on Mitchell et al. (2011), and 43% diagnosed as “split” events, again
370 broadly consistent with reanalysis output which suggests that just under half of SSWs
371 can be categorised as “split” events (Charlton and Polvani 2007). The timing of
372 SSWs during boreal winter is shown in Figure 10. Again, the timings are broadly
373 consistent with reanalysis output, although there are somewhat more displacement
374 events during March than diagnosed from reanalysis.

375

376 **4. Climate Change and Energy Balance**

377 When coupled to a slab q-flux ocean model, IGCM4 has an equilibrium climate
378 sensitivity when doubling CO₂ from its pre-industrial concentration of 280 ppmv of
379 2.1K. This sensitivity is slightly higher than the value of 1.6K in IGCM3 (Joshi et al
380 2003), and is likely due to the changes in cloud physics outlined above.

381 We have not performed simulations of a slab model for this paper because although
382 one effect of a slab ocean is to change the characteristics of model interannual
383 variability (as shown by Winter and Bourqui 2011a), the nature of such changes will
384 depend on the depth of the slab, and how this depth changes seasonally and
385 geographically: for instance in the North Atlantic Ocean the effective mixed layer
386 depth changes from 50 m during summer to 500 m in winter. Moreover, the dynamic
387 influence of the atmosphere on the ocean will also depend on the effective mixed
388 layer depth of the ocean, or depth of the slab, as shown by O' Callaghan et al (2014),
389 as well as causing a dynamical ocean response (Zhai et al 2014).

390 Because interannual variability is sensitive to slab ocean depth, and the IGCM has a
391 constant slab depth, rather than one that varies seasonally and geographically, we
392 have not discussed interannual variability in this paper. However, such a topic would
393 be a source of useful research in the future for a configuration of the IGCM that had
394 such a varying slab ocean model.

395 As a first assessment of coupled model performance, the zonally averaged net surface
396 energy imbalance and wind stress curl in IGCM4 are examined and compared to
397 reanalysis, since large errors in these two fields will give errors in the dynamic and
398 thermodynamic ocean responses respectively. Figure 11 shows that the broad patterns
399 of response are similar in both model and reanalysis. In equatorial regions incoming

400 solar radiation is not quite balanced by outgoing IR emission because of the presence
401 of tropical convection and thick clouds, leading to positive values (see top panel); the
402 intense rainfall associated with such convection is shown in the top panels of Figures
403 4 and 5. In subtropical regions, a lack of cloud leads to more IR emission and negative
404 values in both reanalysis and IGCM4. The pattern of wind stress curl (see bottom
405 panel) is indicative of the combined effects of midlatitude westerlies, and subtropical
406 and tropical trade winds, and is similar in both model and reanalysis apart from the
407 southern ocean westerlies being slightly too equatorward in the model, and the Arctic,
408 where the IGCM fails to reproduce large values associated with mesoscale
409 circulations (e.g. Condrón and Renfrew 2013) that the model cannot represent given
410 its horizontal resolution.

411 To summarise, we have presented the physical details, and major climatological and
412 dynamical features of the IGCM4 climate model. The model provides a fast
413 alternative to conventional state-of-the-art GCMs while retaining the richness of
414 dynamical behaviour allowed by the primitive equations of meteorology. As such the
415 IGCM4 forms a useful part of the “hierarchy of models” approach needed to fully
416 understand climate.

417

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428 **6. Code Availability**

429 The code is available to scientific researchers on request by emailing
430 m.joshi@uea.ac.uk in the first instance. Websites detailing different IGCM
431 configurations are given in section 2.2. IGCM4 requires as a prerequisite a fortran
432 compiler, the nupdate code management utility, and MPI routines for parallel
433 integrations (although IGCM4 is designed to run on one processor).

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435 **7. References**

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574 **8. Table Captions**

575 Table 1: Values of surface characteristics for each surface type in IGCM4(ice, inland
 576 water, forest, grassland, agriculture, tundra, swamp, desert).

Surface Type	Albedo	Snow-covered albedo	Height when albedo is snow-covered (m)	Roughness length (m)
Ice	0.8	0.8	0.05	0.03
Inland Water	0.2	0.6	0.05	0.001
Forest	0.25	0.7	0.1	0.1
Grassland	0.25	0.8	0.1	0.05
Agriculture	0.25	0.8	0.1	0.05
Tundra	0.3	0.8	0.05	0.03
Swamp	0.2	0.8	0.05	0.03
Desert	0.3	0.8	0.05	0.03

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583 **9. Figure Captions**

584 Figure 1: Model layer index vs pressure (for a surface pressure of 1000 hPa) for the
585 35 layer model (black) and the 20 layer model (red). Note that the lowest 19 layers are
586 exactly the same for both configurations.

587 Figure 2: The fraction of oxidised methane (which is linked to CH₄ concentration- see
588 equation 1) derived from HALOE data (top left panel); the analytical approximation
589 which extends to the poles (top right panel); the perturbation to stratospheric water
590 vapour (SWV) (ppmv) in pre-industrial conditions, when CH₄ is 0.75 ppmv (bottom
591 left); the perturbation to SWV (ppmv) if CH₄ is increased to 2.5 ppmv (bottom right).

592 Figure 3: Surface temperature (°C) in IGCM4 (a,b), NCEP-DOE reanalysis (c,d) and
593 difference between IGCM4 and reanalysis (e,f). In all cases the left-hand panels
594 display results for the DJF season and the right-hand panels display results for JJA
595 season. For the reanalysis a mean over the years 1979-2013 is taken.

596 Figure 4: DJF season mean precipitation (mm day⁻¹) in CMAP (a), IGCM4 (b) and
597 difference between IGCM4 and CMAP (c). Subfigure (e) shows the difference
598 between a multi model mean of an ensemble of CMIP5 GCMs integrated using AMIP
599 SSTs and CMAP; (f) as for (e) but for the multi model mean minus one standard
600 deviation; (g) as for (e) but for the multi model mean plus one standard deviation. In
601 all cases the solid line is the 4 mm day⁻¹ contour in CMAP and the dashed line is the
602 same contour in the model of the subfigure. Subfigures (a,b) are based on the top
603 colour bar, subfigures (c-f) are based on the bottom colour bar. The CMIP5 models
604 used in the ensemble are: ACCESS1.0, ACCESS1.3, BCC-CSM1.1, BCC-
605 CSM1.1(m), BNU-ESM, CanCM4, CCSM4, CESM1(CAM5), CCMC-CM, CNRM-
606 CM5, CSIRO-Mk3.6.0, FGOALS-g2, GFDL-CM3, GISS-E2-R, HadGEM2-AO,

607 INM-CM4, IPSL-CM5A-LR, IPSL-CM5A-MR, IPSL-CM5B-LR, MIROC5, MPI-
608 ESM-LR, MPI-ESM-MR, MRI-CGCM3, NorESM1-M, mean over the years 1979-
609 2005.

610 Figure 5: As for Figure 4, but during the JJA season.

611 Figure 6: Outgoing longwave radiation or OLR (W m^{-2}) in IGCM4 (a,b), interpolated
612 OLR dataset (c,d) and difference between IGCM4 and interpolated OLR dataset
613 (Liebmann and Smith 1996) (e,f). In all cases the left-hand panels display results for
614 the DJF season and the right-hand panels display results for JJA season. For the
615 interpolated OLR dataset a mean over the years 1979-2011 is taken.

616 Figure 7: Zonally averaged temperature (K) in ERA (a,b), difference between IGCM4
617 L20 and ERA (c,d) and difference between IGCM4 L35 and ERA (e,f) in colour
618 shading. In all subfigures contours show the total zonal mean temperature field
619 (contour interval is 10 K, 240K contour thicker). In all cases the left-hand panels
620 display results for the DJF season and the right-hand panels display results for JJA
621 season. For the reanalysis a mean over the years 1958-2002 is taken.

622 Figure 8: Zonally averaged zonal wind (ms^{-1}) in ERA (a,b), difference between
623 IGCM4 L20 and ERA (c,d) and difference between IGCM4 L35 and ERA (e,f) in
624 colour shading. In all subfigures contours show the total zonal mean zonal wind field
625 (contour interval is 10 ms^{-1} , negative contours dashed, zero contour dotted). In all
626 cases the left-hand panels display results for the DJF season and the right-hand panels
627 display results for JJA season. For the reanalysis a mean over the years 1958-2002 is
628 taken.

629 Figure 9: Geopotential Height (m) DJF Eddy Fields for: (a), (b) 200 hPa and 500 hPa
630 ERA-40 Reanalysis respectively. The same for (c), (d) IGCM4 L20 and (e), (f)
631 IGCM4 L35. For the reanalysis a mean over the years 1958-2002 is taken.

632 Figure 10: Distribution of sudden stratospheric warmings in boreal winter by month in
633 the IGCM4 (filled grey boxes) and reanalysis (red outline boxes) (top panel);
634 distribution of displacement-type warmings (middle panel); distribution of split-type
635 warmings (bottom panel).

636 Figure 11: Annually averaged net downward zonal surface energy imbalance (Wm^{-2})
637 in IGCM4 (black) and NCEP reanalysis (red) (top panel); Wind stress curl (10^{-7}Nm^{-3})
638 in IGCM4 (black) and NCEP reanalysis (red) (bottom panel).

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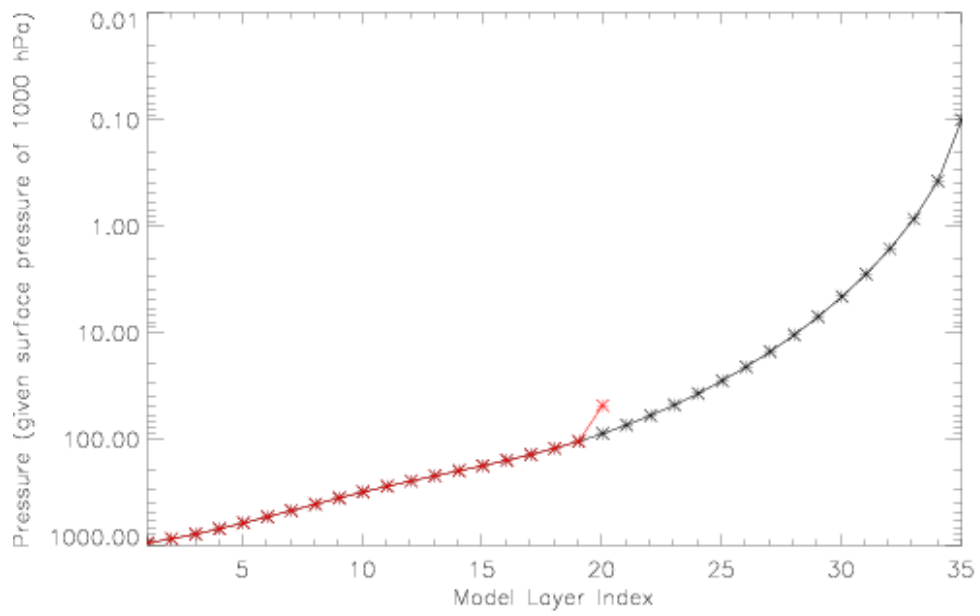
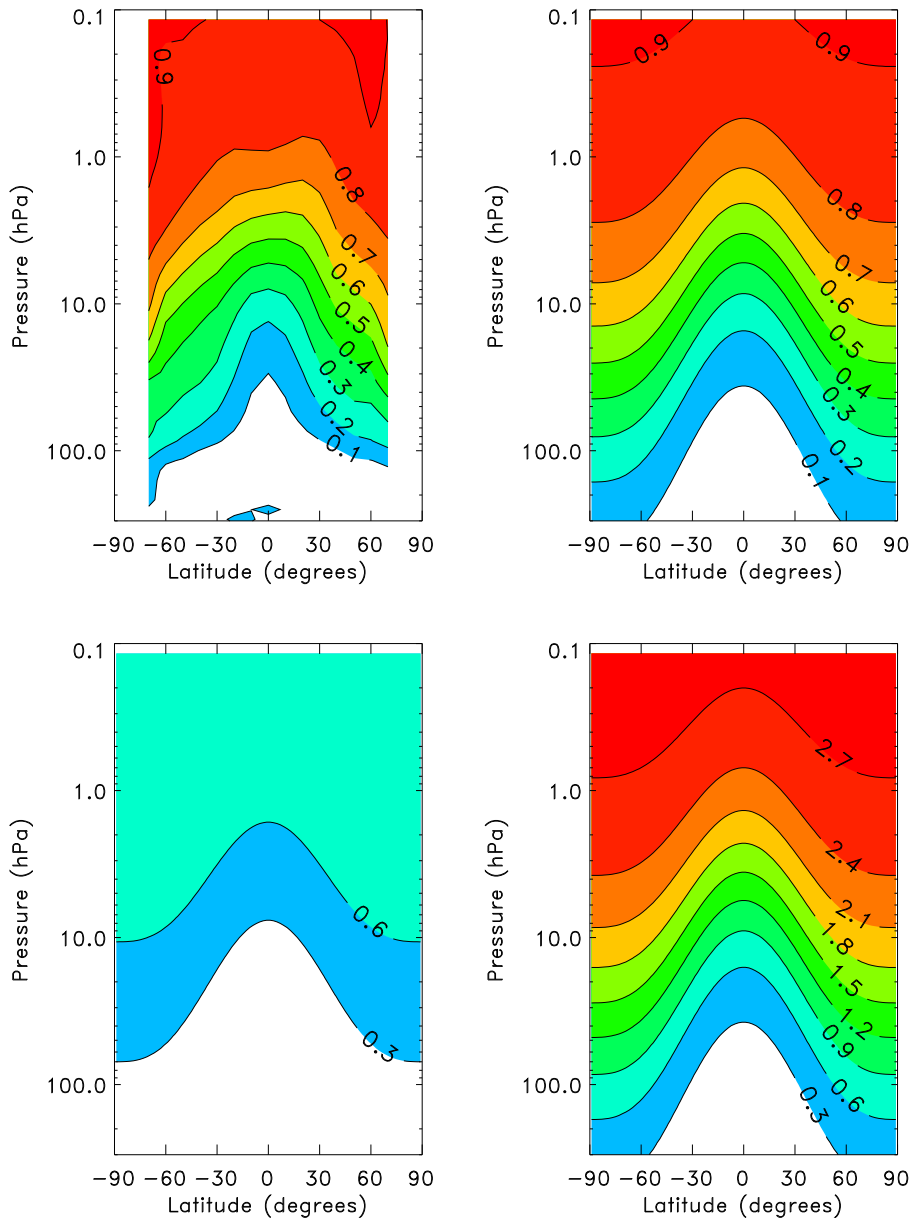


Figure 1



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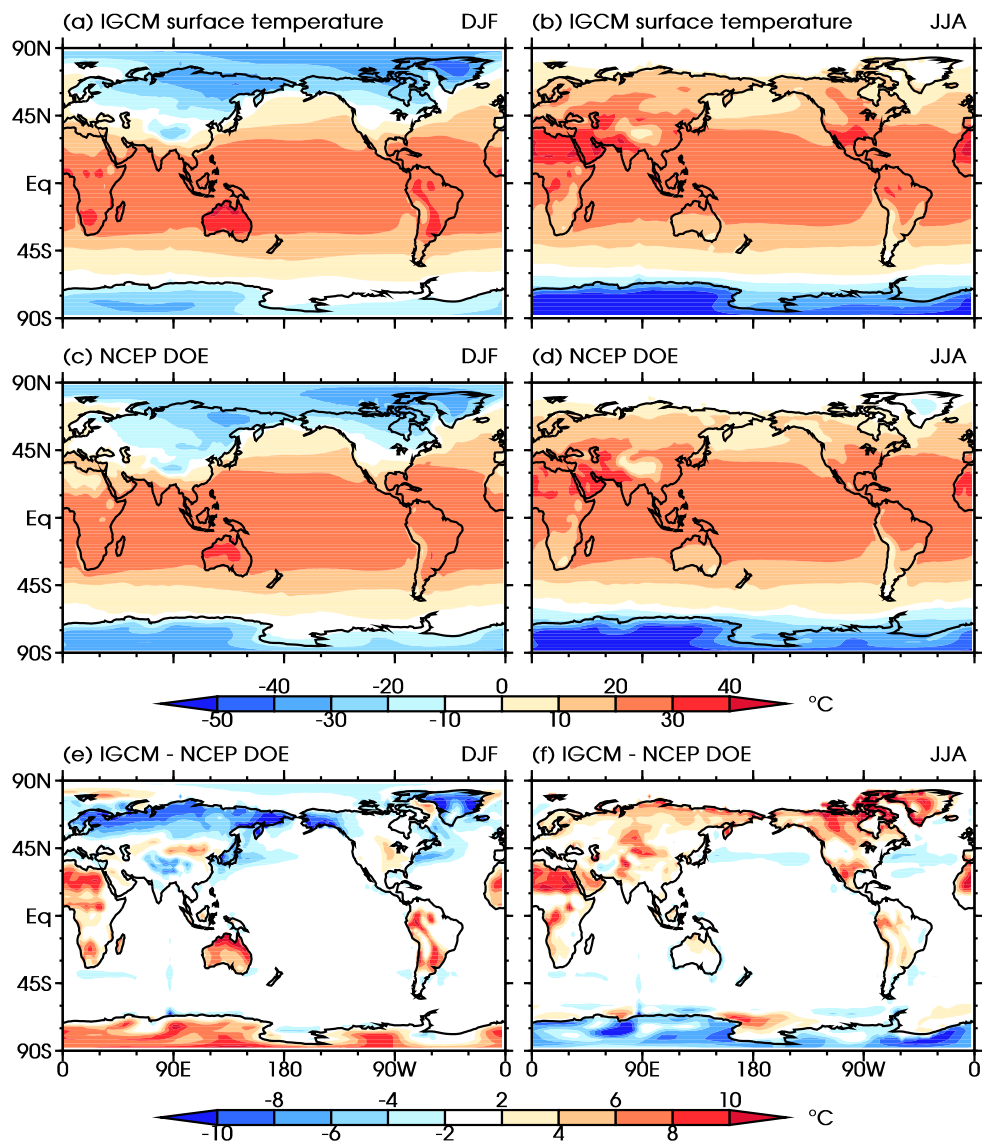
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Figure 2

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Figure 3

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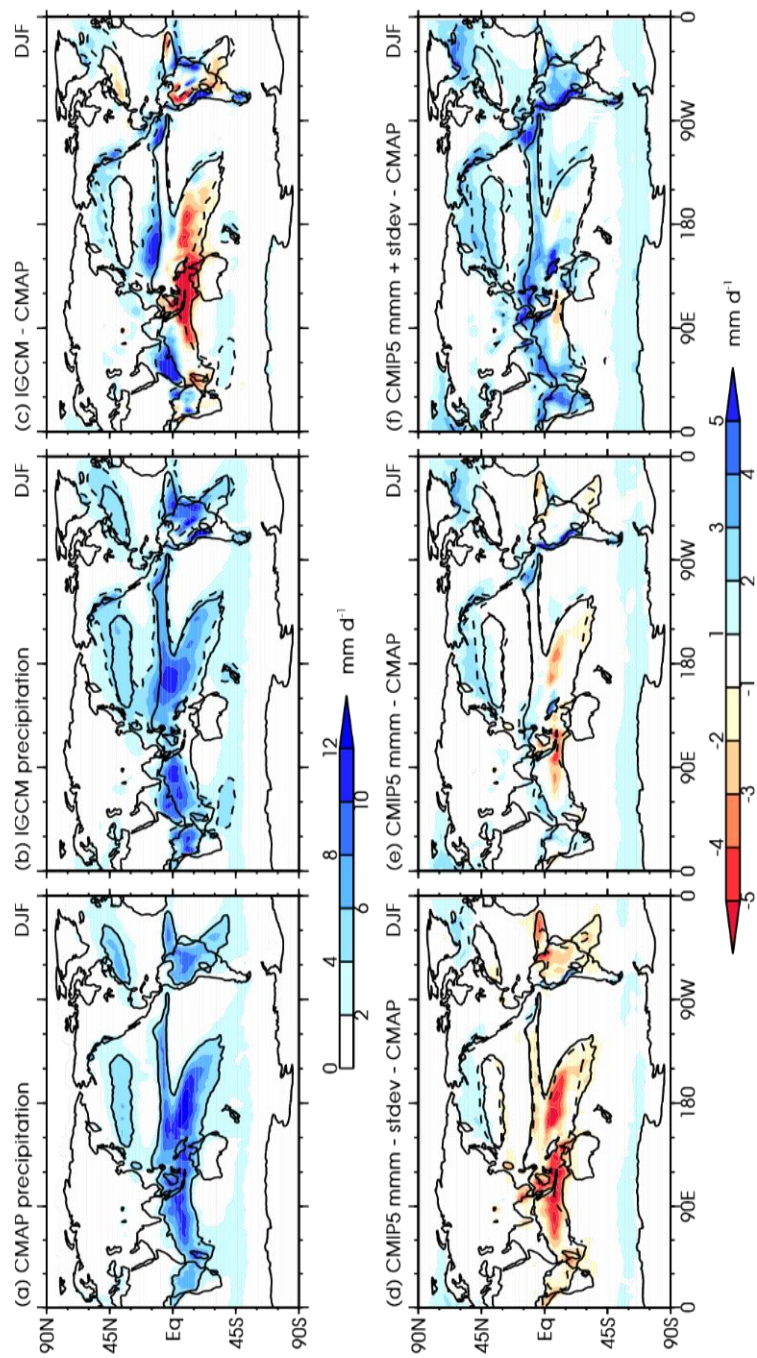
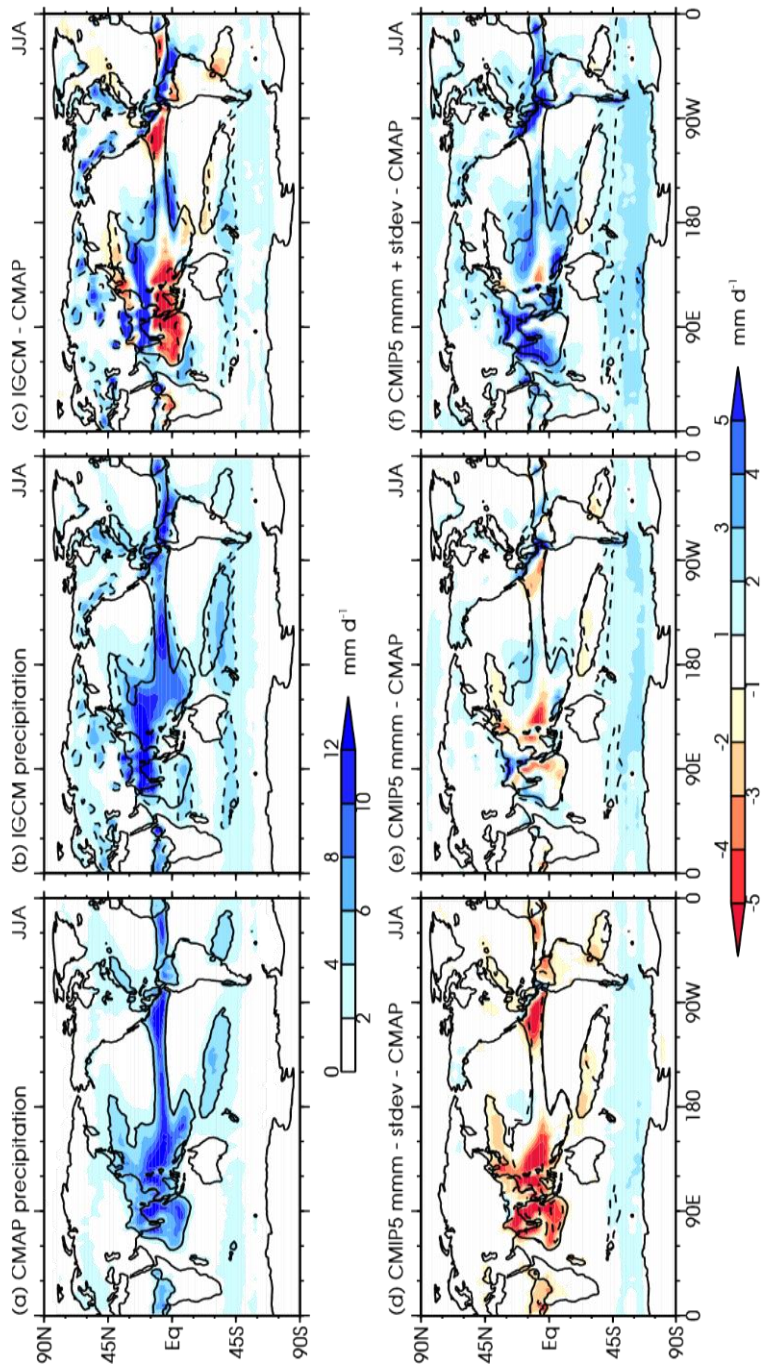


Figure 4



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Figure 5

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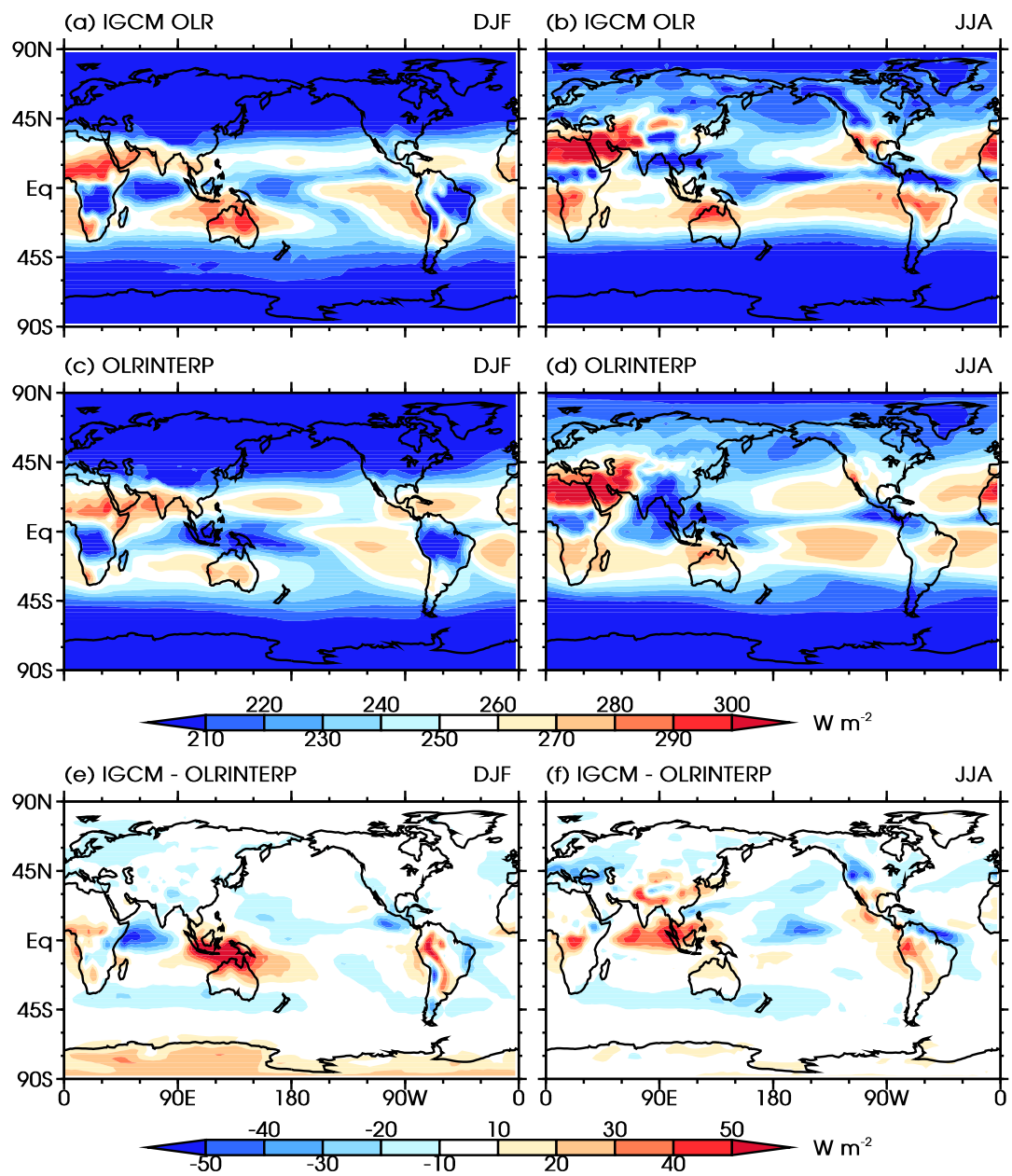
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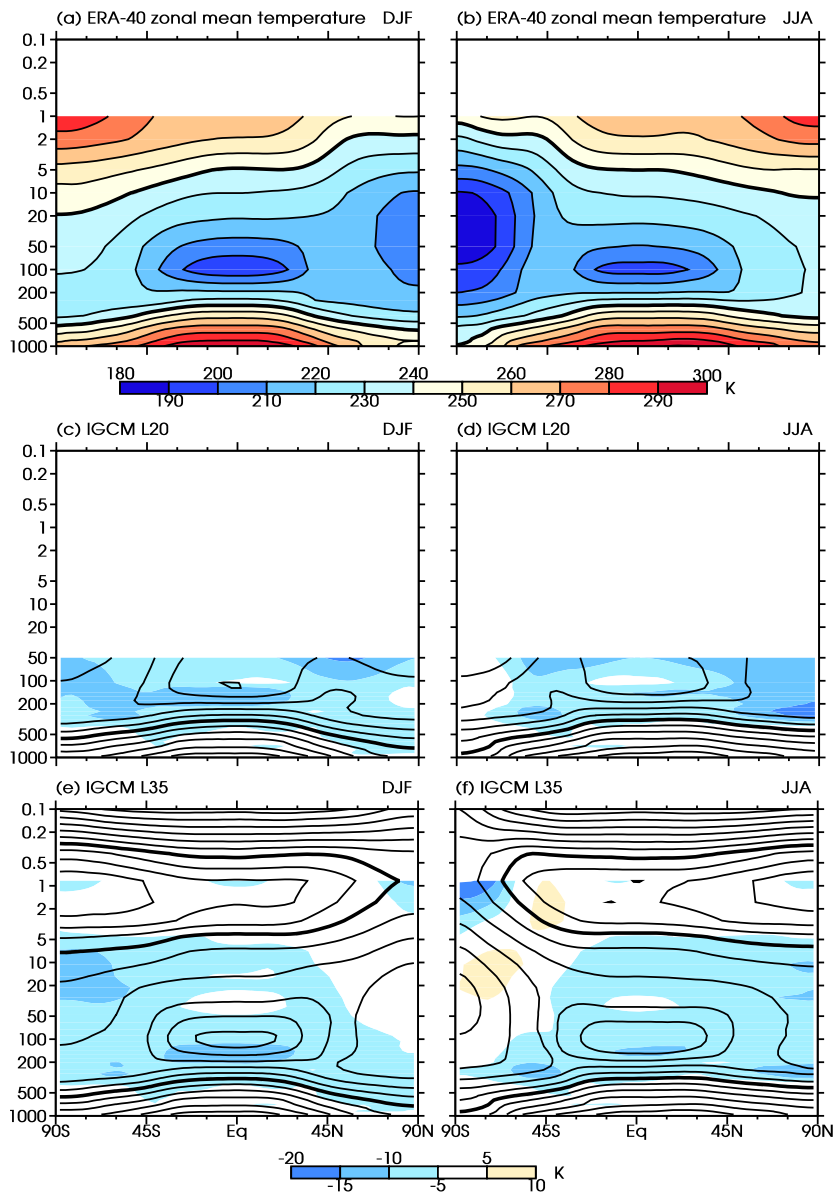
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Figure 6



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Figure 7

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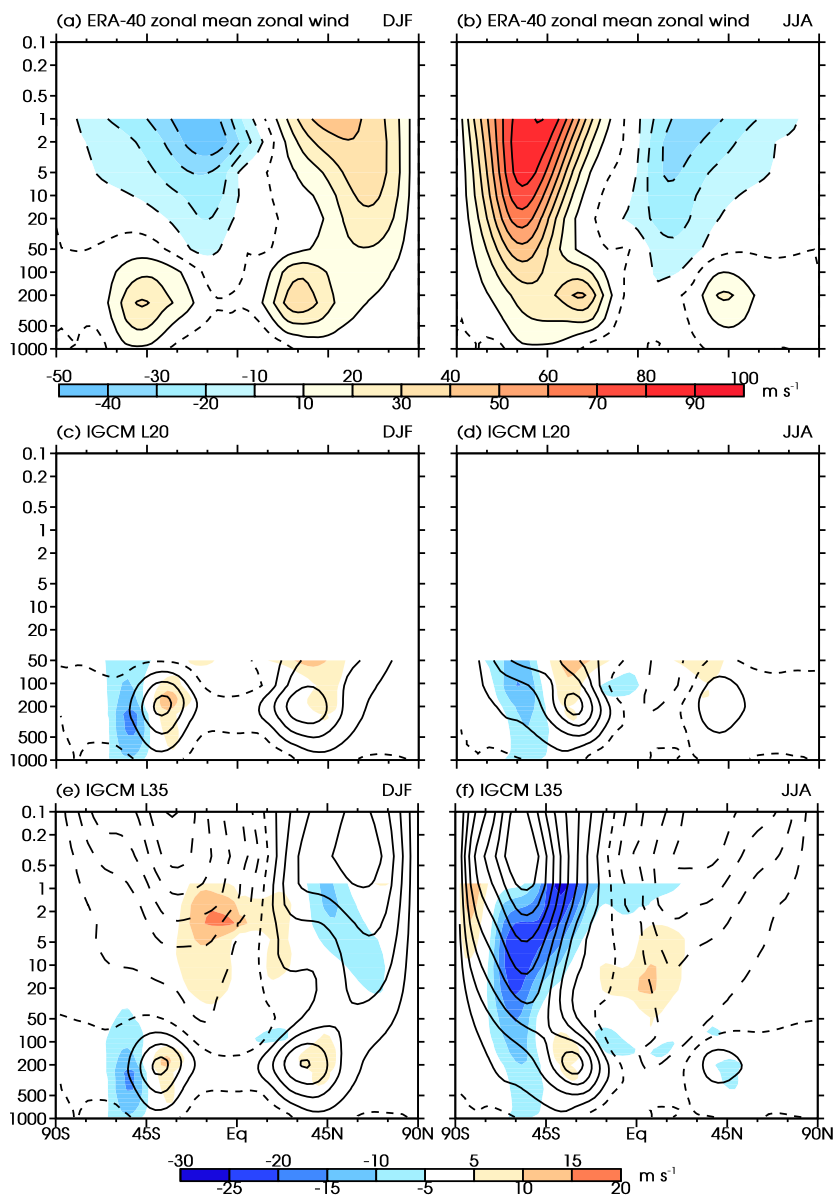
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Figure 8

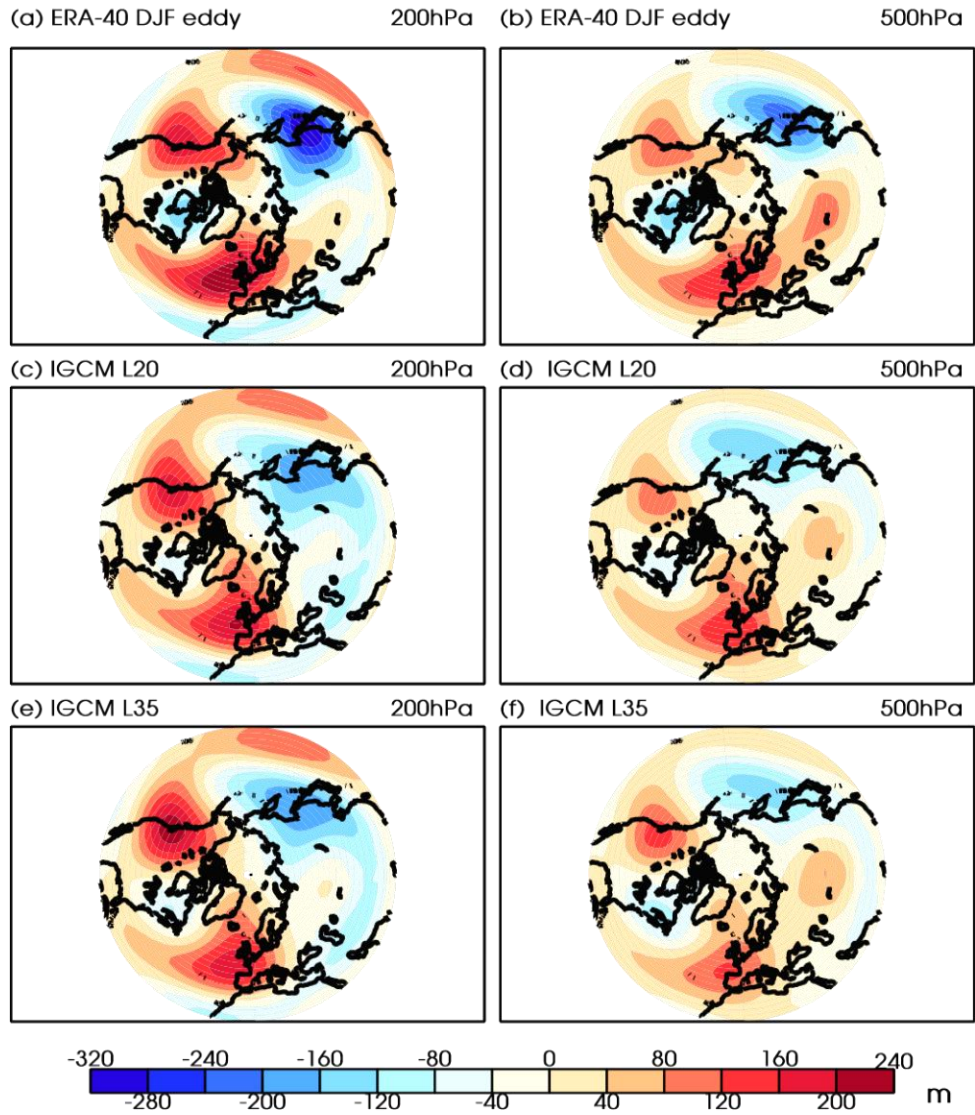
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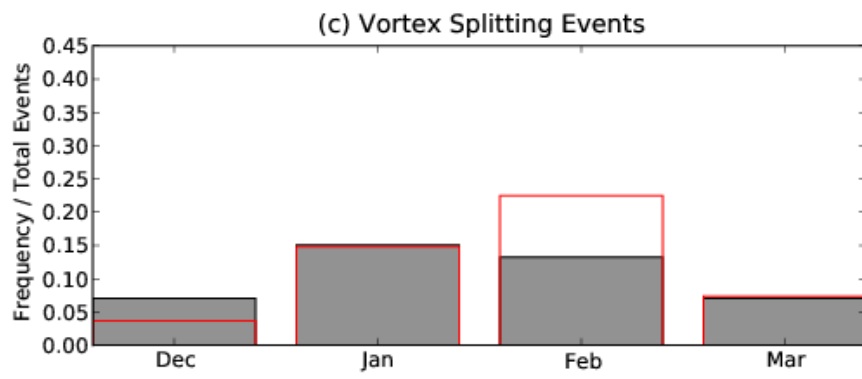
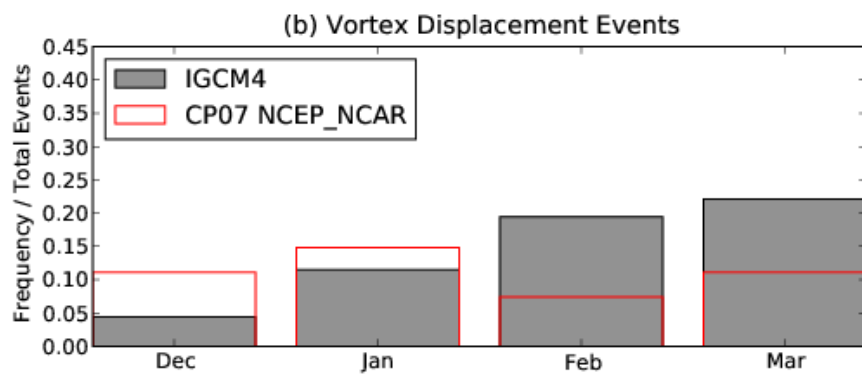
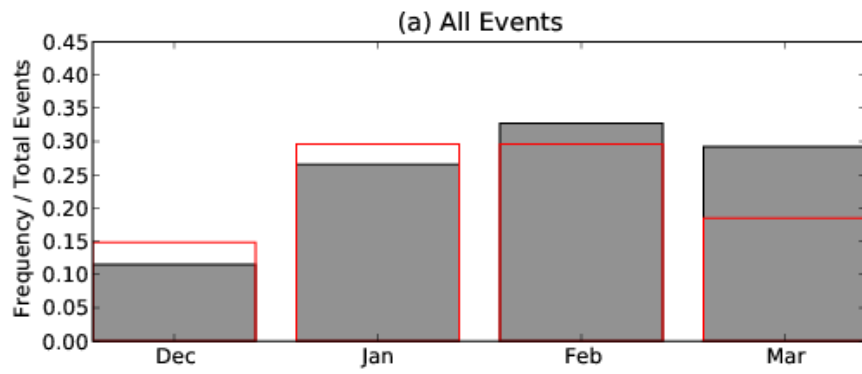


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Figure 9



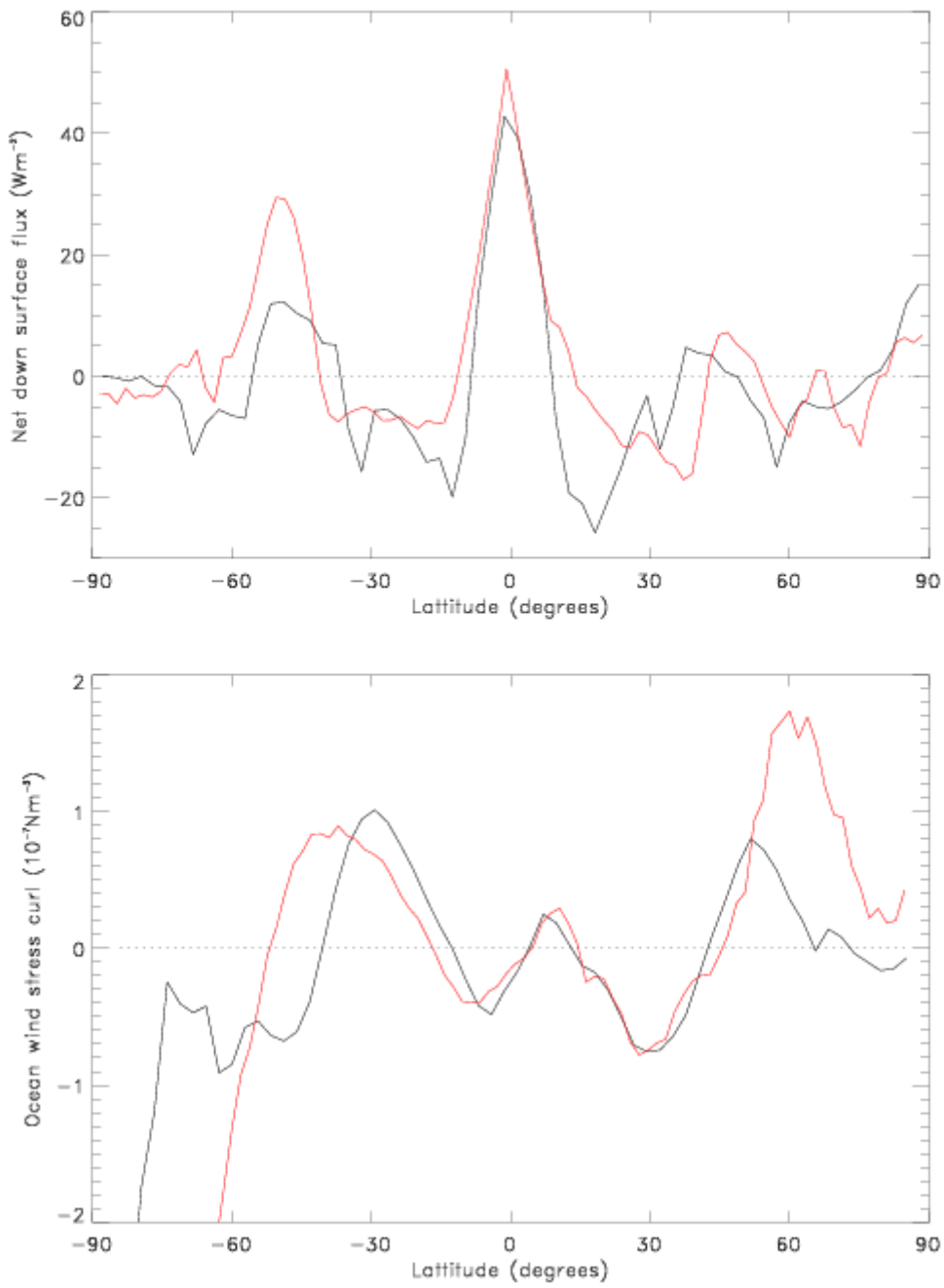
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Figure 10

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Figure 11